

A Guide to AstrolmageJ Differential Photometry

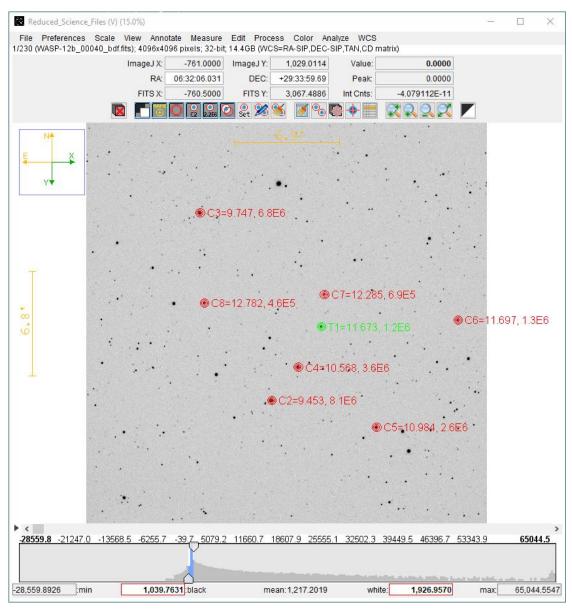


Image Display Interface with WASP-12b Target and Comparison Aperture overlay

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1. Introduction

The BAA VSS Guide to AstrolmageJ Photometry covers:

- Image reduction and photometry processing in AstroImageJ (AIJ)
- Steps in creating BAA VSS Report files from an AIJ measurements data file
- Demonstrate importing AIJ RA/Dec coordinate-based apertures into plate solved images

Chapter 10 in the AstroImageJ User Guide is a step-by-step guide to differential photometry in AstroImageJ (AIJ) based on the WASP-12b example files and is the first step in learning to use the software. The WASP-12b FITS files can be downloaded from the AIJ web site. The chapter is self-contained and can be followed without reading the entire User Guide.

This guide follows on from my article Introduction to AstroImageJ (AIJ) in the <u>June 2018 VSS Circular</u>. The aim is to assist BAA VSS members with some familiarity with AIJ in creating VSS report files using the example WASP-12b FITS files. The guide covers image reduction, importing photometry results and creating a BAA VSS report file using the VSS Photometry spreadsheet.

1.1. Software Summary

Software	Version	Notes
AstrolmageJ	db3.2.1	All User Guide is the main software reference; there are also a number of on-line tutorials
VSS_Photometry	V2.05 (or later)	Excel spreadsheet converts AIP4WIN and AstroImageJ photometry results to BAA VSS Report files.
WASP-12b FITS files	N/A	4GB+ download of raw image and calibration FITS files. From FITS header, these images were exposed with a red filter.

AlJ performs image reduction, photometry processing and calculation of target ensemble magnitude. Results are saved to a text file, *measurements.txt*. Measurements.txt is imported into VSS_Photometry for conversion to VSS Report format.

Software, example files and documentation can be downloaded from the AstroImageJ web page http://www.astro.louisville.edu/software/astroimagej/

1.2. Comments on AIJ Tutorial Guide

- 1. The VSS_Photometry spreadsheet was designed for AIP4WIN photometry files and recently extended to import AIJ data. For the user, one obvious difference is how the two programs label the comparison star apertures, starting with C1 in AIP4WIN and with C2 in AIJ. For example, AIP4WIN would label four comparison apertures in sequence C1, C2, C3 and C4 while in AIJ the target aperture is labelled T1 then C2, C3, C4 and C5. VSS_Photometry follows the AIP4WIN format and in the Results sheet Comparisons are labelled C1, C2 ...
- 2. The aperture file used in this guide, radec.txt, lists J2000 coordinates and Comparison star Johnson V band magnitudes. These values were downloaded from the AAVSO VSP for WASP-12. The WASP-12b images were exposed with a red filter and, because of filter mismatch, the Results sheet in VSS_Photometry shows large (Calc-Quote) differences in the Data Checks section. In section 5, a second set of results for Landolt field SA113, exposed with a Johnson-V filter, report 'OK' Magnitude Checks for each Comparison.

2. Overview

Fig 2.1 illustrates the exchange of time-resolved photometry data between AIJ and VSS_Photometry.

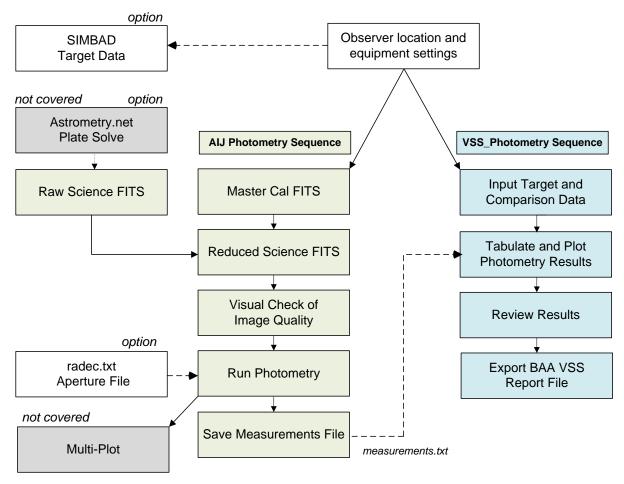


Fig 2.1:Overview of AstrolmageJ + VSS_Photometry Analysis Sequence

Observer settings are covered in Appendix B. The main text walks the user through each stage from initialising target data through to creating a BAA-VSS format report file.

Plate solve option: Chapter 7 of the <u>AstroImageJ Users Guide</u> covers set-up and running Astrometry.net to plate solve FITS images through AIJ. The aperture file *radec.txt* comprises a list of target and comparison star RA/Dec coordinates. Importing these coordinates into AIJ enables the software to place apertures on a plate-solved image.

The WASP-12b example files are already plate solved so this step is not required to follow this guide.

Manual aperture option: Refer to the AstroImageJ User Guide and other tutorials for placing apertures on star images based on finder charts.

AlJ photometry results are saved in text file *measurements.txt*. VSS_photometry imports measurements.txt, computes target magnitude and exports a BAA VSS format results file.

While not required for BAA VSS photometry, AIJ also computes heliocentric (HJD) and barycentric (BJD) times taking DP Coordinate Converter observer and target parameters (see section 3.2 and Appendix B for setting up DP Coordinate Converter data).

3. Science Image Reduction

3.1. Software Setup

Setup directories and configure AIJ to follow this guide as detailed in the following appendices:

Appendix A: recommended directory set-up for example files.

Appendix B: configure AIJ and VSS_Photometry for Moore Observatory (University of Louisville) and equipment parameters

3.2. SIMBAD Coordinates

This section illustrates importing SIMBAD coordinates into the DP Coordinate Convert dialog.

- 1. Open AIJ and from the Toolbar and click pt to open the CCD Data Processor and DP Coordinate Converter windows.
- 2. In DP Coordinate Converter window Fig 3.1, enter "WASP-12b" in the SIMBAD Object ID text box and press <Enter> to populate Standard Coordinates fields.

Optional: click the SIMBAD icon to open the SIMBAD web page for WASP-12b (Extra-solar Confirmed Planet)

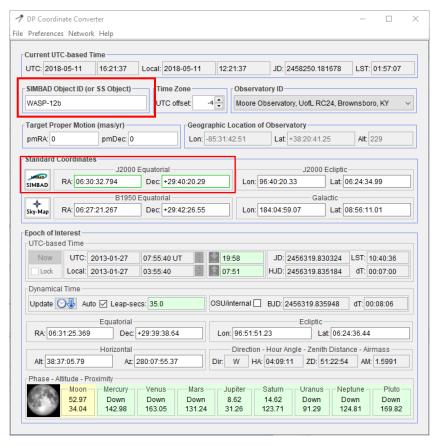


Fig 3.1: WASP-12b Coordinates in DP Coordinate Converter

3.3. Aperture File

Create radec.txt aperture file and save in WASP-12b folder as detailed in Appendix C.

3.4. Reduce Science FITS Files

Image reduction steps in AIJ are:

- (i) build master calibration files and
- (ii) run science image reduction process
- Select the AIJ / CCD Data Processor window and configure for building master calibration files as fig 3.2

Bias / Dark / Flat: Build and Enable selected, med (median) combine

Dark: scale and deBias selected to enable exposure time scaling

Note: the Science Image Processing Enable checkbox is not selected

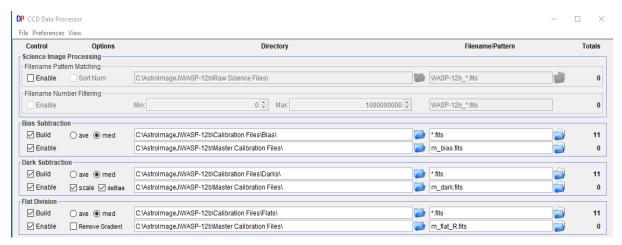


Fig 3.2: CCD Data Processor Master Calibration File Settings
Disable Science Image Processing

2. To import raw calibration files, click the folder icon and navigate to the respective calibration file directory, e.g. C:\AstroImageJ\WASP-12b\Calibration Files\Bias\, etc.

A simple file pattern (*.fits) is sufficient if each directory contains only single file type.

Fig 3.2 shows 11 bias, dark and flat FITS files.

3. To specify master calibration files, select the master calibration directory e.g.

C:\AstroImageJ\WASP-12b\Master Calibration Files\.

In Fig 3.2 the master calibration filenames are m bias.fits, m dark.fits and m flat R.fits.

Totals = 0 indicates that no master calibration files have been saved yet.

Refer section 6.5 AIJ Users Guide for details of AIJ Data Reduction Algorithms

4. Click [START] to initiate processing of calibration files; a Log window opens detailing offset and scaling corrections.

When the process is finished, totals = 1 indicates that the master calibration files have been saved.

5. Deselect Bias, Dark and Flat Build check boxes and select the Enable checkbox in Science Image Processing section. Specify the raw image file directory and matching pattern, Fig 3.3.

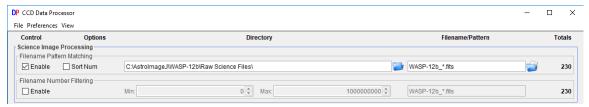


Fig 3.3: Enable Science Image Processing

6. Enable Save Calibrated Images and select 32 bit option.

Enter "Reduced_Science_Images" and "_bdf" in the Sub-dir and Suffix textboxes respectively, Fig

3.4.

NOTE: Option to select Plate Solve (on-line Astrometry.net plate solve, refer chapter 7 in AIJ Users Guide). Not required for WASP-12b files which are already plate solved.

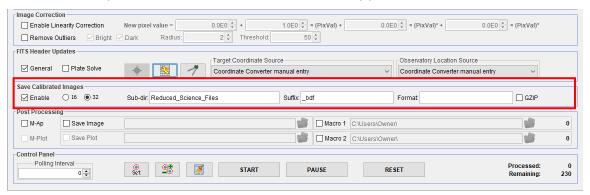


Fig 3.4: Calibrated image options

- 7. Click [START] to commence processing of science images. The AIJ Image Display interface opens with the first science image then cycles through each image. The Processed and Remaining counts in the CCD DP dialog monitor progress and a Log window documents the image reduction process (several minutes for WASP-12b files).
- 8. When processing science images is finished, close AIJ and move the reduced image files:

From: C:\AstroImageJ\WASP-12b\Raw Sciences Files\Reduced_Science_Files

To: C:\AstroImageJ\WASP-12b\Reduced_Science_Files

3.5. Referencing Target File Path

As an alternative to specifying full file paths, the Bias, Dark and Flat directories can be referenced to the Science Image Directory detailed in the tooltip in fig 3.5.

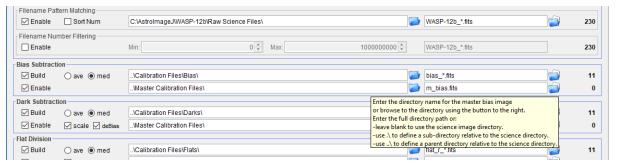


Fig 3.5: Referencing Calibration Directories

4. AIJ Photometry

4.1. Configure Measurement Apertures

- 1. Open AIJ and from the AIJ Toolbar, import image sequence *Toolbar | File | Import | Image Sequence* .. navigate to the Reduced_Science_Folder and select the first file (WASP-12b_00040_bdf.fits).
- The Sequence Options dialog opens, select Sort names numerically and Use virtual stack options.
 Click [OK] to open the first reduced science image in the AIJ Image Display Interface, Fig 4.1.
 The yellow N-E pointers indicate this is a plate-solved image.
 Refer to chapter 5 in the AIJ Users Guide for a detailed description of available functions.

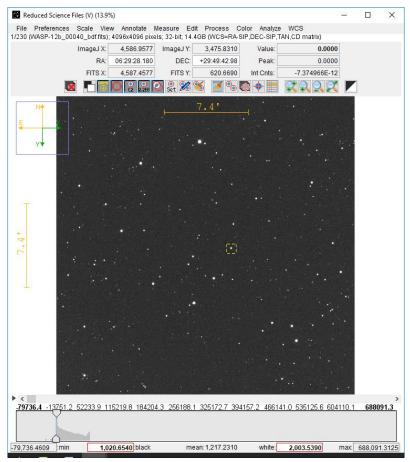
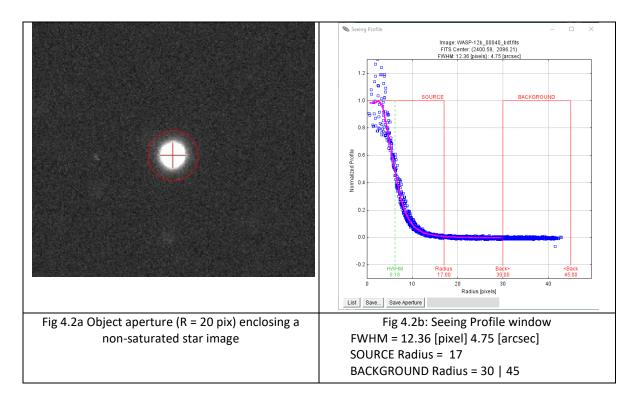


Fig 4.1: Image Display Interface: WASP-12b_00040_bdf.fits

- 3. Option: select View and the appropriate *Invert.*. option so that N-E is orientated up and left as Fig 4.1.
- 4. Initial set-up of apertures sizes to measure Seeing Profile:
 Click change aperture settings to open the Aperture Photometry Settings window.
 Enter object / inner / outer radius = 20 / 24 / 30 in and [OK] to close the window.
 Although the exact values are not critical, the object aperture must be large enough to fully enclose the image of a single star.
- 5. Click toggle aperture display as necessary to *de-select* the sky background apertures so that the "Live Photometer" is a single circle overlay as shown in Fig 4.2a

6. Alt-Left click on a bright non-saturated star image – the star highlighted in Fig 4.1 is a good candidate. The Seeing Profile for the selected star opens in a new window, Fig 4.2b



7. Click [Save Aperture] in the Seeing Profile window to select the suggested aperture sizes then click in the Image Display window to confirm as shown in Fig 4.3.
Optionally, edit Seeing Profile generated aperture radius values to preferred values relative to FWHM ≈ 12.5 pixel.

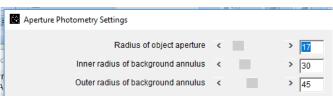


Fig 4.3: Final radius values

8. Click [OK] to close Aperture Photometry Settings window, click clear apertures .. ** to clear any aperture overlays and click ** to toggle display of sky apertures.

4.2. Visual Check of Image Quality

The image stack can be inspected either manually stepping through the stack or by running an animation.

- 1. Use the scroll bar controls highlighted in Fig 4.4 to inspect individual images.
- 2. Click on control left of the scroll bar to start / pause image animation. Right click on this control to set animation speed and other options.

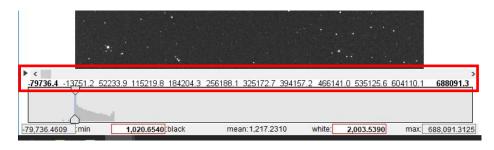


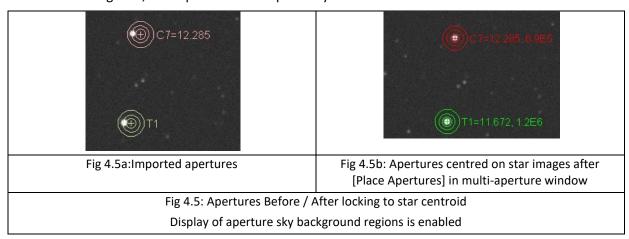
Fig 4.4: Image inspection controls

3. To remove an image from the stack, click , left-most icon.

Note: this function only removes the image from current photometry analysis sequence, the FITS file is *not* deleted from the Reduced_Science_Files directory.

4.3. Run AIJ Photometry

1. In the Image Display window, select *File | Import apertures from RA / Dec list* .. , navigate to and select radec.txt file to import the aperture set. Click zoom to fit .. at to re-centre the image. As shown in Fig 4.5a, initial placement will probably be offset from star centroids.



- 2. Open the Multi-Aperture Measurements window (in Image Display), click [Place Apertures] to centre the apertures on star centroids (Fig 4.5b); the Multi-Aperture Help dialog opens.
- 3. Press <Enter> to start AIJ time resolved photometry processing. Several windows open, including:

Plot of Measurements: can be configured to plot relative flux and other parameters during processing. Refer section 10.2.1.10 in the AIJ Users Guide for more details about Multi-plot. Measurements window: displays a table of photometry results.

NOTE: Apertures track frame-to-frame shifts in star centroids including image rotation after a meridian flip.

- 4. When processing is complete, select *Measurements | File | Save As ..* navigate to WASP-12b director and save as "Measurements.txt", replacing existing file.
- 5. Close AIJ

5. Processing Results in VSS_Photometry

5.1. Setup VSS_Photometry Spreadsheet

- 1. Open VSS Photometry.XLSM and select the Buttons tab.
- 2. Click [Select Photometry Import File] button, navigate to and select measurements.txt in the WASP-12b directory.
 - The software automatically detects the photometry file format and the Type field displays: 'AstroImageJ'.
- 3. Click [Select Export Directory] button, navigate to and select the WASP-12b directory.
- In BAA VSS report file name textbox, enter "BAAVSS WASP-12b.txt".
 Fig 5.1 shows the Buttons sheet with WASP-12b directory and file names, including AAVSO report file.

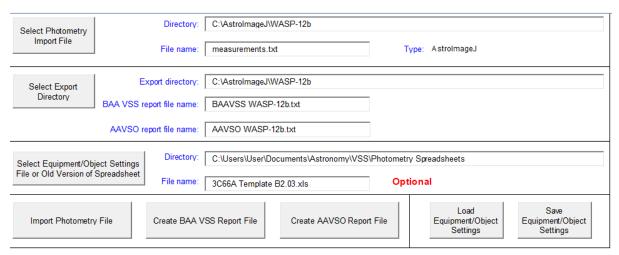


Fig 5.1: Buttons sheet configured for WASP-12b

5. Select the Results sheet, clear previous entries and enter *red-text* data as shown in Fig 5.2. Star Designation and Reference Magnitude values taken from *radec.txt*, see Appendix C.

Observation	n Summary		mns W and X explanation		Red - Requ	iired, Blue -	Optional,
Chart	AAVSO WAS	P-12b	Star Apertu	re			Air
Filter	V		Sky Annulu Inner Radius				Observa Block
Photometry Time Stamp	Start Expos	sure	Sky Annulu Outer Radiu				Julian D
Timing Error	1.00		Automatic populated		<u>'</u>	Check 1	
Star Data & Su	ımmary Calcula	ations	(Individual V	/aria	able Measu	_ /	
			User Inpu	ut	/	/	
	Star		AIP4WIN		Include in	Refere	nce
Туре	Designation		Star	*		Magnitude	Error
AAVSO Results	WASP-12b		Var				
BAAVSS Results	WASP-12b		Var				
Comparison	000-BKG-164		C1	ΠÌ	Yes	9.453	
Comparison	000-BKG-165		C2	_	Yes	9.747	
Comparison	000-BKG-166		C3		Yes	10.568	
Comparison	000-BKG-420		C4		Yes	10.984	
Comparison	000-BKG-421		C5		Yes	11.697	
Comparison	000-BKG-167		C6		Yes	12.285	
Comparison	000-BKG-168		C7		Yes	12.782	
Comparison							

Fig 5.2: Results sheet: Header and Comparison data

5.2. Import Results in VSS_Photometry

1. Select the Buttons sheet, click [Import Photometry File] button and close dialogs when data import is finished.

2. Select Results sheet Fig 5.3

Note: The large Calc-Quoted Differences (red coloured cells) resulting from a mismatch between V-band Reference magnitudes and WASP-12b imaged with red filter (see comments in section

1.2). Refer to Appendix F for an example of reference star photometry results

Star Data & S	ummary Calculations	(Individual Variable Measurements to Right)			delete not space)			DAMANOO SIIRRI AINI AMANO		
		User Input	1	1	/		Results			Data Checks
	Star	AIP4WIN	Include in	Referenc	:e	Average	Average	Std Dev	Average	Calc - Quoted
Туре	Designation	Star	Analysis	Magnitude	Error `	Magnitude	Error	Magnitude	Weight	Difference 1
AAVSO Results	WASP-12b	Var				11.683	0.041	0.009		
BAAVSS Results	WASP-12b	Var				11.683	0.041	0.009		
Comparison	000-BKG-164	C1	Yes	9.453		9.556	0.041	0.002	0.395	0.103
Comparison	000-BKG-165	C2	Yes	9.747		9.739	0.041	0.004	0.323	-0.008
Comparison	000-BKG-166	C3	Yes	10.568		10.464	0.041	0.003	0.139	-0.104
Comparison	000-BKG-420	C4	Yes	10.984		10.791	0.041	0.003	0.092	-0.193
Comparison	000-BKG-421	C5	Yes	11.697		11.539	0.041	0.004	0.033	-0.158
Comparison	000-BKG-167	C6	Yes	12.285		12.257	0.041	0.004	0.011	-0.028
Comparison	000-BKG-168	C7	Yes	12.782		12.697	0.041	0.005	0.006	-0.085

Fig 5.3: Results sheet for WASP-12b photometry V-band reference / R-filter exposures

- 3. To remove a comparison star from photometry analysis, click on the 'Yes' entry in the 'Include in Analysis' column and select 'No' from the drop down options
- 4. To select a Check star for submitting AAVSO report, click on 'Yes' entry then select 'Check'. In Fig 5.4, 000-BKG-420 is de-selected from analysis and 000-BKG-421 selected as the AAVSO check star.

Star Data & Su	ummary Calculations	(Individual Variable Measurements to Right)					
		User Input	/	1			
	Star	AIP4WIN	Include in	Reference			
Туре	Designation	Star	Analysis	Magnitude Error			
AAVSO Results	WASP-12b	Var					
BAAVSS Results	WASP-12b	Var					
Comparison	000-BKG-164	C1	Yes	9.452			
Comparison	000-BKG-165	C2	Yes	9.747			
Comparison	000-BKG-166	C3	Yes	10.568			
Comparison	000-BKG-420	C4	No	10.984			
Comparison	000-BKG-421	C5	Check	▼ 11.697			
Comparison	000-BKG-167	C6	Yes	12.285			
Comparison	000-BKG-168	C7	Yes	12.782			

Fig 5.4: Results sheet Yes / No / Check selections

5.3. Export BAA VSS or AAVSO Report Files

To create a BAA VSS formatted report file, select the Buttons sheet and click [Create BAA VSS Report File] button to save e.g. ..\WASP-12b\BAAVSS WASP-12b.txt.

To create an AAVSO formatted report file, select the Buttons sheet and click [Create AAVSO Report File] button to save e.g. ..\WASP-12b\AAVSO WASP-12b.txt.

6. Summary

Analysing time series photometry images in AstroImageJ is relatively straightforward: the more complex software features, including multi-plot, de-trending data and multivariate transit fit analysis, are not required for creating VSS Report files.

In addition to variable star photometry, other cases where BAA members might consider using AstroImageJ include:

- For observers who are new to stellar photometry, AstroImageJ is free to down-load with several sets of example data on the AIJ home page and other sites.
- AAVSO recommends AstroImageJ for exoplanet photometry. Chapter 7 in <u>A Practical Guide to Exoplanet Observing</u> (Dennis Conti) describes using AIJ for this application. This chapter also works with a WASP-12b data set and covers Multi-plot functions in detail.
- Asteroid light curve photometry: in the Multi-plot Main window, select File | Create Minor Planet Centre (MPC) format .. to create MPC formatted data.
- More experienced observers might consider AstroImageJ as an alternative to AIP4WIN, for example. The user forum is well supported by Karena Collins at the University of Louisville. AstroImageJ release notes are listed here.

Users familiar with other photometry software should find the transition to AIJ fairly straightforward. All users are encouraged to check out the AIJ website referenced in the Introduction section for a functions list, tutorials and link to the AIJ forum.

Appendix A: Example File Directories

After downloading WASP-12b raw image and calibration example files (over 4GB) set up directories as Fig A1.1.

Create and save a blank text file measurements.txt in the WASP-12b folder.

radec.txt is the WASP-12b aperture file used in this guide, see Appendix C.

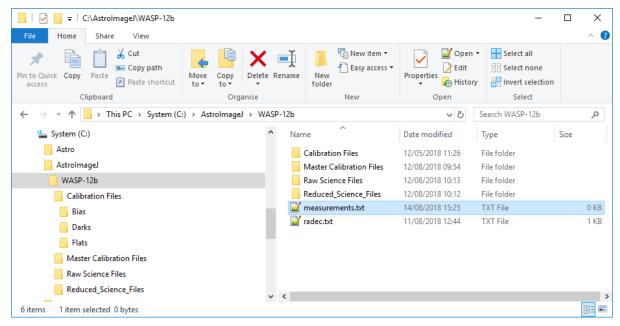


Fig A1.1: WASP-12b directories used in text

Appendix B: Configure Software

Set up observer location and equipment parameters in AIJ and VSS_Photometry, and configure multi-aperture settings in AIJ.

NOTE: Boxed sections detail important instructions for configuring AstrolmageJ software.

- 1. Open VSS_Photometry and select ObsvEqmt sheet
- 2. To follow the tutorial using WASP-12b example files, input University of Louisville into Observer & Equipment fields shown in Fig B1.1 and save changes:

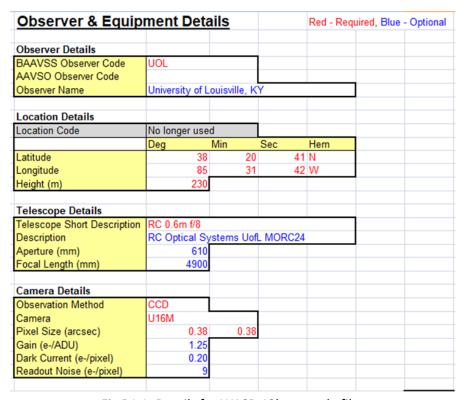


Fig B1.1: Details for WASP-12b example files

- 3. Open AIJ and from the Toolbar click DP to open the DP Coordinate Converter window.
- 4. In the DP CC window, click Preferences and select Use custom observatories list (...). Close AIJ then re-open AIJ and the AIJ / DP CC window
- 5. Click on Observatory ID and select Moore Observatory, UofL, RC24, Brownsboro, KY, Fig B1.1a.
- 6. Informative only: Edit Time Zone / UTC offset for Eastern Standard Time.

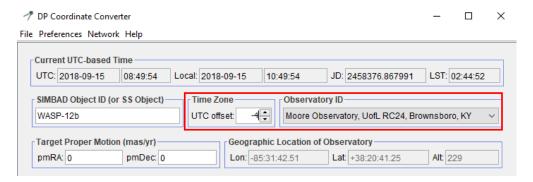


Fig B1.1a: Observatory ID and Time Zone

- 7. To set up a custom observatory, navigate to the AIJ home directory and edit *observatories.txt* to add a new entry formatted as specified in the file header. Close observatories.txt and save changes. The new observatory should now be available from the Observatory ID drop-down.
- 8. Select the CCD Data Processor window. In the FITS Header Updates section, select options:

Target Coordinate Source: Coordinate Converter manual entry
Observatory Location Source: Coordinate Converter manual entry



Fig B1.2: Target and Observatory data imported from DP CC settings

9. Click to open the General FITS Header Settings panel Fig B1.3

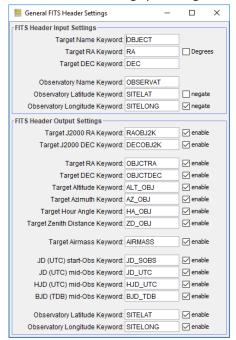


Fig B1.3: Setup FITS Header Keywords

10. Enable all the FITS Header Output keywords as shown. The important keywords are Airmass and JD-based times, the others are informative.

The FITS Header Input Settings section is not used; Target and Observatory details are imported from the DP Coordinate Converter window

11. Close the FITS Header Settings dialog.

12. Click in the Control Panel row (bottom row) of the CCD DP window to open the first of two rather busy dialog windows, Fig B1.4.

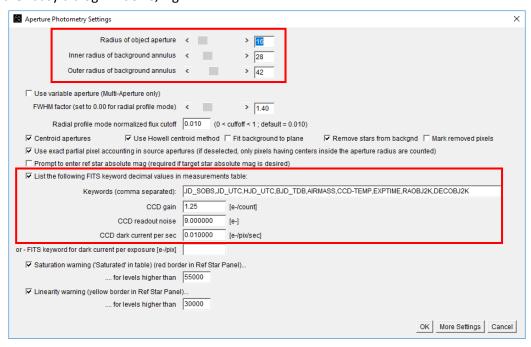
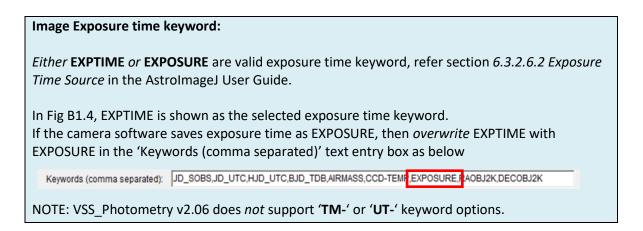


Fig B1.4: Page 1 of 2 Aperture Photometry Settings



- 13. Enter values for CCD gain, CCD readout noise and pixel dark current (e-/pix/sec) listed below, estimate values for the U16M camera used for WASP-12b test images.
- 14. Check FITS keywords in measurements table include:
 - JD_SOBS, JD_UTC,HJD_UTC,BJD_TDB,AIRMASS,CCD-TEMP,EXPTIME (*or* EXPOSURE), RAOBJ2K,DECOBJ2K
- 15. Confirm check boxes are selected as indicated

Other settings:

Aperture radii settings in the upper section, these are set in main text.

Below that are a number of options with default selections.

Keywords to include in FITS header correspond to the General FITS Header settings dialog.

- 16. Click [More Settings] to open the second window. All except the last item are selected (Clear overlay before use).
- 17. Close the Aperture Photometry Settings dialog.

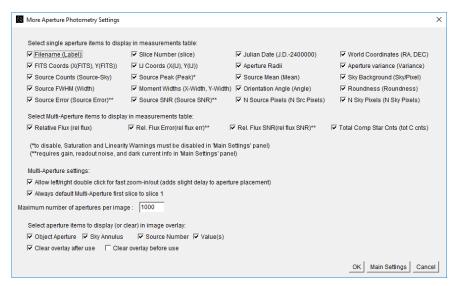


Fig B1.5: More Aperture Photometry Settings dialog Note: Single and Multi-Aperture selections (24)

Appendix C: Target and Comparison Aperture File

Copy / Paste box text into a text editor and save as radec.txt in WASP-12b folder.

```
#WASP-12b, 06:30:32, +29:40:20, 0, 1, 99.999
#000-BKG-164, 06:30:47, +29:35:30, 1, 1, 9.453
#000-BKG-165, 06:31:09, +29:47:47, 1, 1, 9.747
#000-BKG-166, 06:30:39, +29:37:40, 1, 1, 10.568
#000-BKK-420, 06:30:16, +29:33:45, 1, 1, 10.984
#000-BKK-421, 06:29:51, +29:40:46, 1, 1, 11.697
#000-BKG-167, 06:30:31, +29:42:27, 1, 1, 12.285
#000-BKG-168, 06:31:08, +29:41:53, 1, 1, 12.782
#RA, Dec, Ref Star, Centroid, Magnitude
06:30:32, +29:40:20, 0, 1, 99.999
06:30:47, +29:35:30, 1, 1, 9.453
06:31:09, +29:47:47, 1, 1, 9.747
06:30:39, +29:37:40, 1, 1, 10.568
06:30:16, +29:33:45, 1, 1, 10.984
06:29:51, +29:40:46, 1, 1, 11.697
06:30:31, +29:42:27, 1, 1, 12.285
06:31:08, +29:41:53, 1, 1, 12.782
```

File format:

1st character : '#' => comment line ignored by AIJ

RA, Dec: J2000 coordinates in HH:MM:SS and +/-dd:mm:ss formats

Ref Star: 0 => Target, 1 => Comparison

Centroid: 1 => centre aperture on star centroid

Magnitude: 99.999 => no magnitude (target star)

otherwise catalog magnitude (comparison star)

Appendix D: Required FITS Header Terms

The AIJ results file Measurements.txt must contain the following fields. If any are missing then attempts to import measurements data will be aborted with appropriate error message:

	Field	Notes
1 EXPTIME / EXPOSURE Exposure time in second		Exposure time in seconds. [1]
2	JD_SOBS	Julian Date at start of exposure
3	JD_UTC	Julian Date at mid-exposure
4	Label	FITS filename
5	N_Sky_Pixels_T1	No. pixels in T1 sky aperture (outer annulus)
6	N_Src_Pixels_T1	No. Pixels in T1 source aperture (centre aperture)
7	Sky/Pixel_T1	Sky background flux / pixel
8	Sky_Rad(Max)	Sky annulus outer radius (pixel)
9	Sky_Rad(Min)	Sky annulus inner radius (pixel)
10	slice	Image stack index
11	Source_Radius	Source aperture radius (pixel)
12	Source_SNR_C2	Computed SNR for first comp star
13	Source_SNR_T1	Computed SNR for target star
14	Source-Sky_C2	Sky-corrected flux for 1 st comp star
15	Source-Sky_T1	Sky-corrected target flux

[1] EXPTIME or EXPOSURE are valid FITS header fields for AstroImageJ data processing; VSS_Photometry v2.06 or later can import data containing either keyword.

Appendix E: AstroImageJ Windows Installation Packages

Link to AIJ installation packages (if this link fails, then an internet search for "AstroImageJ" should locate the home page and download links):

(https://www.astro.louisville.edu/software/astroimagej/installation_packages/

NOTE: Boxed sections detail important instructions for installing AstroImageJ software.

Index of /software/astroimagej/installation_packages

	Name	Last modified	Size	Description
	Parent Directory		-	
	AstroImageJ installation linux.html	2014-12-07 01:05	39K	
	AstroImageJ installation mac.html	2017-01-20 12:15	5.3K	
	AstroImageJ installation windows.html	2014-12-07 01:23	54K	
D	AstroImageJ script.tar.gz	2014-12-07 18:05	1.2K	
	AstroImageJ user guide 2.0 partial.pdf	2014-12-07 17:50	1.9M	
D	AstroImageJ v3.2.0 20160201 linux.tar.gz	2016-02-01 04:16	7.7M	
	AstroImageJ v3.2.0 20160201 mac.zip	2016-02-01 02:30	10M	
D	AstroImageJ v3.2.0 20160201 windows.zip	2016-02-01 03:52	7.9M	
Ü	AstroImageJ v3.2.0 20170222 windows java7 x32.zip	2017-02-22 05:41	51M	
D	AstroImageJ v3.2.0 20170222 windows java7 x64.zip	2017-02-22 05:39	54M	
?	Xresources	2014-12-07 17:49	412	

Fig E1: AstroImageJ Installation Packages

The highlighted links will download either Windows 32 or 64 bit versions plus a copy of the Java runtime engine. The next figure shows a typical installation:

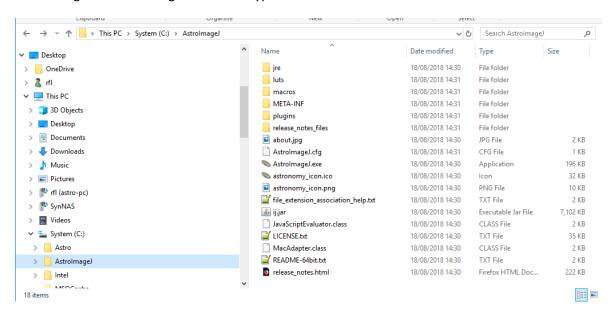
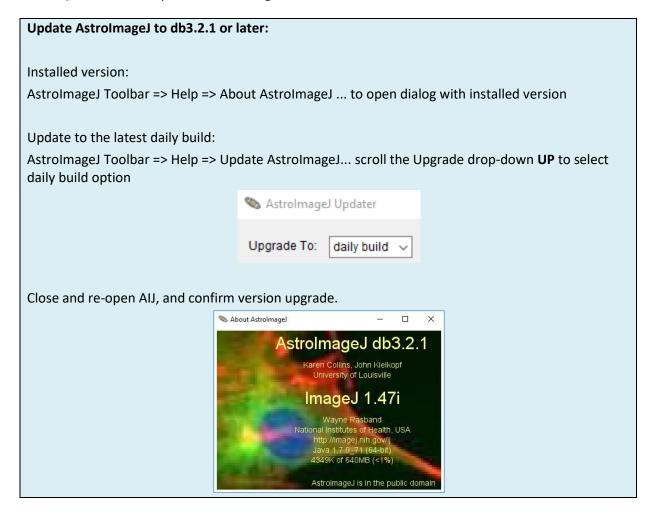


Fig E2: AIJ Installation

Folder jre contains the java runtime engine. File AstroImageJ.cfg is a short text file which is created when AIJ is run for the first time. A Windows Defender Firewall dialog may open, select [Allow access], close and re-open the AstroImageJ toolbar.



Link to the WASP-12b example files; Windows users can download the large *WASP-12b example uncalibrated images.zip* file (4.4GB)

https://www.astro.louisville.edu/software/astroimagej/examples/

Note: AlJ can run from a USB memory stick configured with the installation files and directories in Fig E2

Appendix F: V-Band Photometry of Landolt Reference Stars

Fig F1 shows the results of an imaging series taken with a V-Johnson filter. The Calc-Quoted Differences are less than 0.015 mag and all Magnitude Checks are flagged 'OK'.

The SA113 objects are Landolt reference stars with mmag uncertainties in catalog V-band magnitudes.

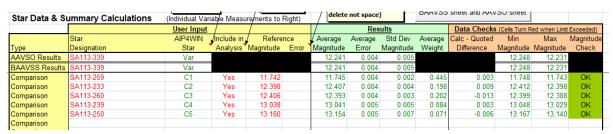


Fig F1: Results sheet for SA113 photometry V-band reference / V-filter exposures

Appendix Z: Document Change Record

Revision Date	Notes
15 Sept 2018	Initial release
20 Oct 2018	Changes to Appendices to assist installing AstroImageJ, updating to db 3.2.1 and Aperture settings required to import data into the VSS_Photometry spreadsheet. Key instructions are highlighted in blue-shaded text boxes, including EXPTIME / EXPOSURE conversion implemented in VSS_Photometry v2.06 Extended list of FITS header terms in Appendix D
	Extended list of 1113 fleader terms in Appendix D
21-Oct-2018	Changes following user comments: Section 2: clarify plate solve is optional and highlight manual aperture placement (latter option is covered in AIJ Users Guide) Section 4.1, optional setting N-E orientation Section 5.2: Add Fig 5.4 and text to illustrate Yes / No / Check selections in 'Include in Analysis' column. Add option to save AAVSO report file and move Landolt reference star photometry to Appendix F