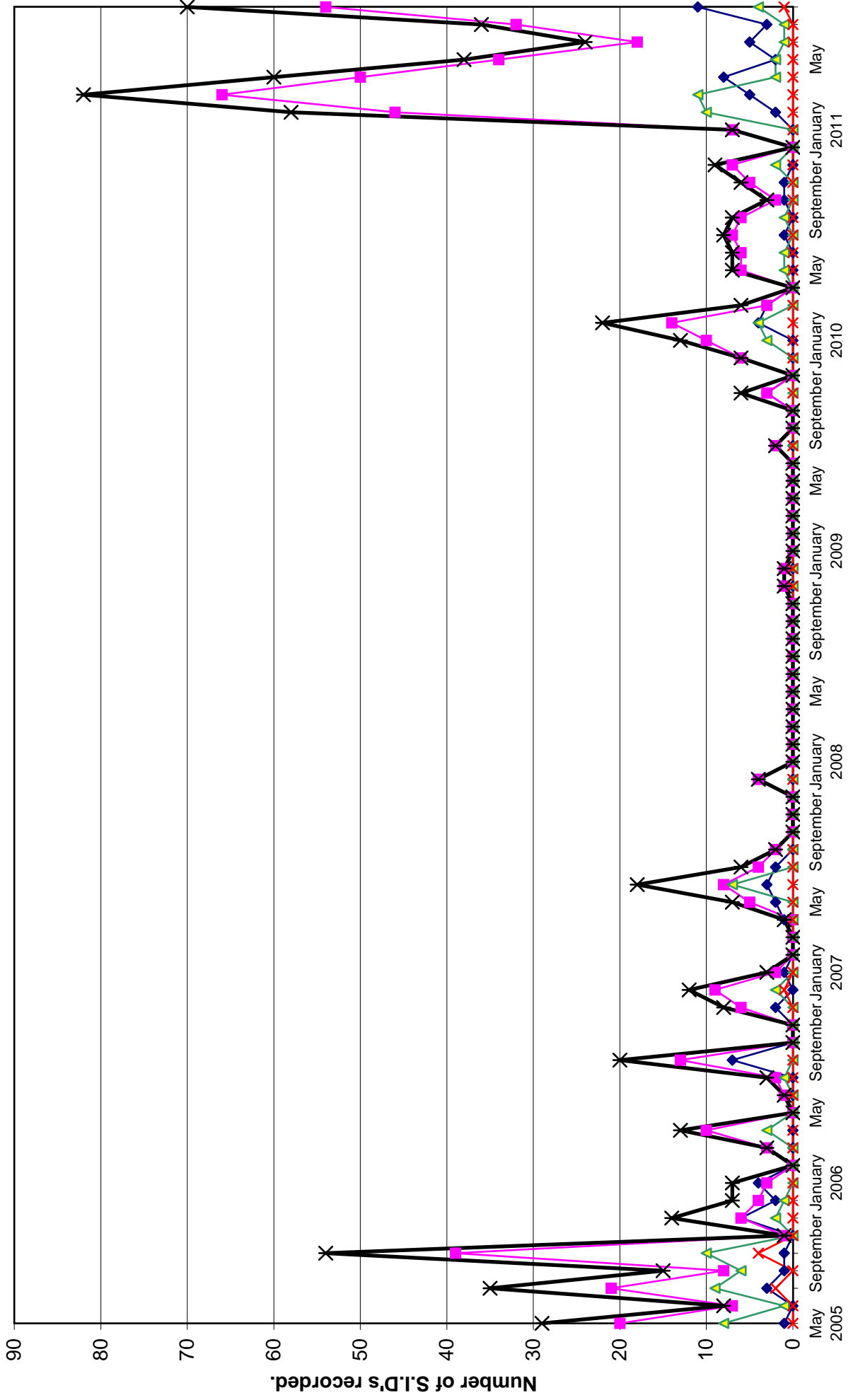


		Colin Clements (23.4kHz/37.5kHz)	Peter Meadows (23.4kHz)	Mike King (20.9kHz)	John Wardle (19.6/23.4kHz)	Peter King (18.3kHz)
		AAVSO receiver, 0.76m screened loop aerial.	Tuned radio frequency receiver, 0.58m frame aerial.	AAVSO receiver. Tuned loop aerial.	PC soundcard, long wire aerial.	Own designed receiver, 1.4m loop aerial.
DAY		START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)
1	C4.0				07:30 07:40 08:13 2	
1	C1.4					08:10 08:15 08:18 1-
1	C2.0				12:35 12:43 13:00 1	12:35 12:40 12:45 1-
1	*					
1	C1.3			13:40 13:42 13:56 1-		16:50 17:05 17:10 1
2	M1.4				06:09 06:21 07:08 2+	11:35 11:40 11:44 1-
2	C2.0					
2	C1.1					
2	B8.5					
3	M1.7					
3	C8.7	07:56 08:02 08:29 2			07:44 07:59 08:45 2+	07:40 08:00 08:07 1+
3	C1.0					
3	*					
3	M6.0	13:09 13:43 15:41 3+			13:19 13:51 15:00 3	13:17 13:50 14:10 2+
3	C2.3					18:52 19:00 19:05 1-
3	C8.5	19:29 19:33 19:42 1-			19:24 19:32 19:40 1-	
4	C3.5	07:59 08:08 08:28 1+			08:01 08:07 08:30 1+	
4	C2.5	09:19 09:26 10:05 2+			09:18 09:21 09:31 1-	09:15 09:20 09:25 1-
4	C1.6				10:06 10:15 10:23 1-	10:05 10:13 10:20 1-
4	C1.2					11:44 11:50 11:55 1-
4	C2.1	13:06 13:10 13:33 1+			13:05 13:11 13:18 1-	13:05 13:08 13:13 1-
4	*					
4	*					
4	C1.0			14:06 14:08 14:34 1+		14:10 14:18 14:25 1-
4	C1.8	14:36 14:41 15:13 2			14:35 14:40 14:51 1-	14:30 14:40 14:45 1-
5	C1.0					10:25 10:30 10:38 1-
5	C2.0	12:41 12:52 13:09 1+			12:40 12:49 13:34 2+	12:40 12:39 12:45 1-
5	C1.8					17:40 17:48 17:55 1-
5	C1.4					19:40 19:45 19:50 1-
5	B8.0					
6	C4.1	08:42 08:51 09:26 2			08:41 08:50 09:05 1	08:38 08:47 08:50 1-
6	C1.3				11:43 11:58 12:10 1+	11:40 11:45 11:55 1-
6	C1.4				12:17 12:27 12:30 1-	12:15 12:25 12:30 1-
6	C1.4					16:55 18:05 18:40 3
6	C1.4					19:20 19:25 19:30 1-
7	C1.6					08:25 08:30 08:38 1-
7	C2.1				09:25 09:30 09:48 1	09:25 09:28 09:30 1-
7	C1.3				14:01 14:09 14:18 1-	14:00 14:05 14:10 1-
8	C2.2				16:34 16:38 16:53 1	16:30 16:35 16:40 1-
8	M3.5	18:07 18:10 18:54 2+			18:03 18:11 19:14 2+	18:00 18:10 18:20 1
8	C7.7					
9	X6.9				07:54 08:03 10:00 3+	07:49 08:05 08:10 1
9	C2.2				13:33 13:45 14:05 1+	13:30 13:45 13:58 1+
9	C2.4			15:56 15:58 16:52 2+	15:51 15:58 16:14 1	15:45 15:55 16:05 1
9	C2.0				18:12 18:25 18:33 1	18:05 18:10 18:44 2
10	C2.4				09:33 09:40 09:54 1	09:30 09:38 09:45 1-
10	C6.1	10:34 10:47 11:56 2+			10:32 10:43 11:30 2+	10:20 10:45 10:55 2
10	C1.2					14:05 14:10 14:12 1-
10	C1.5				15:17 15:28 15:42 1	14:50 15:25 15:40 2+
11	C1.3				07:42 07:50 07:57 1-	
11	?				09:37 09:46 09:56 1	
11	?					
11	C6.2	09:38 10:22 11:37 3			09:56 10:06 10:40 2	09:35 10:23 10:35 2+
12	B3.6					
15	C3.4	11:14 11:24 11:47 2			11:15 11:25 12:00 2	11:10 11:20 11:35 1
15	B8.6				14:00 14:18 14:32 1+	
17	C3.4				10:26 10:43 11:20 2+	10:20 10:45 11:10 2+
17	C2.6				16:19 16:32 16:49 1+	16:15 16:30 16:45 1+
18	B9.7					
18	C1.1					15:00 15:15 15:30 1+
20	B5.7					
20	B8.4					
21	B6.1					
21	B5.9					
21	*					
21	C1.5				18:23 18:44 18:52 1+	
22	C1.1					
24	C1.1					16:30 16:33 16:35 1-
25	B9.6					
26	C2.3				13:10 13:19 13:31 1	
29	C3.7				07:08 07:24 08:13 2+	
29	?					
29	?					
29	C2.9				11:30 11:38 11:46 1-	
29	?					
29	?					
29	C1.1					
30	C1.5				13:51 14:00 14:31 2	
30	B8.8					
30	*					10:15 11:20 11:50 3
30	B7.5					
30	C1.5					15:55 16:00 16:05 1-
30	*					
31	C2.2				11:19 11:29 11:57 2	

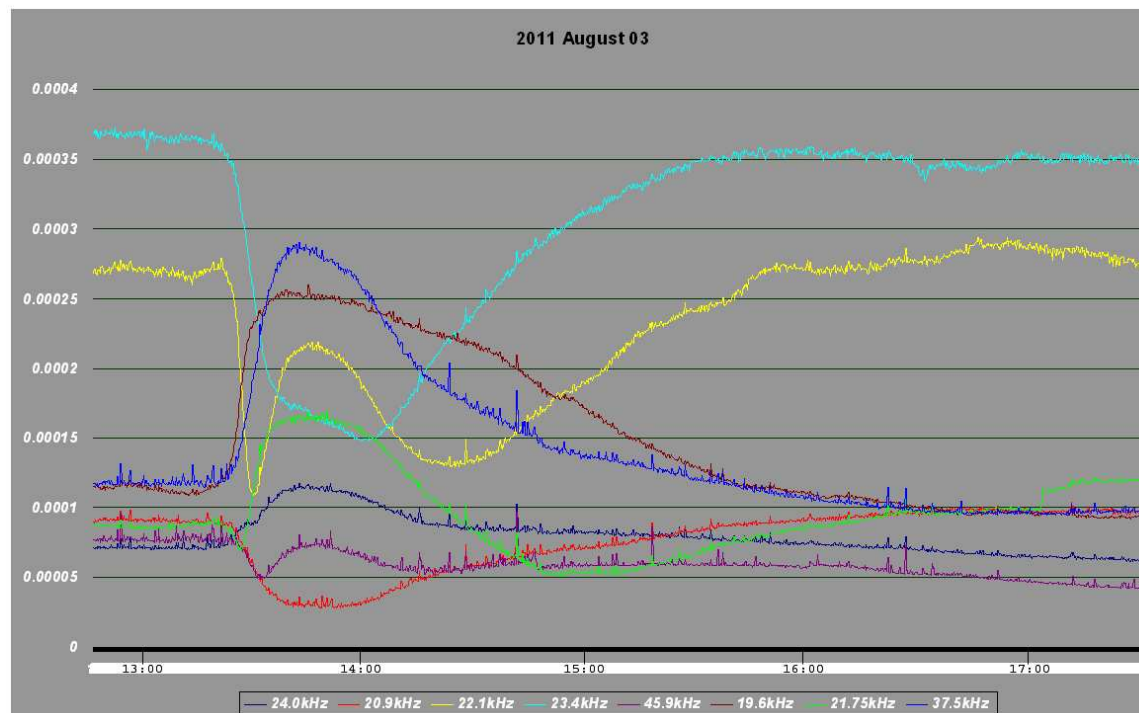
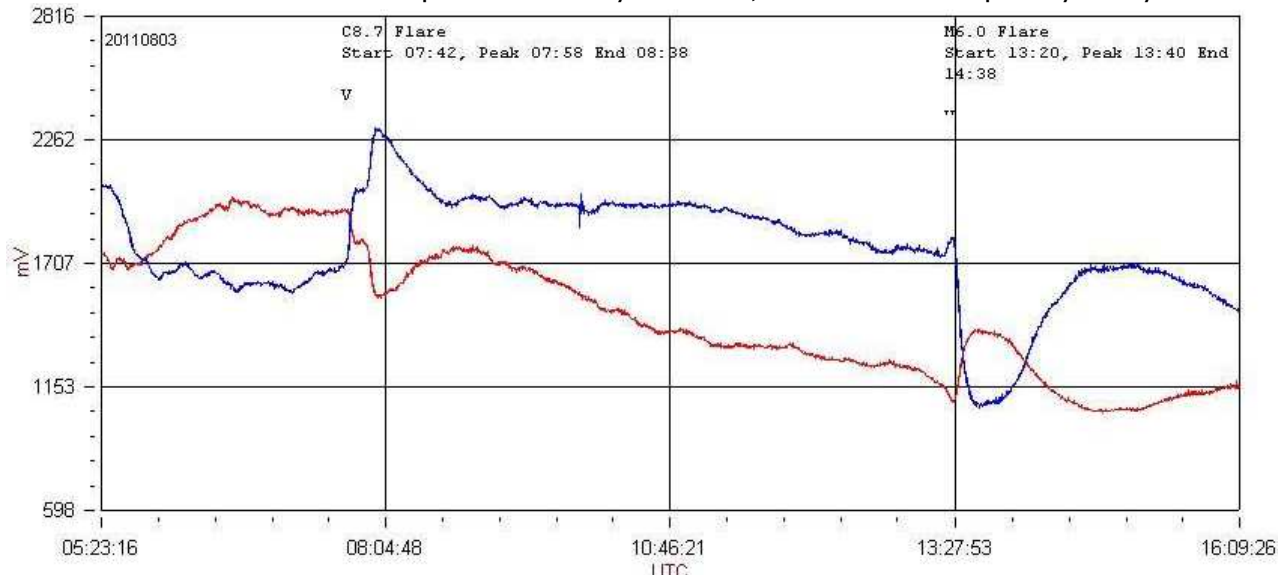
VLF flare activity 2005/11.



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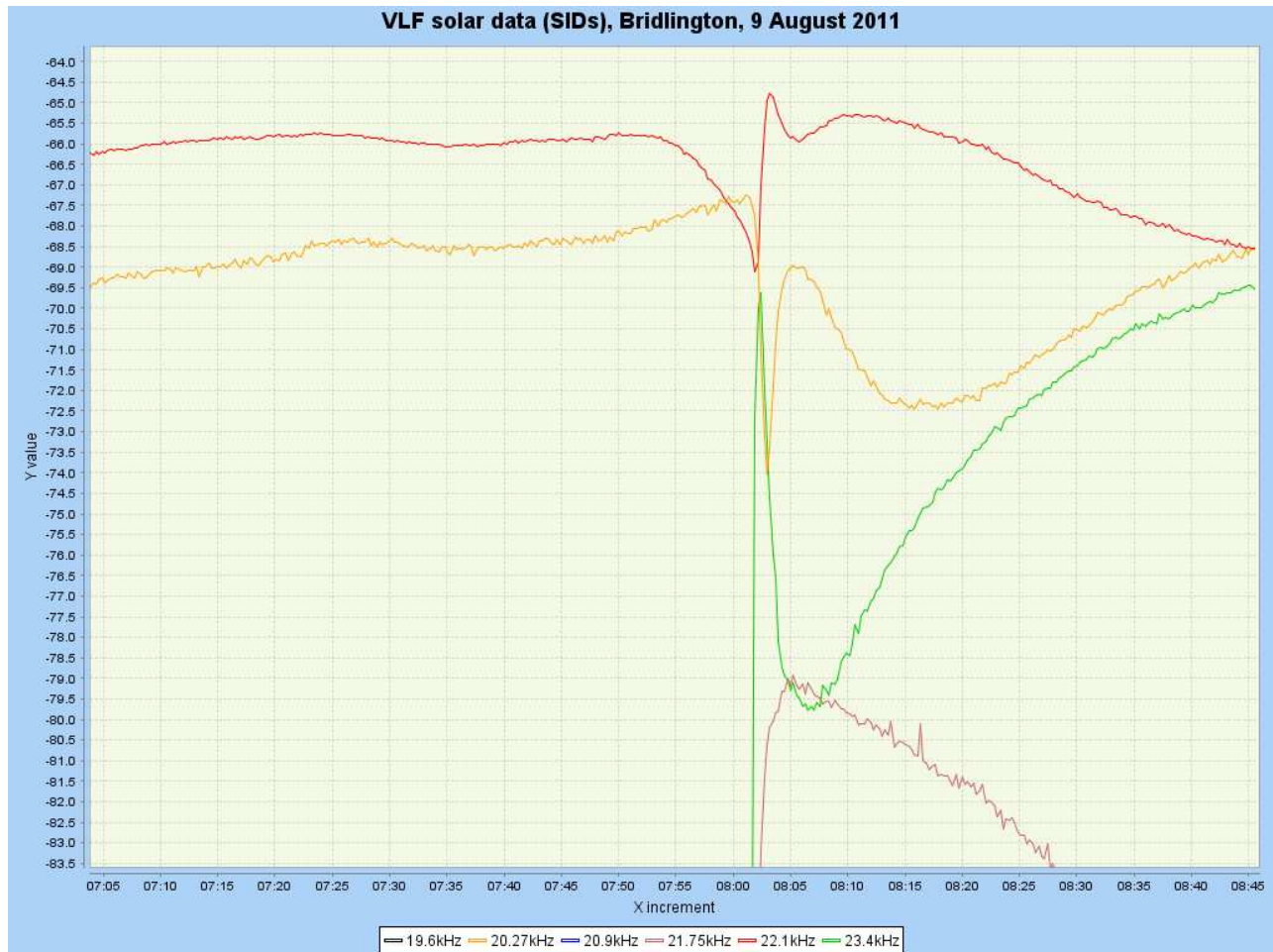
This month has seen the first X-class flare recorded by the group since 2006 December 5th. There was an X2.2 flare on 2011 February 15th, but at 01:56UT it was not recorded. There is also a bumper crop of C and M-class events, with 11 B-class recorded as well.

An M6.0 flare on the 3rd produced a very clear SID, as shown in this plot by Martyn Kinder:

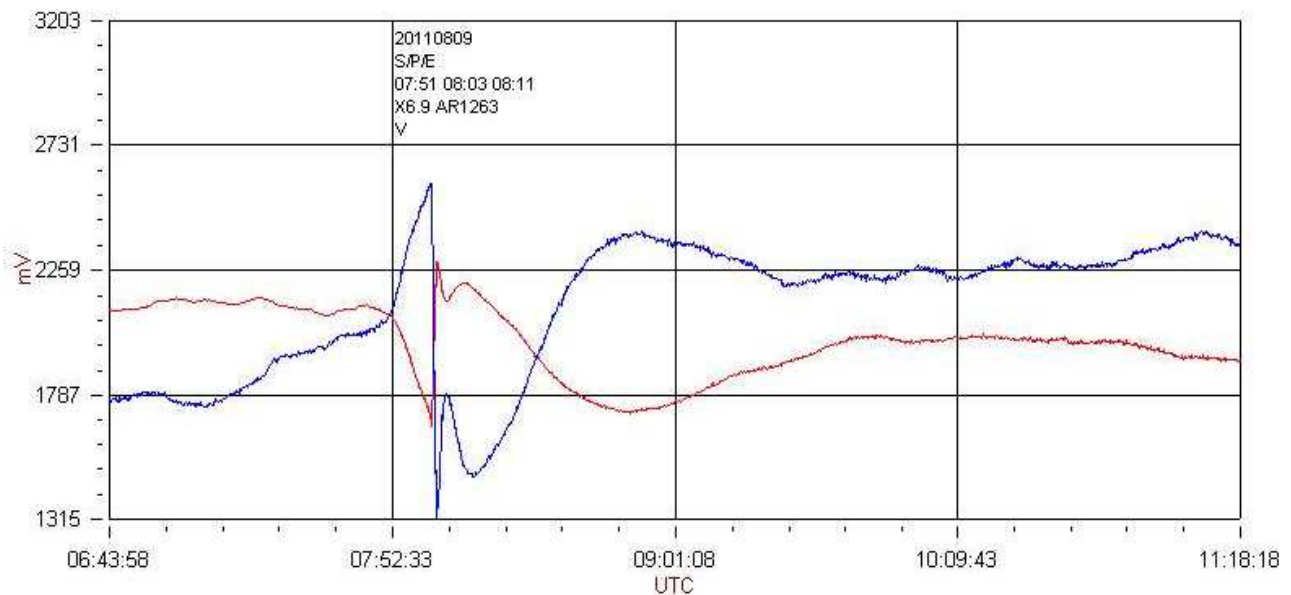


Mark Edwards also included this plot of the event, as recorded with spectrum lab software.

The start of the X6.9 flare on the 9th occurred just before the 23.4kHz signal re-appeared after its usual morning break. The SID was typical of those produced by energetic solar flares, in that the ground/sky wave interference pattern shifted a peak and trough over observers. The effect is shown in the recording made by John Wardle:

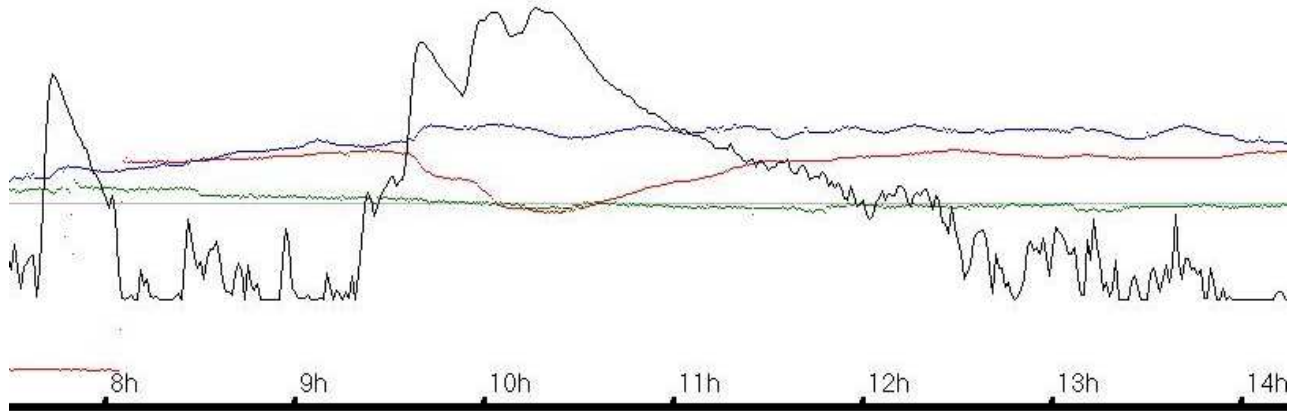


John is using the Starbase system for controlling his receivers and logging the results. The 'spike and wave' SID makes timing the event difficult, although thinking of the moving interference pattern helps. Martyn Kinder also sent a plot of this event:

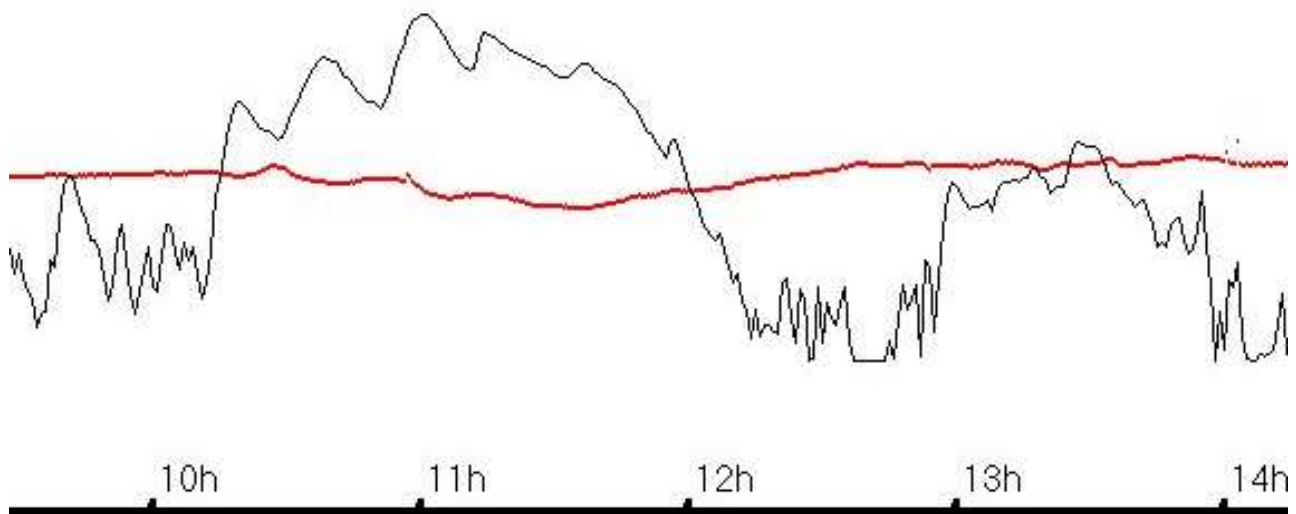


The C6.1 flare on the 11th was much smaller, but created a more complex SID. The background X-ray flux was already at a fairly high level by 09:30UT when the first of 3 peaks

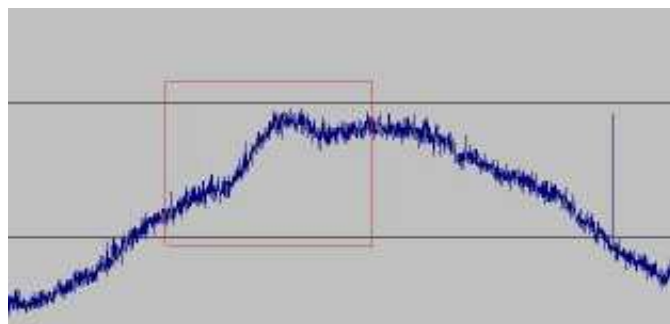
occurred. Two further peaks followed before the flux returned to lower levels again. My own recording shows 3 SIDs (red is 23.4kHz, blue is 22.1kHz):



The GOES data has been added in black to help clarify events. I have listed all 3 peaks in the activity summary to include all of the timings received, although the activity chart includes it as just a single flare.



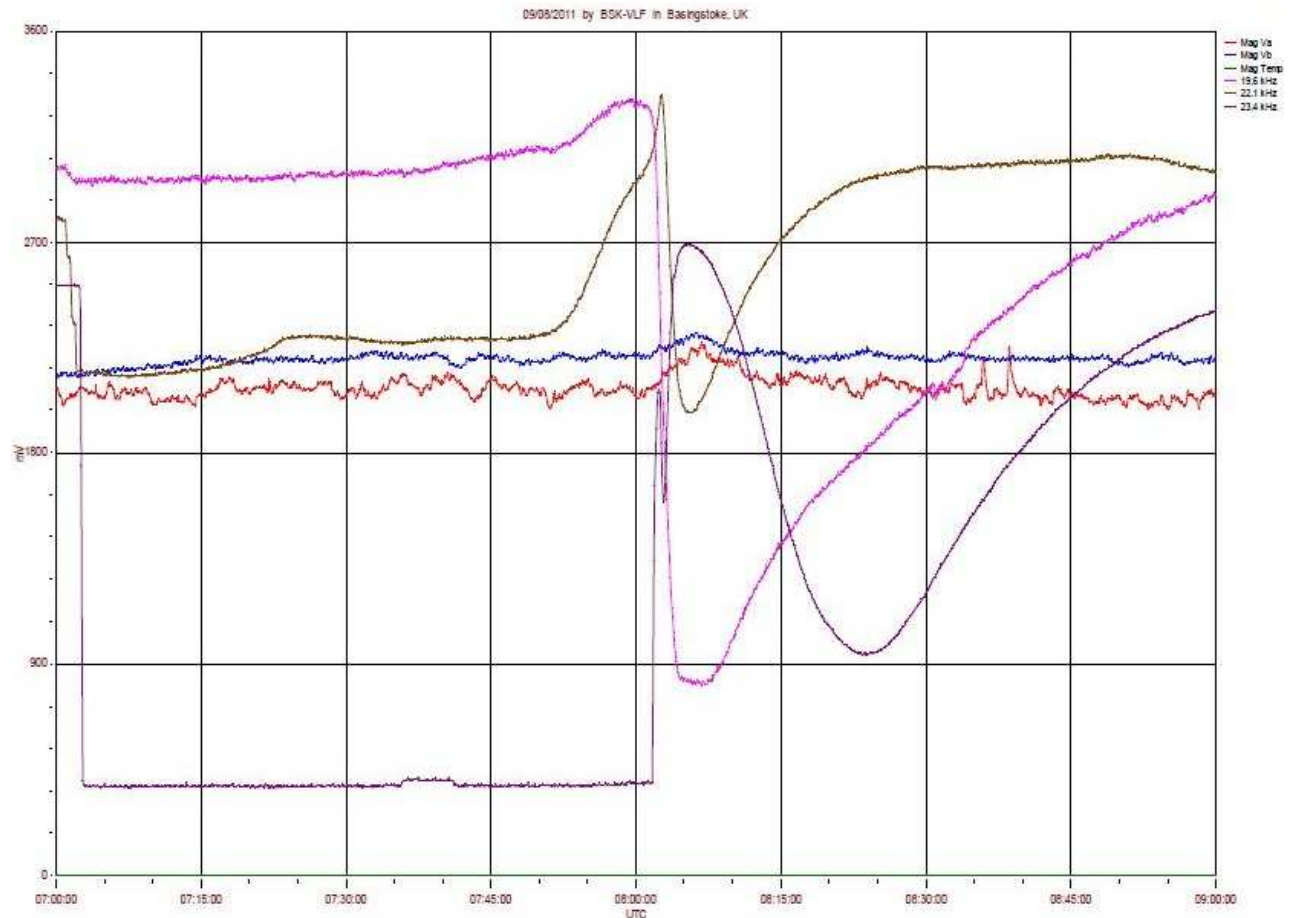
August 29th produced more similar events, shown above. At C2.9, I have barely recorded an identifiable SID for any of the six peaks shown in the GOES X-ray flux.



Mark Edwards noted that his signal at 24.0kHz was responding to some fairly weak B-class flares. The B3.6 on the 12th is probably the weakest recorded by the group so far, but evidence is provided in his recording shown above.

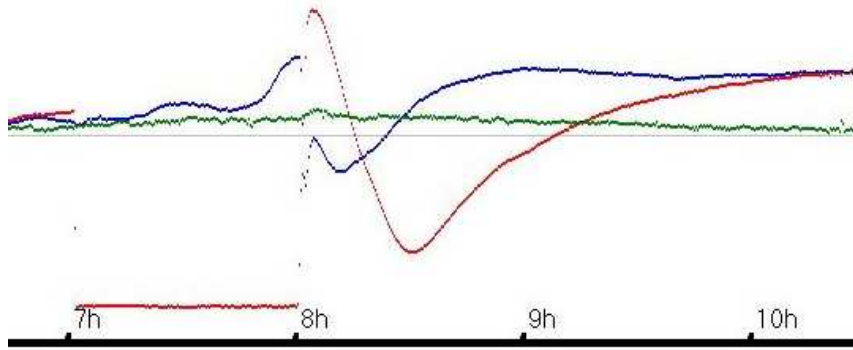
Magnetic Data

The SID from the X6.9 flare on the 9th was accompanied by a sudden flare event, or SFE. This is a magnetic disturbance occurring at the same time as the SID, due to the flare. Normally, the solar magnetic field associated with the flare would travel through space with its plasma at a much slower speed than the X-ray photons causing the SID. However, the large disturbance to the ionosphere will alter the large currents flowing in this region. The effect is that of a motor where a moving electric field (the ionosphere) will cause a changing magnetic field. This interacts with the Earth's own magnetic field to give a small disturbance. The effect is not always seen, but the chart from Paul Hyde shows it quite well:

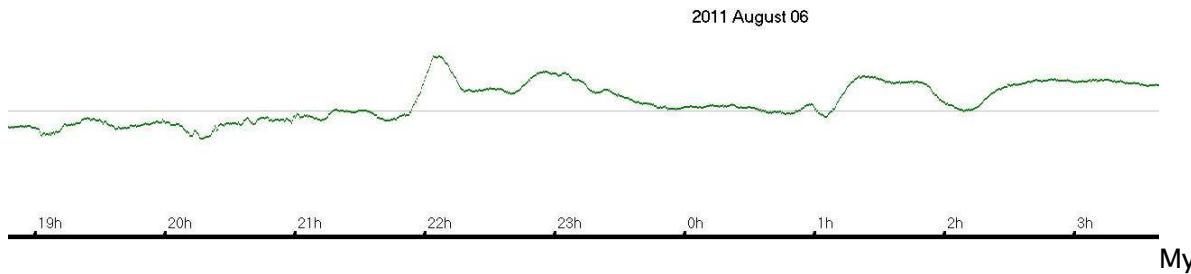


This sort of connection has been recorded previously by the Radio Group, but has so far remained unidentified. Paul drew my attention to it, and I found that I have very similar results. In Paul's chart, the brown and pink traces are the VLF SID, the red and blue traces being the two magnetic components. A small increase in both magnetic components can be seen, peaking around the maximum point of the SID. This is the time of maximum ionospheric disturbance, and thus also maximum magnetic disturbance from the moving electric current.

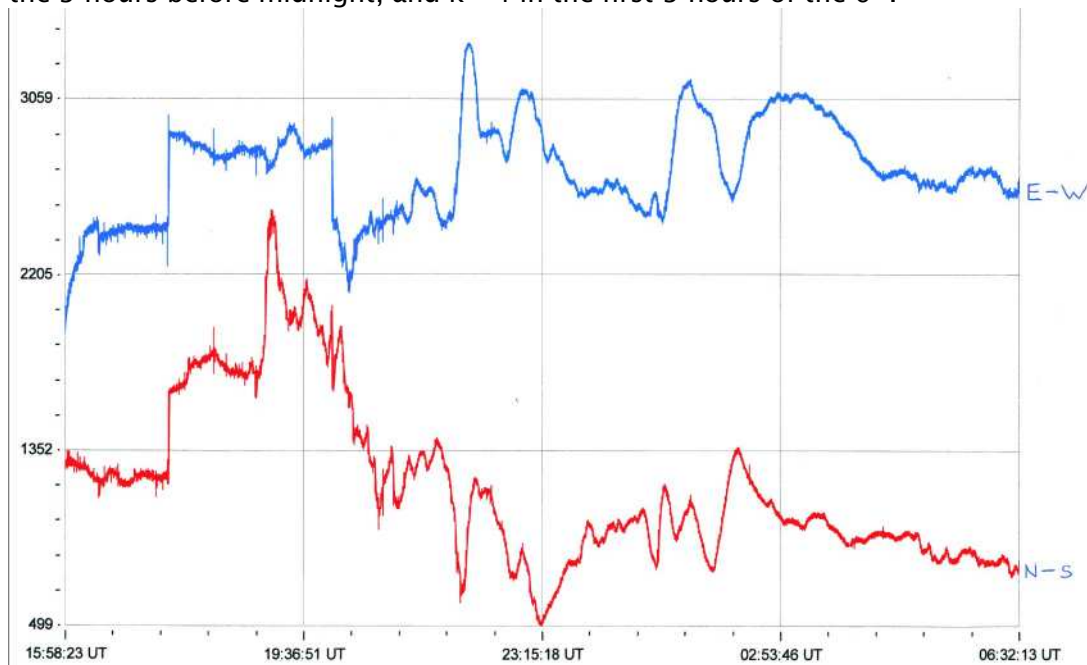
We will be keeping an eye open for similar events during solar cycle 24, so please send in any recordings of SFE candidates that you find. My own chart is on the next page, green being the magnetic signal, blue 22.1 kHz.



The vertical scale of my plot is rather compressed compared to Paul's, but the bump in the green magnetometer trace is clear enough. The active period shown on the 5th was a result of a CME produced by an M9.3 flare at 04UT in the 4th, and thus not shown in this summary. The remaining disturbances were due to coronal hole high-speed streams, mostly returning after a complete solar rotation.



magnetic recording of the CME arrival on the 5th/6th is shown above. I estimate a k-index of 5 in the 3 hours before midnight, and $k = 4$ in the first 3 hours of the 6th.



The recording by Colin Clements, above, shows good agreement in the East-west (blue) component, despite some local disturbance from about 16:15 to 20:00.

ROTATION	KEY:	DISTURBED.	ACTIVE	B, C, M, X = FLARE MAGNITUDE.	Synodic rotation start (carrington's)
2407	F	18 19 20 21 22 23 24 25 26 27 28 29 30 31			2092 1 2 3 4 5 6 7 8 9 10 11 12 13
2408	F	14 15 16 17 18 19 20 21 22 23 24 25 26 27 28 29 30			2093 2010 February 1 2 3 4 5 6 7 8 9 CC MCCMMC C
2409	F	10 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 26 27			2094 2010 March 28 1 2 3 4 5 6 7 8 CC CBM CC C
2410	F	9 10 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 26			2095 2010 April 27 28 29 30 31 1 2 3 4 CC B
2411	F	5 6 7 8 9 10 11 12 13 14 15 16 17 18 19 20 21 22			2096 23 24 25 26 27 28 29 30 May 1
2412	F	2 3 4 5 6 7 8 9 10 11 12 13 14 15 16 17 18 19 20			2097 21 22 23 24 25 26 27 28
2413	F	29 30 31 1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16 17			2098 18 19 20 21 22 23 24 C MCCC
2414	F	25 26 27 28 29 30 1 2 3 4 5 6 7 8 9 10 11 12 13			2099 14 15 16 17 18 19 20 21 C CC
2415	F	22 23 24 25 26 27 28 29 30 1 2 3 4 5 6 7 8 9 10			2100 11 12 13 14 15 16 17 18 C C C C
2416	F	19 20 21 22 23 24 25 26 27 28 29 30 31 1 2 3 4 5			2101 6 7 8 9 10 11 12 13 14 C
2417	F	15 16 17 18 19 20 21 22 23 24 25 26 27 28 29 30			2102 1 2 3 4 5 6 7 8 9 10 11 C B
2418	F	12 13 14 15 16 17 18 19 20 21 22 23 24 25 26 27 28 29 30			2103 31 1 2 3 4 5 6 7 C CC CM
2419	F	8 9 10 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 26			2104 27 28 29 30 1 2 3 4 C CC C
2420	F	5 6 7 8 9 10 11 12 13 14 15 16 17 18 19 20 21 22 23 24			2105 25 26 27 28 29 30 31
2421	F	1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16 17 18 19 20			2106 21 22 23 24 25 26 27 C C
2422	F	28 29 30 31 1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16 17			2107 18 19 20 21 22 23 24 C C C C C C C C
2423	F	24 25 26 27 28 29 30 31 1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16			2108 17 18 19 20 21 22 C C C C C C C C C C
2424	F	23 24 25 26 27 28 29 30 31 1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16			2109 17 18 19 20 21 22 C C C C C C C C C C
2425	F	19 20 21 22 23 24 25 26 27 28 29 30 31 1 2 3 4 5 6 7 8 9			2110 10 11 12 13 14 15 C C C C C C C C C C
2426	F	16 17 18 19 20 21 22 23 24 25 26 27 28 29 30 31 1 2 3 4 5 6 7 8 9			2111 3 4 5 6 7 8 C C C C C C C C C C
2427	F	12 13 14 15 16 17 18 19 20 21 22 23 24 25 26 27 28 29 30			2112 1 2 3 4 5 6 7 8 C C C C C C C C C C
2428	F	9 10 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 26 27 28 29			2113 30 31 1 2 3 4 C C C C C C C C C C
2429	F	5 6 7 8 9 10 11 12 13 14 15 16 17 18 19 20 21 22 23 24 25 26 27			2114 28 29 30 31 C C C C C C C C C C

Magnetic data from Paul Hyde, Colin Clements, Gonzalo Vargas & John Cook.

Radio Astronomy Group General Meeting.

We will be holding a Radio Group meeting on Saturday, November 12th, in Northampton. All are invited to attend. There will be two professional speakers, talking about UK involvement in the square kilometer array and the LOFAR telescopes. Radio group members will be talking about various radio sources, software defined radio for the amateur, magnetic and solar observations, etc. There will be a charge of £12 for BAA members (£15 non-members) to include lunch, tea and coffee, or £7 (£10 non-members) without lunch. Full details will be found shortly at www.britastro.org/radio Early booking is advised, so that we can gauge seating arrangements, etc. Previous meetings have been well received. I look forward to the opportunity to see some of you there. The venue (Humfrey rooms, Castilian Terrace) is close to the town centre, and 10 minutes walk from the railway station.