

Value of Micrometers (T.B.P.P x W.H.S) Screen = 25.75^{''}

Scale of Camera 6" focus, 26.75 fms. → F/4.45

1° = 467 = 1.186 cm.

1^{''} = 2.141

1 cm = 0.842

5 cm = 26.25

1 mm = 5.25 0.1 = 3/5^{''}

End of 1966 — 1967

R.L. Waterfield's Observatory
ASCOT Beck.

1966 Saturn Passage of ring through center.

Very few observations were made
That showed Saturn pass to ^{the} passage through the ring plane (Oct 29) on Oct 16.8
When the ring was of course very easy as fine continuous middle
S. Mc. Neil got an observation of Oct 19.8 (10 days before passage) and
despite v. poor seeing Ring was easy in 6"

Bad weather prevented further observation until Oct. 29.8 — over 12 hours
after the passage & no sign of the ring could be seen.

Bth J & S.M.N. showed the planet several times between Nov 2 and Nov 26
but seeing was rather very poor & no sign of the ring was ever seen.
The center reached its maximum elevation above the dark side of the ring on
Nov 23 — only $0^{\circ}26'$. at the same time the Sun was about $2^{\circ}5'$ above the
light side of the ring. Undoubtedly the failure to see the dark side of the ring was due
to this very small angle of $0^{\circ}26'$ of the center's inclination.

The center again passed through the plane of the ring to its light side on Dec 18.0
& we hoped to be able to see the ring by the evening of 18th Dec — 3/4 days after.
Unfortunately the nights of 18th & 19th Dec were both overcast, and we
could not observe the planet until Dec 20 17^h 30^m. Despite v. poor seeing
the ring was then very easy as a fine continuous middle.

M.B. I had showed previous passages through the plane of the ring:

1921 at RO Greenwick; Phillips, Healdy; Skafed; W.H.S. W.F. Howard &

1936. Healdy and at Praha (returning from Christ eclipse)

1951 ASCOT.

1967 Occultation (1a)

Jan 21. Z.C. 0621 May 6.2 D.D. Moon 11.1 P.A. 128° Pred. U.T. 19^h 37^m 12.3 (-)

NA +0.64 Observed L.S.T. 3^h 37^m 0.75 (RW) +0.2 Clock error +0.56 Slow

Cor " " 3^h 37^m 1.3

∴ U.T. of Düsseldorf was 19^h 37^m 40.3 [O-C = -13.7 (No a+b)]

good quality.

1967 Occultation (2a, 1b)

Feb 21 Z.C. 1169 May 5.4 S.D. Moon 12.4 P.A. 96° Pred. U.T. 18^h 42^m 35^s

Observed L.S.T. 18^h 43^m 57.5 (RW) +0.1 Clock error +2.87 Slow

" " 18^h 43^m 59.5 (SM-N) +0.1

Corrected LST 18^h 44^m 0.4 (RW)

" " 18^h 44^m 0.4 (SM-N)

NA +0.51 ∴ U.T. Düsseldorf 18^h 42^m 35.2 (RW) O-C +0.2

NA +0.51 18^h 42^m 35.2 (SM-N) O-C +0.2

good quality

1967 Occultation (3a, 2b)

Mar 19 Z.C. 0994 May 6.5 S.D. Moon 8.8 P.A. 121° Pred. U.T. 23^h 52^m 25^s

Observed LST 11^h 37^m 5.25 +0.2 (RW) 11^h 37^m 6.25 for SM-N Clock +5.39 Slow

Corrected LST 11^h 37^m 10.64 (RW) 11^h 37^m 11.64 (SM-N)

NA -0.59 ∴ U.T. Düsseldorf was 23^h 52^m 24.2 (RW) (O-C -0.8)

NA -1.12 23^h 52^m 26.2 (SM-N) (O-C +0.2)

good quality

1967

March 5-8 (nights of Sun-Mon 5-6th) OaO White 6" Triplet.

T = 49°-47° F3-11 Transp 7/10

1967c. Comet head was comet discovered in Bern Feb. 11. (but factored local light)

(1967c) Poor weather prevented observation before to-night. The comet had been in t. high declination (81° at discovery) and is moving very rapidly Southwards.

Began an exposure during on a moving web in PA 90° γ at a rate of 40.25 in 10 minutes. S.M. Will guided the telescope in 1 1/4 minute steps (5" each) and exposed for 21 minutes 15 sec.

Exp. 6^h 35^m 1^s → 6^h 56^m 16^s + w2 Clock error = -0.4^s Fast

Comet's Mid L.S.T. = 6^h 45^m 38.1^s Plate centered at 5^h 8.5^m + 29° 6'.0

A quick visual search with 6" showed no sign of comet.

The plate showed a strong image of the comet. The coma moderately condensed is about 45" in diameter and there is a broad straight tail 3' long directed towards PA 70° N γ - more or less away from Sun.

Integrated photometric magnitude estimated at 10.0. Because of scarcity of guide stars comet is about 1° off plate center.

Plate monument C Wtd 1967c

Twelve stars were measured & 4 independent reductions made - their mean giving the adopted position :-

Mar 5.831043 $\alpha = 5^h 7^m 22.67$ $\delta = +28^\circ 55' 35.2$ (1950.0)

(UT = 19^h 56^m 42.1 L.S.T. = 6^h 45^m 38.1)

Range $\alpha = 0.11$ $\delta = 2.3$

1967. Occultation (4a)

Mar 20 Z.C. 1131 Mag 7.2 S.D. Moon 9.8 P.A. 110° Pnd. U.T. = 23^h 30^m 49^s

Observed by R.H.W. 6" (S.M. & Hill with 4" hot star before & in the 'wind' vibration)

Observed L.S.T. 11^h 19^m 22.5 + cor Clock error + 5.95 Slow.

Corrected " " 11^h 19^m 28.5

NA < 1.77 ∴ U.T. of Bieleman was 23^h 30^m 49.0 (O-C = 0.0)

No correction - probably very small

good quality

1967 Occultation 5 (36)

Mar 22 Z.C. 1393 Mag 6.7 S.D. Moon 11.7 P.A. 163° Pnd. U.T. 22^h 20^m 46^s

Observed by S.M. & Hill 6" (R.H.W. in London)

Observed L.S.T. 10^h 16^m 57.5 + cor Clock error + 6.4 Slow

Corrected " " 10^h 17^m 3.9

∴ U.T. Bieleman was 22^h 20^m 42.8 quality fair - doubt of clock error

NA - 0.83

[O-C = -2.2]

gafed timer sound.

1967 Occultation 6 (5a, 4b)

Mar 23 Z.C. 1514 Mag 6.1 S.D. Moon 12.7 P.A. 70° Pnd. U.T. = 21^h 24^m 55^s

NA + 0.68 Observed L.S.T. 9^h 24^m 57.8 + cor (R.H.W.) Clock error + 6.16 Slow

NA + 0.38 9 24 58.75 + cor (S.M. & Hill)

Cor. " " 9^h 25^m 3.96 (R.H.W.) and 9^h 25^m 4.91 (S.M. & Hill)

∴ U.T. Difference was 21^h 24^m 55.5 (R.H.W.) good quality.

21 24 56.4 (S.M. & Hill)

[O-C + 5.5 R.H.W.]
[O-C + 1.4 S.M. & Hill]

1967

March 27.8 (night of Mon-Tues 27⁵.28⁵)

Oa-O plate 6" Triplet

Count Wild 1967c

Transp. 5/10 δ Wild IV

Exposure 20 minutes L.S.T $8^h 7^m 1^s \rightarrow 8^h 27^m 18^s + 10$

Mid L.S.T. of Exposure = $8^h 17^m 4.5^s$

Clock +3.0 Slow

Guided by S.M. Neil who drove on a moving web in PA 12° at rate of $31.2''$ in 20 minutes with $\frac{1}{4}$ or 5 minute steps of $7.8''$.

Plate centered at $\alpha = 5^h 15^m 0$ $\delta = +6^\circ 29'$ approx.

The plate shows that the count has very greatly faded since Mar 5.7
Photographic mag. est. at 12.0. The coma is about $25''$ in diameter & essentially circular with no sign of any tail.

Plate measurement C. Wild 1967c

March 27.8

Main stars were measured & 3 independent reductions made; the mean of these three gave the adopted position:-

March 27.834295 $\alpha = 5^h 14^m 1.14$ $\delta = +6^\circ 24' 12.2$ (1950.0)

UT: $20^h 1^m 23.1$

LST = $8^h 17^m 4.0$

Range $\alpha = .07$ $\delta = 2.2$

Mar 31.8 C. Wild (1967c) (night of Fri-Sat 31-1st) Oa-O plate 6" Triplet Transp. 4.

The transparency was only ^{good} estimated - at 50 good as Mar 27.8. We gave an exposure of 25 min. Corrected time of mid exposure was L.S.T $8^h 57^m 32.3^s$. Drove on moving web in 5 minute steps of $7.4''$ each. Plate center $5^h 17^m + 4^\circ 30'$ (S.M.N. drove)

Careful search showed no signs of comet which is estimated less than 13.0 mag

1967 Occultation 7 (5h)

April 17 Z.C. 1088 May 5.6 DD Moon 7.0 PA 34° Pub. U.T. 0^h 12^m 47^s

Observed by S. McMill 6" (only) (Time signals RSW)

Observed LST 13^h 48^m 16.5^s + cor (SMN) Clock error -10.3 Feet.

Corrected time 13^h 48^m 6.2^s LST.

PA -0.01

∴ U.T. Disappearance was 0^h 12^m 52.9^s (O-C + 5.9)
 good quality

1967 Occultation 8 (6h)

April 17 Z.C. 1089 May 6.8 DD Moon 7.0 P.A. 151° Pub. U.T. 0^h 18^m 39^s

Observed by S. McMill (Time signals RSW)

Observed L.S.T. 13^h 54^m 4.8^s + cor (SMN) Clock error -10.3 Feet

Corrected time 13^h 53^m 54.5^s LST

PA -0.71

∴ U.T. Disappearance was 0^h 18^m 40.2^s (O-C + 1.2)

1967.

June 6:0 Periodic Comet Tempel (2)

This was the first attempt I made to photograph the comet

June 6:0 (night of Mon.-Tuesday, June 5-6) Oa-0 plate

Exposure 30 min. Transh 4-5/10 Temp = 61°-61° Foun 7.

LST 16^h 45^m 2.0 → 17^h 15^m 2.0 + wt (clock 2.02 fast)

Corrected LST mid Exposure = 17^h 0^m 0.0

The telescope was driven with a fixed web. The comet was moving slowly: 10-minute motion $\Delta = +2.5$ $\delta = -1.4$ (i.e. $\pm 3''$ in PA \nearrow)

There were no signs of the comet visually in 6"; on the plate it was a faint object, but when found was well centered & suitable for measurement. The estimated photographic magnitude was 12.5

Measurement 9 stars were measured on plate & 3 individual reductions made

The straight mean was taken: June 6.005320 18^h 25^m 28.31 $\delta = -2^{\circ} 55' 7.8$ (1950.0)

June 13:0 (night of June 12-13, Monday-Tuesday) P. Tempel (2) Oa-0 plate

Exposure 25 min. Transh 5/10 T = 51°-49° Foun 10.

LST 16^h 50^m 1.0 → 17^h 15^m 1.0 + wt (Clock 3.2 slow)

Cor. LST mid Exp. 17^h 2^m 34.2 i.o. V.T 23^h 42^m 42.1 June 12.

Star on fixed web. Comet slow moving.

Good image of comet on plate. Photographic magnitude estimated 12.0

Measurement 12 stars were measured & 4 individual reductions made. The straight mean was taken: June 12.987987 18^h 27^m 23.01 $\delta = -3^{\circ} 39' 19$ (1950.0)

1967.

1967 Occult. 9. (6a)

June 16. ZC. 1855 Mag 7.1 Moon 8.7 days D.D. P.A. 101°

Observed by RLW. Observed LST $15^h 27^m 49.5$ +003 Clock error 0.80 min.

Corrected Observed LST = $15^h 27^m 48.70$

∴ U.T. Difference was $21^h 52^m 28.5$ (good quality)

O-C = -0.9

1967 Occult. 10 (7a)

June 18 ZC. 2104 Mag. 7.5 ~~S.D.~~ Moon 10.7d. P.A. 92°

Observed by RLW. Observed LST $17^h 6^m 7.9$ +007 (Clock 1.83 min)

∴ Corrected LST = $17^h 6^m 6.07$

∴ U.T. Difference was $23^h 22^m 37.9$ (good quality)

O-C = -4.7

1967 Occult. 11 (6a) (7b) Michael Hendin staying. Observed by him in 6" as ladder

Aug. 26 Z.C. 0457.

too high for me to observe comfortably

Mag 6.5 Ref. D.L. Moon 20.9th. P.A. 295°

I had previously calculated the distance of the dark limb, when star was to disappear, radially from the terminator - this was 676" or 26.2 Arcs. of micrometric screw. The distance was subdivided by this amount. One of two webs^(a) was placed tangentially to the bright limb at P.A. 115°. Then by moving the telescope parallel to the position webs, until web (a) was on the terminator the point of appearance was indicated by the intersection of web (b) & position web.

The star appeared within a few seconds of sec of this point & M.J.H. got a good observation.

Observed M.J.H. LST $8^h 14^m 46$ (S.I.) before U.T. $23^h 35^m 00$ [2 kept to avoid clock shift] [was found to have stopped.]

i.e. Reference observed at $23^h 26^m 46.9$ UT good quality. O-C = +3.4

1967

June 30.0 P. Tempel (2) (night of June 29-30, Thursday-Friday) Pa-0 Plate

Exposure 20 mins. Transh. 6/10 Temp. 57°-55° Foun 8

G.S.T. $17^h 46^m 8^s \rightarrow 18^h 6^m 8^s + wr.$ (Clock 6:30 fast)

Mid Exposure corrected LST $17^h 56^m 2.0^s$ i.e. U.T. = $23^h 29^m 10.6^s$

Telescope driven with fixed web

The comet was ^{fairly} easy object in 6" and its visual magnitude estimated at 11.5. There was a strong image on plate & magnitude estimated at 11.5 photoelectric.

Measurement of Plate. Nine stars were measured, but it was found that one of them $130 - 8^{\circ} 46' 38''$ gave a poor position. This was discarded; and the result taken from 3 reductions comparing 8 stars - each ^{of} 8 stars being given equal weight.

June 29. 978595 $18^h 29^m 20.31^s - 7^{\circ} 43' 55.1''$ (1950.0)

July 9.0 P. Tempel (2) (night of July 8-9, Sat.-Sun.) Pa-0 Plate

Exposure 20 mins. Transh. 6-7/10 Temp. 58°-56° F=8.

LST $18^h 10^m 2.0^s - 18^h 30^m 2.0^s + wr.$ (Clock 3:30 slow)

or LST mid-Exposure LST $18^h 20^m 5^s$ U.T. $23^h 17^m 46.5^s$.

Telescope driven with fixed web. Comet fairly easy in 6" Vis Mag 10.5-11.0 Good image on plate - no tail etc. - Phot. mag est. 10.8

Measure of Plate. Nine stars measured only 8 stars used in 3 reductions - each of 8 stars given equal weight.

July 8. 970677 $18^h 30^m 12.34^s - 11^{\circ} 22' 15.5''$ (1950.0)

1967 occult. 12 (8a)

Nov. 9. Z.C. 3227 Mag 6.4 Moon 7.6 days. $\delta. \delta.$ PA. 105° .

Observed by R.H.W. L.S.T. Dist. = $23^h 36^m 1.5^s + cor$ (Clock 2.255hr)

Corrected LST Dist. = $23^h 36^m 3.75^s$

\therefore V.T. Dist. Difference was $20^h 25^m 20.7^s$ good quality,
O.C. = -4.8

Note on Closing & Dismantling of the Observatory at Ascot.

The above occultation (Nov. 9) was essentially the last observation made at Silverd Marsh, Ascot. There were at the time no comets for observation; and what with my impending move from Ascot to somewhere in Dorset-Somerset region there was little opportunity for doing any other astronomical work.

During the summer many comets however were seen in Somerset & Dorset with a view to finding one remote from artificial lights & with a reasonably good horizon.

Finally I decided on a wooden bungalow with about $\frac{1}{4}$ acre of garden on the edge of the hamlet of Woodston near North Cadbury in Somerset. This was remote from lights & it being surrounded by agricultural land seemed likely to remain so for a long time. And the horizon was reasonably good. The site was purchased & alterations were started on the bungalow in October.

The observatory at Ascot began to be dismantled & packed on Dec. 5. R.H. Fry, T.H. Siddons & Ian Percall gave great help. The Newton box was removed & everything packed by Mr. Percall & his assistants via the I.C.S.T. agricultural research station bus & moved to Woodston in 2 journeys on Dec. 19 & Dec. 20.

1967.

July 17.0 P. Tempel (12) [night July 16-17, Sunday-Monday] 0a-0 plate

10 min Exposure Transp. 6-7/10 Temp 63° F.

LST 19^h 11^m 5^s.0 → 19^h 21^m 5^s.0 (+40) (clock 6^h 32^m show)

Mid Exposure Corrected LST 19^h 16^m 11^s.2 = UT 23^h 42^m 16^s.2

Telescope driven fixed web. Comet was fairly easy to see on plate but seemed to be of same magnitude as on July 9. Phot. mag = 10.8 Not seen in 6" - low altitude → difficult horizon for drawing ladder.

Measurement. 10 stars were measured - but only 9 used in reductions

We gave each of nine stars equal weight -

July 16.987687 18^h 32^m 13^s.15 -15° 21' 48".7 (1950.0)

July 27.0 P. Tempel (12) (night July 26-27, Wednesday-Thursday) 0a-0 plate

Exposure 10 mins Transp 4/10 sv. low alt. Temp 65-63° F. Form 6

LST 18^h 16^m 5^s.0 → 18^h 26^m 5^s.0 +40 (clock 2^h 9^m fast.)

Mid Exposure Corrected LST 18^h 21^m 2^s.1 = U.T. 22^h 7^m 57^s.0

Telescope driven on fixed web. Extremely low altitude - telescope & camera only just above clear of walls. Unable to look for comet visually.

In spite of v. low altitude & poor transparency image of comet was, though faint, sufficiently contrasted for good measurement. Est. Phot. mag 10.5

Measurement Eleven stars were measured & into 3 reductions made

each star received an equal total weight

July 26.922188 18^h 37^m 52^s.83 -20° 52' 42".8 (1950.0)

1968 Jan-Feb. Mon. Interim house after move. Asst. to G. S. G. G.

The Observatory, Woolston, near North Cadbury, Somerset.

The household furniture, including the library and various smaller instruments, were moved in furniture vans from anot on December 22, 1967. The Observatory & the larger equipment was moved down separately on Dec ~~19~~¹⁴ & ~~23~~¹⁹ in the lorry of the Agricultural Research Station of the Imperial College of Science & Technology. The College very generously supplied the lorry & driver & provided the labour for packing all the equipment onto the lorry. On arrival at Woolston all the instruments were ^{stored} ~~parked~~ in the Bumpelow garage, with the exception of the Observatory roof & window masts which were parked & covered on the front lawn.

Ingulf & Jan Percall moved into the Bumpelows on Dec. 30 and during January & February 1968 the builders (David Chant) completed the slight alterations made to the bungalow. Early in February David Chant started laying the foundations of the Observatory & making the concrete paths to it from the bungalow.

The base of the Observatory was a concrete raft about 8" deep & about 20 feet x 20 feet with a central pier about 2 feet deep & another on one side - the first for the telescope, the second for the seasonal clock. The walls of the building were made of concrete blocks (as opposed to the original wooden ones); & the internal measurements of the building were increased by about 18" to x feet. The walls were lined on the inside with insulating boards & an air space; & a wooden floor was placed over the concrete base. The building was surrounded by

1968 Feb. Mar.

Michigan period ASOOT to Worcester

a concrete highway about 3 feet wide, lining the perimeter of the concrete raft, with a low wall of concrete about 2 feet high.

As the dimensions of the building had been increased by about 18" each way the lower part of the roof, supporting the circular steel rail to the upper roofing part, had to be enlarged to fit onto the larger building beneath. This was all carried out successfully by David Cham! & his men. [The lower part of the roof with the circular rail had to be divided into 2 halves for the road journey; but when it was reassembled the circularity of the rail had not been maintained.] The building was provided with one window due South, and a sliding door on the West side. Electric light system was laid on from the buslines.

The building was practically completed by the evening of Friday Feb. 23rd. and that same night the lower section of the pulley was set in place. On the morning of the 24th we started erecting the rest of the telescope. Soon after midday RM Fry arrived unexpectedly & gave most valuable help in the assembly of the polar & declination axes, clamps, circles & slow motions. The reconstruction was essentially complete that same evening. During that week-end I managed to get the orientation of the polar axis sufficiently accurate; ~~and~~ ^{on} the night of February 27 & 28 I exposed my first plate on Comet Keys-Seli.

A lot remained still to be done. So far the orientation of the polar axis had only been done roughly with the circle visually. The accurate adjustment photo-graphically was

Position and altitude of declination axis of 6" equatorial
at WOLSTON, SOMERSET.

$\lambda = +2^{\circ} 30' 16''$ west $\phi = +51^{\circ} 2' 47''.5$ Height 97 metres.
(10 minutes 1.3 seconds)

[Difference from Greenwich $d\lambda = +2^{\circ} 525$; $d\phi = -0.434$]

1968 Jan-Mar Inclin Period Asort to Woodston.

undertake gradually mainly during moonlight as the coast occupied occupied all the time being the dark of the moon. The wireless magnet carrying the 150 foot bifilar reel required some repair & the lead-in had to be replaced (after 20 years at Asort). This was completed by Mar 18th which signals were obtained regularly from M.S.F. (using the Vernier method of coincidences with the Sidereal clock). We also had some trouble with the Sidereal clock; however after cleaning carefully it began to perform well. From Feb to till Mar 18 we had to rely on the Chronometer (one loan from Admiralty via Dick Woolley). This was of highest quality & during January & February I had started rating it; and the times taken from it could be relied on to less than 0.1 sec. From Mar 18 we used the sidereal clock which could be relied on to 0.01 sec.

The 6" Triplet with Camera had been moved intact and I was glad to find that the focus had undergone no change.

The delimitation of the location of the new site was made from the 6" x 15" O.S. maps of the region. Owing to my inexperience it was difficult for me to determine the position of the telescope in relation to the landmarks on the maps. We used a crude approximation at first; but with Stevenson's came to stay from April 20-29 & was able to determine the precise location on the map & also with his accurate Aneroid to find the altitude of the declination axis above the 300 foot contour at the bottom of our drive. The figures finally adopted were:—
See sketch page.

C. ~~Hayashi~~ Ikeya-Seki 1967 n

1968 Feb 27-28 (night of Tuesday-Wednesday)

This was the first clear night after the preliminary adjustment of the orientation of the polar axis, and was the first opportunity for photographing the new comet ~~Hayashi~~ Ikeya-Seki discovered at end of December.

Feb 28-1. Exposure 10 mins. Gac-o Plate Transp 6-7/10, T=33° F. 15.

L.S.T. Exh $10^h 58^m 5^s \rightarrow 11^h 8^m 5^s + 01^m$. ~~More~~ ~~with~~ ~~moving~~ ~~web~~.

Orz Mid LST. $11^h 3^m 25.7$

*Clock 20.7 slow

More with moving web. P.A. $37^\circ \rightarrow 34.5$ in 10 min in $\frac{1}{4} = 8.3$ steps.

Visual Comet was easy in 2" finder. Visual mag est. 7.0

Photographic Coma Outer diameter $25'$ inner condensation $10'$ diameter.

Tail Main part - bluish tapering spine $7.5'$ long in P.A. 280°

Wide fan $210^\circ-290^\circ$, varying up to $3'$ long. Integ. Mag ± 7.0

Plate center $\pm 17^h 15^m + 40^\circ 42'$

1968 Mar 40 (Mar 3-4 night of Sunday-Monday)

C Ikeya Seki 1967 n Exposure 30 mins. Transp 5-6/10 T=34° F 14.5

L.S.T. $12^h 12^m 30^s - 12^h 42^m 30^s + 01^m$ (clock 16.6 slow). Mid Gz. LST $12^h 27^m 46.5$

More with moving web in P.A. $90^\circ \rightarrow 20$ min $M = 73.4^\circ$ in $2\frac{1}{2}$ min = $\frac{1}{8} = 9.17$ steps.

Plate center $\alpha = 17^h 17^m \delta = +47^\circ 46'$

Photographic image. Outer coma $10'$ in diameter of light. Central condensation $\pm 3'$ diameter

Main Tail: Straight rather diffuse spine extending $22'$ from center of coma in P.A. 275° .

The two "tufts" of a diverging envelope arise from N & S parts of coma towards

P.A. $300^\circ \rightarrow 210^\circ$ & are very diffuse & wide extending about $7'$ from center of coma.

The axis of the tufts & envelope is not symmetrical with the tail being in P.A. $250^\circ \pm$

Integrated Mag ± 7.0

1968. Mar 27-28 and March 31

Checks

Adjustment of Polar Axis. When the telescope was first erected at Webster on Feb. 29 the approximate orientation of the polar axis was done by means of declination circle readings. This was sufficiently good for the history of counts in all but very high declinations.

On March 27 a Plate was exposed on the Pole and 2 short exposure made with an interval of 60 minutes showed that the error in orientation was $153.5''$ - about $16''$ in zenith & $101''$ in altitude. This was corrected during daylight on afternoon of March 31.

1968

Mar 12:0 (night of Mar 11-12, Monday-Tuesday)

C/Hege-Schi 1967n Exposure 4 min. 3 days before Full moon Transh 6/10. Moon high up.

LST $10^h 58^m 1^s \rightarrow 11^h 2^m 1^s + \text{cor}$ Clock 20.5 sec Slow

Corrected mid LST = $11^h 0^m 21.5^s$ Stars with fixed web

Visual Comet easy in 6" depth microscope Est. mag ± 7.0

Photograph Central nucleus almost stellar but elongated N-S by $\approx 15''$

Central condensed part 1' diameter. Overall diameter $2\frac{1}{2}'$.

This short exposure (accelerated by near Full-moon) was taken by Position measurement.

Plate center $17^h 17^m + 59^\circ 56'$ efflux.

N.B. Though considerably foggier it could have had even longer exposure (say $\rightarrow 6$ min) without gross fog.

This was v. surprising (after a.s.w.) with such a bright moon.

Mar 26:0 (night of Mar 25-26, Monday-Tuesday)

C/Hege-Schi 1967n Exposure 32.5 min Transh 7/10. T=45°. F=12.

LST $9^h 25^m 1^s \rightarrow 9^h 57^m 32^s + \text{cor}$ (Clock 2.5^s Slow)

Corrected mid LST = $9^h 41^m 19.0^s$ Stars with moving web in PA $73^\circ 37'$ Δ

20 min Motion = 2.68 Rows = $69'' \cdot 0$ in $\frac{1}{8} \text{hr} = 2.5 \text{ mins} = 9.6 \text{ steps}$.

Plate Center $16^h 39^m + 80^\circ 41'$ (1968.3)

Photo Good plate Core diameter (through center of coma & perpendicular to axis of tail) Extent, overall 5'-7'

" " Condensed part 2'

Tail: Main part is a diffuse central spine 27' long in PA. 225°

Two "tufts" of a conical envelope bisect either side of coma symmetrical with axis of tail.

Est mag. about 7.0.

Mar 27 Kott. P. Bar. Region

10 Orientation of axis.

(April 3-4. After the plates were exposed on the P.M. - 2 short exposures given with 60 minute exposure. The overlap of the 2 images appeared complete but owing to my poor seeing the images were not exact; and the test should be repeated at some future time. However the adjustment is accurate enough for practical purposes.

1968

April 2.8 (night of April 2-3, Tuesday Wednesday)

C. Ikeya-Schi. 67m Exposure obtained during a short gap in cloudy sky.

Exposure 3 minutes Ga-O plate. Still considerable twilight & high 5. day moon
Drive on fixed web - drive not recorded during exposure.

LST $8^h 38^m 31^s \rightarrow 8^h 41^m 31^s + 65$ Clock 0.9 fast.

Corrected Mid LST = $8^h 40^m 0.1^s$.

Plate exposed primarily for position.

Photographic Coma small diameter $\pm 90''$

Central Condensed Heat $30''$ diameter.

Two radial slits: the stronger in PA $135^\circ \pm 45''$ long.

weaker in PA $30^\circ \pm 60''$ long

Plate center $\Delta = 10^h 21^m \delta = +87^\circ 0'$.

April 4.9 (night of April 4-5, Thursday Friday)

[An attempt was made to expose on C Ikeya-Schi during short gaps in cloud.

In the hurry to take advantage of this gap the telescope was accidentally set
in the wrong side of the hole - & the cover not closed until plate was developed!]

2 minute Exposure. Free drive. Mid cor LST = $10^h 31^m 1.4^s$ Plate Center $20^h 29^m + 85^\circ 30'$

1968 (1)

Observed 1968. April 6. Z.C. 1211 May 6.2 Greenwich D.L. P.A. 111°

Observed L.S.T. of Greenwich 12^h 55^m 20^s.0 + cor (clock 1^s.6 slow)

∴ Cor. L.S.T. of G.S. was 12^h 55^m 21^s.6

or U.T. 24^h 4^m 30^s.4 sec. (April 7 0^h 4^m 30^s.4)

The quality of observation was fairly good.

This was the 1st occultation observed at Washinton. We intended to observe one immediately before by 1206.Z.C. - but Stephells made an error in the longitude correction & missed it - but just had time to put it aside for the second one, 26 1211.

Preliminary reduction from N.A. Office given - 0^s.15 for ^{of this observation} correction of distance from corrected limb. O-C for Hambrook - 4^s.6

468

April 5.8 (night of April 5-6, Friday - Saturday) Very transp 7/10 - but mostly 2nd quarter

Ikeya-Schi Exposure 3 mins Plate Oa-0 White
1967ⁿ

LST. $9^h 40^m 30^s \rightarrow 9^h 43^m 30^s + w$ (clock 10^s slow)

\therefore Corrected Mid LST = $9^h 42^m 1^s$

Star with fixed web. Plate center $\alpha = 7^h 39^m \delta = +83^\circ 38'$

Photographic: Unfortunately the stars in web imperfect: definite softness of star images.

Coma Overall diameter = 2'.5. Central condensation 20" diameter.

Two radial spikes: 60" long in P.A. 135° & 110° : neither are quite, tho' very nearly, radial.

April 6.8 (night of April 6-7, Saturday, Sunday)

C Ikeya-Schi Exposure 2 mins. Plate Oa-0. Very Transp 7/10. But half moon by
1967ⁿ

LST $9^h 48^m 1^s - 9^h 50^m 1^s + w$. (Clock 1.6 sec slow)

\therefore Corrected Mid LST = $9^h 49^m 2^s.6$

Star with fixed web. Plate center $\alpha = 7^h 42^m \delta = +83^\circ 38'$

Rather short exposure & weaker image. Coma about 2.0 diameter. Central cond about 20". ^{Circular} No detail. L

~~April 10.8~~

April 11.8 (night of Apr 11-12, Thursday, Friday)

C Ikeya-Schi Exposure 2 mins. Plate Oa-0 Transp 5/10 - 1 day before Full moon.
1967ⁿ

LST $11^h 30^m 1^s - 11^h 32^m 11^s + w$. (Clock 2.9 sec slow)

Corrected Mid LST $11^h 31^m 8^s.9$

Star fixed web. Plate center $\alpha = 7^h 3^m \delta = +78^\circ 56'$

Rather weak image of coma. Essentially same April 6.8 - circular.

1968

April 13.9. (night of 13-14th, Saturday Sunday)

C/ Ikeya-Schi. Exposure 2 mins Plate 04-0 Transp 4/10. T=42°
1967 n. LST 10^h 56^m 31^s → 10^h 58^m 31^s + cor Clock 3.6 sec slow.

Corrected mid LST 10^h 57^m 34.6^s

Drone on fixed web.

Plate center 6^h 54^m + 77° 5'

Center Plate, film image rather woolly. Comet image weak - near bright star.

April 14.9. (night of 14th-15th, Sunday Monday) P

C/ Ikeya-Schi Exposure 17 1/2 mins. 0a-0 Plate. Transp. Fair but more full rising.

1967 n. LST 10^h 57^m 30^s - 11^h 15^m 0^s + cor (Clock 3.9 sec slow)

Corrected mid LST 11^h 6^m 18.9^s

Drone with moving web in 2 1/2 minutes, 5.5 steps in PA 78.5

The focus is poor & the film image v. woolly. First Plate of New Batch with gross fog band rising diagonally across plate (all plates in box affected - returned to Kodak!) Fortunately comet is just clear of fog band.

Comet image strong. Central condensation 1' diameter Outer coma 3.5 diameter

With diffuser Van Lail PA 30° - 110° with steeper condensation in PA 90° and 45°

- the strongest in 45° PA is about 8' long Plate centered near 6^h 53^m + 76° 15'

April 26.0 (night of 25th-26th, Tuesday-Friday)

C/ Ikeya-Schi Exposure 45^m 5^s Plate 0a-0 Transp 7/10 T 65-50° F = 8.5

1967 n. LST 11^h 58^m 1^s - 12^h 43^m 6^s + cor (Clock 1.1^s Fast) Cor Mid LST = 12^h 20^m 39.4^s

Drone moving web in 1/4 = 2 1/2 minutes = 4.7 steps in PA 90°

Plate centered near 6^h 45^m + 68° 0'

Continued over →

Ocult (2) 1968

May 2 Z.C. 1035. May 68 Stephenson Dark limit P.A. 71° .

Observed LST of δ_{ii} $12^{\text{h}} 14^{\text{m}} 3^{\text{s}} 0 + \text{cor}$ Clock $3^{\text{s}} 0 \text{ slow}$

Corrected LST $\underline{12^{\text{h}} 14^{\text{m}} 6^{\text{s}} 0}$

U.T.SJ Observed δ_{ii} is $\underline{21^{\text{h}} 40^{\text{m}} 40^{\text{s}} 9}$

Quality of observation very good.

Velocity reduction from P.A. Office since distance of star from the direction is

+0.35 from corrected limit.

O.C. from Star chart is +1.9.

1968

April 26:0 continued C / Ikeya Selin 45 min. exposure.

Path, good plate: from good, trials were, disappointing good.
Tail in a faint diffuse broad fan, as before, PA $110^\circ - 45^\circ$. Main
extension is around PA 60° to a distance of about 12'. Outer coma
has a diameter of $\pm 4'$ and the central condensation about 1' diameter - this
is exactly placed, the coma extending more in the direction of the tail.

May 4:1 (night of May 3-4, Friday-Saturday)

Transit 3/4/10

C. Tajo-Honda-Kamamoto 1968a was discovered in Japan on April 30.8

I made a 5 minute exposure through high thin passing cloud - the
altitude I had set - with plate centered $\alpha = 0^h 13^m \delta = +35^\circ 15'$
But there was no sign of the reported 7.0 mag. comet.

Exp. 5 min LST $16^h 50^m 15^s \rightarrow 16^h 55^m 8^s$ (approx U.T. $2^h 15^m$)

May 6:0 (night of May 5-6, Sunday-Monday)

1) C/T-H-Y 1968a. Much passing cloud

Another attempt - an exposure of $1\frac{1}{2}$ minutes - was made during a
gap in the clouds. LST $15^h 20^m 10^s - 15^h 21^m 38^s$ centered on
 $\alpha = 0^h 57^m 6^s \delta = 44^\circ 20'$ again no sign of comet on plate.

It was later found that there was an error in the ^{Daily Motion given in} ~~Plate~~ Card
positioning moved from R.A. - so the comet was not shown on the above
2 plates.

2. C / Ikeya-Selin A 4 min exposure was given LST $13^h 31^m 30^s - 13^h 35^m 30^s + w 2^s 5^s$

and centered LST = $13^h 31^m 32^s$. Plate centered around $6^h 45^m + 62^\circ 15'$

Star fixed with good star map for mount. Central had 1' diam. Overall given 3' diameter. Round

May about 8:0 No tail.

1968.

May 11-12 (night of Saturday-Sunday)

C/Toyo-Honda-Yamaoka 1968a

0a-0 Plate Transp. good Full moon

Two short exposures were made to make another attempt to locate the comet. An exposure of 4 minutes was made about 2 hours before midnight, & one of 7 minutes (the extreme limit for full moon) about 2 hours after midnight.

(1) Exposure 4 min LST 12^h 56^m 31^s - 13^h 0^m 31^s + 60 (clock 1.6 slow)

Comet mid LST 12^h 58^m 32.6 (approx V.T. 21^h 49^m)

(2) Exposure 7 min LST 16^h 55^m 31^s - 17^h 2^m 31^s + 60 (1.6 slow)

Comet mid LST 16^h 59^m 2.6 V.T. = 18^h 49^m 26.1

(1) The first plate was moderately fogged; and at first sight no sign of comet could be found, but as soon as the much stronger image was found in the longer, second exposure, the image in the first exposure was soon detected.

(2) The second plate was very heavily fogged; but the image of the comet is quite obvious:

It is a large circular haze fading out to the hazy background showing extraordinarily slight condensation towards the center. The overall diameter on the longer exposure is about 3' - it is of course considerably less on plate one. The overall magnitude was estimated as at least 8 - difficult because of full moon & fogging. Only the second plate was measured for position - long line taken care for estimating the center of the diffuse coma.

Both plates were centered around 2^h 48^m 0^s and +60°.

C/Kege Sbi 1967n. 0a-0 Plate Transp. good Full moon.

Exposure 6 min. had LST covered 14^h 27^m 31.6 (approx UT 23^h 17^m)

~~Comet small diameter 5'5" centered at 14^h 27^m 31.6~~ The plate is chiefly fogged & no definite sign of the comet is seen in its position.

Plate center 6^h 47.5^m +59° 30'

1968

May 17-18 (nights of Friday, Saturday)

C/Tago-Kanda-Yamamoto Plate 0a-0 Transparency 6/10

This is the first opportunity of observing this comet without moonlight.

It was an easy diffuse object visually in 6" Mag est. at 7.5 in the 8 min Exposure from LST (mid, corrected) $14^{\text{h}} 45^{\text{m}} 46.2^{\text{s}}$

Done with fixed web. Plate centered around $3^{\text{h}} 50^{\text{m}} + 61^{\circ} 30'$

The image is clear. Coma is a very diffuse circular patch, with very little central condensation, except for peripheral fading, ^{or} remarkably uniform brightness all over. Overall diameter 1.5

May 17-18 continued.

C/Kiyasu-Sehi Plate 0a-0 Transparency 6/10

Exposure 6 min Corrected mid LST $15^{\text{h}} 57^{\text{m}} 31.2^{\text{s}}$

Done with fixed web. Plate centered around $6^{\text{h}} 52.5^{\text{m}} + 57^{\circ} 0'$

Image of comet very weak & diffuse. Overall diameter of coma 3'. There is an almost stellar central condensation.

May 23-24 (nights of Thursday-Friday). Thin cloud passing, low altitude Transp? 4-5/10

C/Tago-Kanda-Yamamoto Exposure 12 min. Mid LST $15^{\text{h}} 1^{\text{m}} 33.5^{\text{s}}$

Done fixed web - rather poor due owing to clouds. Plate centered around $4^{\text{h}} 25^{\text{m}} + 60^{\circ} 20'$

Coma is 3' in diameter. There is now greater concentration towards center & a fairly small central nucleus. Magnitude estimated 8.0

Oculation (3) 1968-

May 30. 2 C. 1131. May 72 Δ h. Δ h. Δ h. P.A. $\pm 100^\circ$

Observed disappearance LST $14^h 39^m 19.5^s$ (cal 2.2 feet)

Corrected RST $14^h 39^m 17.3^s$

Observed U.T. $22^h 15^m 23.0^s$

owing to the third haze at the very low altitude the star was, & became increasingly more so, difficult to see. There was therefore considerable uncertainty as to the exact time of disappearance and occultation.

Preliminary residuals from N.A. office give the distance of star from occulted limb at observed time as 5.31 - enormously outside our ordinary observing errors & showing that the star disappeared \approx $2-3$ sec before occultation.

7 June 11 N.P.S. Check on orientation of RST axis

2 shot exposures ($3/4$ min) were given ~~at~~ by passing Δ end of a clock-dial of 1 hour 52 minutes: an arc of 28° . The separation of the 2 images was $18''$ indicating an error of $37''$ in axis, requiring correction of $36''$ in alt. $\Delta 8''$ in azimuth. This was carried out next day on Pole Star.

1968

May 30-31 (night of Thursday-Friday) Plate Oa-0 Transp $\frac{6}{10}$: but daylight twilight & low altitude.
C' Tapo-Aruda: Yammoratu. Exposure 20 minutes Midnight LT $17^{\text{h}} 1^{\text{m}} 28^{\text{s}}.8$ (May 31.0)
82007 in P.A. 38° \rightarrow in 5 minutes ($5^{\text{h}} 9^{\text{m}}$) steps. Altitude around $4^{\text{h}} 50^{\text{m}}.5 + 5^{\text{h}} 8^{\text{m}}.30$
Fairly strong images of round crater some 3' diameter now faintly marked central $\boxed{T=52^{\circ} F=9}$
condensation about 0.5 diameter. Integrated mag. 8.5

June 19-20 (night of Wednesday-Thursday) Plate Oa-0 Transp $4-5/10$ Much Twilight & v. low Alt.
of haze & obs. a last attempt was made to get this sunset before it got
too near the sun. The only possible times were just before & just after the
lower culmination about midnight; and being midsummer there was very
considerable twilight at this time near the NW & NE horizon. Two
exposures were made one just before & the other just after midnight.

Exposure (1) 13 minutes at mid LT $17^{\text{h}} 3^{\text{m}} 32^{\text{s}}.8$ About half the camera
lens was obstructed by wall of observatory. Telescope clear. Scope fixed west.

Exposure (2) 17 minutes at mid LT $18^{\text{h}} 11^{\text{m}} 34^{\text{s}}.3$. On this side of the
hill the full aperture of the camera lens was unobscured by the wall, but the
telescope lens was completely obstructed. It was therefore impossible to guide &
I had to rely on a fine drive - slight doubling of images resulting.

Both plates were centered around $7^{\text{h}} 11^{\text{m}}.5 + 48^{\circ} 15^{\circ}.0$

On both plates one can make out very faintly a large very
diffuse blur which is undoubtedly the comet.

1968.

© Whitaker Thomas 1968 b.

June 20-21 (night of Thursday-Friday). New record of discovery independently by Whitaker & Thomas of comet on June 15 & 17 respectively, and giving an estimated daily motion of $+0.06 \rightarrow +2^{\circ}44'$ & magnitude 9. It was thought best to make a fairly long exposure with fixed web on the probable position & later to make a 22 exposure driving on comet.

Exposure (1) © Whitaker Thomas Oa-0 plate Transp. 5/10 Partly Cloudy.

Exposure 22 min 29 sec. Mid cor. LST = $17^h 12^m 18.0$ Screen fixed web

Exposure (2) © U-T. 10 min 2 sec. Mid cor. LST. $18^h 55^m 4.4$

Screen in PA $85^{\circ}20'$ in $1\frac{1}{4}$ min - 8" - steps. Transp. 6-7/10
Both plates centered around $15^h 18^m + 18^{\circ}$ in cloud.

Exposure (1) shows a very strong heavily trailed image of the comet

Exposure (2) shows an outer diffuse coma with overall diameter of about 4' & a moderately heavily condensed central condensation of about 0.5 diameter which is eccentrically placed towards 270° in outer coma (i.e. towards trailing).
Estimated integrated magnitude 8.0.

© Whitaker Thomas 1968 b.

June 23-24 (night of Sunday-Monday) Oa-0 plates.

Transp. 6/10.

Exposure (1) 15 min. Mid cor. LST. $16^h 47^m 58.3$. Screen in P.A. $85^{\circ}40'$ in $2\frac{1}{2}$ minute 9.6 steps

Exposure (2) 10 min. Mid cor. LST $17^h 53^m 58.3$ Screen with fixed web.

Both plates centered around $15^h 22^m + 22^{\circ}$

The image on Plate (1) is very strong. Outer diffuse coma 4' diameter & faint heavily condensed central part 1' diameter. In Plate 2 the images are all surprisingly faint (? not fully opened shutter). In both plates the focus is poor.
The mag. est. visually in 6" & spectroscopically as 9.0


1968

June 25-26 (night of Tuesday - Wednesday) Visual duration June 25.9
C. Whitaker Thomas. The connection showed visually only in 6". It was
very diffuse & difficult with visible condensation. Est visual mag 9.5

June 29-30 (night of Sat. - Sun) Oa-O Plate Transp 8/10 Twilight ++ T=56-56°F 8
© Whitaker - Thomas 1968b.

Explosion (1) R.L.W. 22.5 min Cor. LST mid exp = $18^h 16^m 16.9^s$
Strong image of comet. Outer diffuse cone 3' diameter has moderately concentrated
inner about 0.5. Then appears to be a small circulus connects to C.C.
- not trailed like stars - in PA $20^\circ \Delta 30''$

Explosion (2) (Beth & Keovil) 22.5 min Cor. LST mid Exp $19^h 41^m 15.7^s$
Thin plate, which in air well defined, shows no sign of the ? comes to coma.
Otherwise same as plate (1).

Both plates were driven with moving web towards PA 82° 
in $2\frac{1}{2}$ minutes (5.9) steps. They were centered at $\alpha = 15^h 25^m 0^s$ $\delta = +28^\circ 59'$
Est mag visual & photo = 9.5 (1968.5)

July 1-2 (night of Monday - Tuesday) Oa-O plate Transp 3-4/10 T=70°F 4.
© Whitaker Thomas 1968b. also passing cloud, which stopped exposure.


Explosion 16 min 1 sec. Cor. LST mid explosion = $18^h 23^m 1.3^s$
Done on moving web towards PA 80° 
in $2\frac{1}{2}$ min (5.0 steps)
Plate centered on $\alpha = 15^h 26^m 30^s$ $\delta = +30^\circ 38'$ (1968.5)

Plate fogged & much cloud scattering round star in exp.
Only great wealth faint & diffuse image of comet.

1968

July 4.0 (night of Wed-Thursday) Plate Oa-0 Transp 5/10 T=55°-52° F=9.

C/Whittaker-Thorne^{1968b} Exposure 22.5 min Cor LST mid Exp = 18^h 46^m 19.^s7

Drone with moving web towards PA 78° ↘ in 2½ min (4.3) steps.

Plate entered at $\lambda = 15^{\text{h}} 27.8$ $\delta = +32^{\circ} 5'$ (1968)

Outer cone very diffuse 2.5 diameter; moderately concentrated central condensation 0.50
Est. photographic mag = 9.8

July 9.0 (night of Monday-Tuesday) Plate Oa-0 Transp 3/10 in zenith. V. low altitude

C/Honda 1968c New comet discovered July 6.7 by Honda in Japan

A search plate was exposed for 7 minutes: defocus was full moon, poor
transparency & thick haze near horizon & very low altitude (just clear of wall)

Moreover an old plate (Zeyan Id) was only one available.

As might be expected plate was hopelessly fogged, very few stars & no sign of comet.

Exposure 7 min Cor LST mid exposure = 20^h 51^m 29.^s2. Plate entered at
5^h 8.8 + 40° 42' Fixed web. Observed Plate J Keovil.

July 18.0 (night of Wednesday-Thursday) Plates Oa-0s

I C/Whittaker-Thorne 1968b Transp 7/10 T=58° F=9

Exposure 40 min RW Cor LST mid Exposure 18^h 30^m 0.^s3. Plate entered 15^h 39.3 + 37° 55'

Drone with moving web towards PA 53° 22' ↘ in 10 min (7.7) steps.

Comet shows an extremely diffuse outer cone 4-5' diameter & poorly concentrated central part 0.50

Est. photo mag 11.0-11.5

II C/Honda 1968c Transp zenith 7/10. Bdr. low alt. 3rd equator Moon & Twilight +

Exposure 5 min. Cor. L.S.T. mid Exposure 20^h 9^m 30.^s3. Plate entered 5^h 10.6 + 44° 27'

observed Plate J. Keovil. Drone with fixed web. The comet is small but

1968

July 18 continued C/Honda 1968c

heavily condensed. Outer coma to faint 1.5 diam. Strong central condensation ⁽³⁰⁾ 30 diam
The comet was first picked up in 4" Wray by P. J. K.
Visual mag in 4" estimated (P. J. K. + R.W.) 8.0 Photo mag est. 8.0

July 24.0 (night of Tuesday - Wednesday) Plate Oc-0. Transp 7/10 T=55° F=9
C/Honda 1968c Two exposures were made (1) long & (2) short.

Exposure (1) 30 min G.L.S.T mid exposure 20^h 22^m 6.6 Swiss moving web

Exposure (2) 4 min G.L.S.T mid exposure 20^h 57^m 6.8 Swiss fixed web

Exposure (1) web was moved towards P.A 90° δ in 5 min steps of 3.9

Plate center at $\alpha = 5^h 10^m 10^s$ $\delta = +47^\circ 31.0$ (1968.5)

Visually the comet was bright in 6" and vis mag estimated at 7.5

Plate (1) very strong image of comet. Inner coma heavily condensed about 1.5 diam. Outer
coma traceable to diameter of 5'. There is a wide but very faint & diffuse
fan tail from PA 260° - 340° composed of streamers the most definite
being 7' long in 280° & 4' long in 260° & 340°. There is also a broad
sunward "beard" about 2' long in 130° - this is rather stronger than the others.
Mag Photo est about 7.5.

Plate (2) This short exposure shows good stream image & the comet water heavily
condensed o.s. center & overall coma to 1.5.

July 28.0 (night of Sat. - Sunday) Plate Oc-0 Transp 4/10 T=56.55° F=9

C/Whitaker-Turram 1968b. Exposure 50^m 4^s G.L.S.T mid Exp = 20^h 4^m 7.2

Plate centered at 15^h 49^m 20^s +39° 37.0 (1968.5), Swiss moving web

towards PA 28° δ in 10 min (6.0) steps. continued \rightarrow

1968

July 28.0 continued C/Whittem-Thomas 1968 6.

Cored image faint. A very faint diffuse outer coma of about 2.5 diameter with central condensation - almost stellar, but elongated PA 140° - 320° about 30" long. Mag of central condensation about 14; integrated mag. about 12.0-12.5

(night of Sunday-Monday)

July 29.0 / C/Honda 1968 c Oa-O plate Transp 4/10, incoming cloud $T=56^{\circ}$ $F=9$
Exposure 3 minutes - Fixed web. Cor. LST mid Exp. = $20^{\text{h}} 55^{\text{m}} 38.3^{\text{s}}$ Plate centered at $5^{\text{h}} 8^{\text{m}} 24^{\text{s}} + 50^{\circ} 13.5$ (1968.5). Strong image of comet & good star images. Outer coma out to about 1.0 diam. dense central condensation about 20" diam. Est. integrated mag. via photo 7.0

July 31.0 (night of Tuesday-Wednesday) Oa-O plate Transp 3-4/10 $T=57^{\circ}$ $F=8$
C/Honda 1968 c Exposure 40 mins cor LST mid Exp. $21^{\text{h}} 19^{\text{m}} 6.3^{\text{s}}$ Plate at $5^{\text{h}} 7^{\text{m}} 36^{\text{s}} + 51^{\circ} 29.5$ (1968.5). Door moving web towards PA $84^{\circ} 30'$ Δ in 5 min (8"1) steps. Strong image of comet. Central condensation 1' diam. Outer coma about 3.5 diameter. Very faint diffuse tail towards P.A. 280° about 6' long. The poor transparency was doubtless the reason why this plate shows less extension & detail than the 30 min exposure of July 24.0. Est mag photo = 7.0

Aug 14.0 (night of Tuesday-Wed) Oa-O plate Transp 6/10 but hazy cloud. $T=58^{\circ}$ $F=8$
C/Honda 1968 c Exposure 2 min. Cor LST mid Exp. $19^{\text{h}} 29^{\text{m}} 0.5^{\text{s}}$ Plate at $4^{\text{h}} 51.6 + 64^{\circ} 3.5$ (1968.5)
Door fixed web. (2 min motion in $6^{\text{h}} 03$ in PA $79^{\circ} 43'$ Δ) Strong image. Outer coma 2' diam. Heavily condensed central condensation 30" diameter. Visibly bright in 6". Vis photo mag est = 6.0

Occlusion (4) 1968

Aug 16.0 Z.C. 0486 May 5.2 Reckman ΔL (hr) P.A. 216°

The age of moon was 21.0 days. The radial distance of the star from the Terminator at appearance was found to be 14.4 (a 33.6 Revolution of Moon Sun). The position web was placed radially across the moon, on distance wire was placed so that its intersection with position wire was on Terminator; the film, at distance of 33.6 revs, was, at its intersection with the position web, on the point of the limb where star would reappear. Good quality, reappears without error.

Obs. L.S.T. was $22^h 43^m 7.5$ i or clock $6:0$ obs

Obs. on LST = $22^h 43^m 8.5 \rightarrow$ Aug 16 $1^h 15^m 9.8$ U.T. Reappearance.

Preliminary results from NA office gives distance of star from corrected limb as $+0.54$

1968

Aug 19.0 (night of Sunday - Monday) Oa-0 Plate. Transp 6/10 T = 56°-53° F8.

C/Honda 1968c Exposure 25 min. Cor. LST mid Expt = $19^h 32^m 33.1^s$

Plate centered at $4^h 36^m 5^s + 70^\circ 47' 0''$ (1968.5). Drove with moving web towards P.A. $73^\circ 0'$ \nearrow in $2\frac{1}{2}$ min (10"19) steps.

Strong mag of comet - inner condensation region, concentrated 2' diameter. Outer cone traced to about 6' diam. Diffuse fan tail with fairly sharp & narrow central spine about 13' long in P.A. 280° Plate mag est ± 5.5

M.B. Two large condensation flows S but clear of comet - but others smaller, are mass!

Aug 23.0 (night of Thursday - Friday) Oa-0 Plate Transp 6/10 T = 66°-63° F6.

C/Whitaker Thomas 1968b Exposure 90 min. Cor. LST mid Expt = $20^h 25^m 22.6^s$

Plate centered at $16^h 22^m 8^s + 40^\circ 28'$ (1968.5). Drove with moving web towards P.A. 0° \leftarrow in 5 min (4"16) steps. Good plate: focus, clear & negligible fogging. But no sign of comet. Presumably fainter than mag 14.

Aug 26.0 (night of Sunday - Monday) Plate Oa-0 Transp 6/10.

C/Baldwin - Clayton ^{1968d} New comet discovered on Aug 24 during S-W Astromin conference in New Mexico. LST mag of this on Aug 25 evening giving 2 approximate positions, and exposed a plate the same evening - about 2 hours before midnight.

Exposure 20 minutes Cor. LST mid Expt $20^h 9^m 8.7^s$ Plate centered at $18^h 58^m + 32^\circ 15'$ (1968.5)
Drove with moving web towards P.A. $27^\circ 40'$ \rightarrow in 5 min (14"1) steps. Strong mag of comet: heavily condensed central condensation 30", very faint outer cone to about 3' diameter. Integrated mag est = 10.0

1968

Aug 27.0 (night of Monday-Tuesday) Oa-O plate. Transp 6/10 $T_{65} = 64^{\circ}$ FG

C/Bally-wrhan - Clayton 1968d Exposure 7.5 minutes. Cor LST mid Exh = $20^{\text{h}} 10^{\text{m}} 52.4^{\text{s}}$

Drone with moving web towards PA $11^{\circ} 9'$ \rightarrow in $2\frac{1}{2}$ minutes (6.5) steps.

This short exposure shows a faint near stellar central condensation with very faint, diffuse outer coma about 2' in diameter. Phot mag int. exp. Est. 10.0

Sept. 1.0 (night of Saturday-Sunday) Oa-O plate Transp 7/10 $T_{57} = 55^{\circ}$ FG.

1) C/Bally-wrhan - Clayton 1968d Exposure 20 min Cor LST mid exposure $20^{\text{h}} 37^{\text{m}} 35.5^{\text{s}}$

Plate centered at $18^{\text{h}} 33^{\text{m}} 0^{\text{s}} + 32^{\circ} 55'$ (1968.5). Drone with moving web towards

P.A. = 8° \rightarrow in $2\frac{1}{2}$ min (5.4) steps. Strong image of comet. Central condensation

heavily condensed 30" diameter; faint outer coma to about 2' diameter. Appears

slightly fainter than on 25.9: Est photo mag = 10.5

2) C/Honda 1968c Oa-O plate Transp 7/10 $T = 55^{\circ} - 53^{\circ}$ FB

Exposure 17 min, Cor LST mid Exh = $22^{\text{h}} 28^{\text{m}} 35.5^{\text{s}}$. Plate centered at

$19^{\text{h}} 26^{\text{m}} 35.5^{\text{s}} + 76^{\circ} 42.5'$. Drone with moving web towards P.A. $63^{\circ} 50'$

\rightarrow in 1 minute (7.8) steps. Very strong image of comet: Central

condensation (heavily condensed) about 2' diameter, outer coma about 6' diameter.

Near linear tail 17" long towards PA 130° . The coma is hooded, curving

backwards symmetrically on either side of narrow tail. App mag = 5.0

1968

Sept. 7.0 (mid of Friday - Saturday) Oa-0 plate Transp? Full moon

C/ Honda 1968c Exposure 2^m 4^{sec} LST center of Mid Explan 19^h 34^m 8^s.0

Plate centered at 18^h 26^m 25^s + 53° 7' (1968.5) Drizz on rising web towards P.A. 80° 2' ↓ in 1 minute step of 8". Owing to full moon & drizzle only 2 minutes exposure the plate is fairly heavily fogged. The image is considerably stronger than on Aug 14 (with similar exposure) Central condensation 45" outer cone 4' diameter Integrated mag est at 5.0.

Sept. 10.0 (mid of Monday - Tuesday) Oa-0 plate Transp 4/10 Bright last Quarter moon

C/ Bally - Weber - Clayton 1968d. Exposure 15 mins. Cor LST mid Exp = 20^h 1^m 33^s.9

Plate centered at 18^h 7^m 2^s + 32° 53'.5 Drizz with rising web towards

P.A. 0° → in 5 min (6.6) steps. The plate is fairly heavily fogged by moonlight. The central image is clear: Central condensation stiller 10" diam outer cone faint to 2.5 diameters. Est integrated mag = 11.0

Oct. 15.0 (mid of Monday - Tuesday) Oa-0 plate Transp 7/10

II C/ Bally - Weber - Clayton 1968d. Exposure 35 mins. Cor LST mid Exp = 21^h 26^m 30^s.7

Plate centered at 17^h 30^m 0^s + 32° 2'.0 Drizz with fixed web (v. slow rising).

Good plate: no fogging good star images. Comet very strong - but small

Extremely heavily condensed central condensation - almost stiller - 10" diameter

Very faint & diffuse outer cone 5' in diameter. Integrated mag est. 11.8.

I C/ Honda 1968c Exposure 10 mins. Cor. LST mid exposure 20^h 20^m 12^s.2

Plate centered at 18^h 8^m 30^s S = -6° 0'. Drizz with fixed web. Image comet

strong: heavily condensed central condensation 30" diameter Faint outer cone 3' diameter

Integrated photometric magnitude estimated at 7.8 mag

1968

Oct. 29.0 (night of Monday-Tuesday) Plate Oa-0 Transp. 7/10. $T = 50^\circ$

C/Wild 1968 f - a new comet discovered in Bernese on Oct. 17. reported to be mag 15 with daily motion -2.65 in α ; $-17'$ in δ . It was decided

to give a 40 minute exposure + drive on fixed web owing to uncertainty in Daily

Motion: Exposure 40 minutes; cos. LST mid Exposure = $1^h 34^m 4.6^s$

Plate centred at $2^h 28^m 50^s + 36^\circ 3'$. ~~Swen~~ with fixed web. Owing to the

presence of at least 2 very faint nebulae in the close vicinity of comet's

position there was at first doubt as to which of 3 objects was the comet

However the matter was decided by reference to Franklin Adams charts & N&C

which showed two of the 3 objects were nebulae, and that the 3rd (midway

in brightness between the other two) must be the comet. Precise measurements

later confirmed this. The image of the comet is an extremely faint &

diffuse patch of light ~~0.5 x 1.0~~ (0.5 x 1.0) being elongated towards P.A.s

60° and 240° . (The dimensions & P.A.s of the elongated image correspond

with the reported daily motion given above). No significant condensation

could be detected in the image. The integrated mag was estimated at 14.0

Nov. 4.0 & 5.0 (nights of Sun-Mon, & Mon-Tues) Plate Oa-0 Transp. 4/10 Twilight

C/Honda 1968c. On both these evenings attempts were made to photograph

the comet before it went too far south. Evening twilight & haze was however

resulted in no image being obtained, & the plates being badly fogged

- (1) Nov. 3.75 5 min exposure Midexp cos LST $21^h 4^m 29.0^s$ Centred at $18^h 15^m - 14^\circ 4'$
- (2) Nov. 4.75 6 min exposure Midexp cos LST $20^h 58^m 0.2^s$ Centred at $18^h 18.8^m - 14^\circ 20'$

Occlusion (5) 1968

1968 Nov 8 Z.C. 0885 Mag 5.6 Ruckman Dark hill PA 274°

The calculated radial distance from reappearing star to terminator was $3' 27'' = 8.04$ Pur of minutes saw. The star reappeared close to the intersection of the webs

Observed reappearance was LST $24^h 58^m 25.0^s$ + correction (clock 1.5 Feet

\therefore corrected time was $24^h 58^m 23.5^s$ LST

This corresponds to:

Time of reappearance = $21^h 55^m 50.3^s$

Preliminary residuals from NA office inquire by time as $21 55 50.8$, and give resulting distance of star from corrected limb as $+0.66$

1968

Dec. 12:75 (Thursday evening) Plate 0a-0 Poor Transp., Haze

An attempt to get a flat plate of C/ Bally Urban Clayton 1968d was made; but there was no success owing to poor sky & comet getting low in the west. Exposure $38^m 20^s$. On LST mid Exp. $23^h 57^m 27.8^s$ Plate centered at $17^h 56^m 0 + 37^{\circ} 50'$ 1968. Down with moving web towards PA 54° Δ in 10 min. (6:7) steps. No sign of comet on plate.

Dec. 16:75 (Monday evening) Plate 0a-0 Poor Transp. Haze. Much passing cloud.

C/ Bally Urban - Clayton 1968d. This was the final attempt on this comet & also failed. Total of interrupted exposure $35^m 45^s$. Approx time of recorded LST of mid exposure $24^h 33^m 47^s$ down & moving web towards PA 55° Δ in 10 min. (6:8) steps. Plate centered at $17^h 59^m + 38^{\circ} 49'$. No sign of comet.

Plate Measurement Jan 8.75 (Thomas)

Plate I Jan 8.78362 4^h 33^m 59.62 +81° 30' 12.2 (1950.0)

6 stars used 81°156, 81°157, 81°161 $\delta_s = .346960, .345277, .307763$

BD 81°151, 80°142, 81°165 $\delta_s = .232037, .343222, .424721$

Mag. $L = 1.2$ $S = 1.3$ δ used mean from two reductions (3 other stars measured but discarded)

Plate II Jan 8.86615 4^h 33^m 51.40 +81° 29' 52.2 (1950.0)

6 stars used BD 81°156, 81°157, 81°161 $\delta_s = .339994, .368841, .291165$

BD 81°155, 80°142, 81°165 $\delta_s = .310225, .356009, .333766$

(07) Mag. $L = 0.9$ $S = 0.8$ δ used mean from 2 reductions.

Plate Measurement Jan 14.75 (Thomas)

Plate I Jan 14.78746 4^h 27^m 54.78 +81° 2' 47.0 (1950.0)

6 stars used 80°136 80°138 80°143 $\delta_s = .361173, .133510, .505317$

BD 81°144 81°157 80°144 $\delta_s = .183296, .575647, .241057$

Mag. $L = 0.7$ $S = 1.0$ δ used mean from 2 reductions (3 other stars measured but discarded)

Oculations (1) 1969 Jan 23 ZC 0132 Mag 6.9 Diffusion D.L. PA 118°

Obs LST $\delta_i = 3^h 30^m 29.0 + cor$ (clock 1/6 slow) $\therefore cor$ LST $\delta_i = 3^h 30^m 30.6$

ie. $19^h 28^m 43.1$ UT. Pulin NA residual from Cor limb -0.68

Oculations (2) 1969 Jan 28 ZC 0746 Mag 6.8 Star. D.L. PA 45°

Obs LST $\delta_i = 1^h 58^m 10.0 + cor$ (clock 1/4 slow) $\therefore cor$ LST $\delta_i = 1^h 58^m 11.4$

ie. $17^h 36^m 59.5$ UT Pulin NA residual from Cor limb -0.44 .

1969.

Jan 8.75 (mid of Wednesday - Thursday) Oa-0 platin

C/Thomas 1968 J This new comet was discovered on Dec. 19 by Thomas at Lowell Observatory. Though I got the news about 2 days later the weather prevented any observations before 6. night. Comet reported to be of mag 13 and having a daily motion of about 2" in 10 minutes. R. South was here & we gave 2 exposures.

Exposure I 50 min stopped by cloud. Transp 6-7/10 Observed R.L.W.

on 189 mid Exp. $1^h 50^m 5^s$ (U.T. $18^h 48^m 24.4^s$) ~~Plate~~ Driven on fixed web

Exposure II 20 min Transp 7/10 ^{some passing cloud} ~~(all cloud present)~~ Observed R. South

on 189 mid Exposure $3^h 50^m 7^s$ (U.T. $20^h 47^m 14.9^s$)

Both exposures: plate driven on a fixed web.

Plate I, the longer exposure shows heavily condensed central condensation 20" diameter with faint outer coma 1.5 diameter. also 2 nearly linear streaks (1) about P.A. $90^\circ 2'$ long & (2) about P.A. 150° 1.5 long. integrated mag est 12.5

Plate 2, shorter exposure shows little more than the central condensation & only a suggestion of surrounding halo - but probably more suitable for measurement

Jan 14.75 (mid of Tuesday Wednesday) Oa-0 platin Transp 7/10.

C/Thomas 1968 J
Exposure I 40 minutes on 189 mid Exp. = $2^h 20^m 9.2^s$ Transp 7/10 - no cloud.

Exposure II 38.5 minutes on 189 mid Exp. = $3^h 37^m 44.2^s$ Transp 7/10 ^{only passing cloud}

Plates centered at $4^h 29^m 7^s + 81^\circ 3'$. Driven with fixed web.

The view of comet is small, but strong: heavily condensed central condensation about 20" diameter. Key faint & diffuse outer coma 1.5 in diameter - on both plates.

On Plate I there is the doubling of the nucleus - due to a faint star, which is char of comet on Plate II. integrated mag est. at 13.0

Plate measurement Mar 7.8 (Thomson)

Mar 7.84068 $\cdot 5^h 45^m 13.58 + 73^\circ 47' 18.1$ (1950.0)

6 stars used $73^\circ 29.7, 74^\circ 25.9, 73^\circ 30.6$ $D_s = .466162, .151693, .882144$

BD $74^\circ 25.4, 73^\circ 30.1, 73^\circ 30.9$ $D_s = .367800, .341316, .290885$

Range: $d = 0.21$ $\delta = 1.0$ δ limit mean from the 2 reductions

Oscillation (3) 1969 March 27 Z.C.1211 Mag 6.2 Dir. 8h PA 54°

Obs. LST in $8^h 18^m 49.0 + \cos$ (clock 3.35 sec) \therefore Cor. LST in $= 8^h 18^m 52.3$

i.e. $20^h 8^m 35.3$ V.T. δ limit derived from Cor. limit (N.A.O) -0.23

Plate measurement Apr 5.0 (Thomson)

Apr 4.88202 $6^h 55^m 48.37 + 68^\circ 11' 21.9$ (1950.0)

9 stars were measured, but one star was discarded; one of the remaining 8 stars was used twice over (i.e. in 2 of the 3 reductions)

BD $68^\circ 45.7, 67^\circ 46.8, 69^\circ 40.5$ $D_s = .551636, .294395, .153969$

BD $67^\circ 46.1, 68^\circ 46.0, 67^\circ 47.6$ $D_s = .165713, .656826, .177460$

BD $67^\circ 46.1, 68^\circ 45.9, 68^\circ 46.4$ $D_s = .268820, .491336, .239844$

Equal weight was given to each star in the first run: so BD $67^\circ 46.1$ (twice used) was given in the two cases only half the weight of the other 7 stars

Range $k = 4.2$ $\delta = 1.0$

1969

C. Thomas 1968 J

Mar 7.8 (night of Friday - Saturday) Oa-0 plate Transp. $\frac{5}{10}$ $T=43^\circ$ $E=13$
After several weeks of bad weather an attempt was made last night (Mon 6.8)
to photograph C/Thomas; but owing to its manner to the zenith there were difficulties
in finding a guide star before moon rose. Tonight (Mon 7.8) we got a good
guide star - but the position was too difficult to use the micrometer &
I had to guide with a fixed web (in hydrogen eyepiece & star diagonal).
This meant considerable trailing of comet (as it was moving $6''$ in 10 mins
towards PA $46^\circ \searrow$), so that the plate was only of use for position
measurement of comet. The plate shows trailed comet (60 minute exposure \rightarrow $37''$ trail)
but the relative size of central condensation & other comae appear much as in drawing
and the photometric integrated mag was estimated at $13.0 - 13.5$.

Exposure 60 mins. Co. LST mid exposure $7^h 2^m 1.0$. Plate centered at
 $5^h 39^m 12^s + 73^\circ 58'$. Observed R.L.V. & R.H. South for part of exposure.

April 5.0 (night of Friday - Saturday) Oa-0 plate Transp. $\frac{6-7}{10}$

C. Thomas 1968 J Exposure 60" Co. LST mid Exposure $gh 52^m 5.8$ Plate
centered at $6^h 48^m 5 + 68^\circ 53'$ drawn with moving web towards
PA $46.2 \searrow$ in 10 minutes ($9''$) steps. Comet appears as a small cross
with fairly sharp stellar condensation at intersection of arms - the straight arm
being $45''$ long in PA $70^\circ - 250^\circ$ - and the whole cross enclosing a faint halo of
light $40'' - 45''$ in diameter. Integrated photo mag estimated at 13.8

Plate Measurement Apr. 8.0 C/Thomas

April 7.89156 $7^h 3^m 8^s.73 + 67^\circ 31' 46''.1$ (1950.0)

9 stars were measured $67^{\circ}473, 67^{\circ}476, 68^{\circ}464$ $D_s = .169589, .718132, .112328$

B.D. $68^{\circ}459, 67^{\circ}468, 67^{\circ}482$ $D_s = .1266715, .314090, .419195$

B.D. $67^{\circ}461, 66^{\circ}488, 69^{\circ}405$ $D_s = .273157, .449147, .277695$

Fixed mean from 3 reductions. Range $\alpha = 2''.5$ $\delta = 2''.1$

Plate Measurement Apr. 9.0 C/Thomas

April 8.91127 $7^h 5^m 36^s.94 + 67^\circ 18' 16''.0$ (1950.0)

9 stars were measured $67^{\circ}476, 66^{\circ}488, 67^{\circ}480$ $D_s = .653802, .197073, .149125$

B.D. $67^{\circ}475, 68^{\circ}464, 67^{\circ}482$ $D_s = .611936, .166600, .221414$

B.D. $67^{\circ}473, 66^{\circ}490, 67^{\circ}483$ $D_s = .594527, .244055, .161417$

Fixed mean from the 3 reductions. Range $\alpha = 0''.9$ $\delta = 0''.7$

Plate Measurement Apr. 15.

Old plate with 90 minute exposure shows 16.0 mag star easily, and 16.4 fairly clearly.

1969

T 55.55° F9

April 8.0 (night of Monday-Tuesday) Oa-O plate Transp 6/10 deteriorating & 8" steps exposure
 E/Thomas 1968 J Exposure 35 minutes (stopped by cloud) dist mid Exp. 10^h 17^m 42^s. 4
 Plate centered at 7^h 1.0 + 67° 30'. Drove with moving web towards PA 44° ↘
 in 10 minutes (8") steps. Could image close to faint star. Central
 condensation near stiller 20" diameter with surrounding haze about 40" diameter
Est integrated mag about 13.8

April 9.0 (night of Tuesday-Wednesday) Oa-O plate Transp 7-8/10 T=55-50° F9
 E/Thomas 1968 J Exposure 90 min 10^h 50^m 7^s. 2 cor L&T mid Exp. Plate was
 centered at 7^h 8.0 + 67° 15' (1969) Drove with moving web towards
 P.A. 45° ↘ in 10 min (8" 0) steps. The plate is good with strong central
 image. Central condensation, very heavy, 20" diameter, faint diffuse outer zone
 about 3.0 diameter. integrated mag est at 13.8

N₂ Polar Seyma
 April 15.0 (night of Monday-Tuesday) Oa-O plate Transp. 7/10 T 45-42° F12
 Exposure 90 min cor L&T mid Exposure 11^h 34^m 56^s. 8
 Plate centered on Pole Star 2^h 2^m 30^s + 89° 7'

As there was some error in the clock-rate I drove in RA only on plates
 This reduced the clock-error from the Pole Star to all higher declinations
 to a negligible amount. Stars of mag 16.0 (including one of 16.02)
 were clear & could measure with 1/100" scale. Fainter stars including one of
 16.19 ^{mag} 8/16.40 were faint but quite definite.

All around pole, star-images show no distortion from circular aberration.
 Plate fogging is only very slight. The ^{mid} exposure was about 22 cm V.T.
 no width.

Plate Measurement May 19-0 of Thomas

May 18.98435 $8^h 33^m 35.99$ $+57^\circ 46' 30.8$ (1950.0)

Min stars measured (6 used for null)

not used \rightarrow B.D $58^\circ 1136, 57^\circ 1155, 58^\circ 1140$ $Ds = .127287, .718909, .153805$

used \rightarrow $\left\{ \begin{array}{l} \text{BD } 58^\circ 1130, 57^\circ 1154, 57^\circ 1166 \quad Ds = .409291, .331267, .259443 \\ \text{BD } 58^\circ 1125, 57^\circ 1159, 58^\circ 1145 \quad Ds = .365874, .454655, .179471 \end{array} \right.$

Discard mean from the 2 reductions used. Range $\Delta = 0.6$ $\delta = 0.8$

[Range from the 3 reductions (including discarded one) $\Delta = 1.5$ $\delta = 0.8$]

Oscillation (4) 1969 May 19. ZC1013 Mag 6.9 Dis D.L. PA = 139°

This oscillation was observed in duplicate by H.H. Morgan using

4" Wray and R.L.W using 6" Cohn

Orbit Dis RHW $13^h 23^m 1.5$

H.H.M. $13^h 23^m 2.0$

Orbit Dis (clock 0.93 slow $12^h 23^m 2.43$ RHW)

$13^h 23^m 2.93$

This gives U.T $21^h 43^m 32.4$ RHW and $21^h 43^m 32.9$ H.H.M.

Prob. residual from Or Limb from NAO: -0.16 RHW

-0.36 H.H.M.

1969

May 19:0 (night of Sunday-Monday) Oa-0 white. Transp 7/10 T_{43-40} F12
C/Thomson 1968 J Exposure 100 minutes $\text{LST mid Exp} = 15^{\text{h}} 33^{\text{m}} 1.2^{\text{s}}$
Plate centered at $8^{\text{h}} 35.2^{\text{m}} + 57^{\circ} 42'$. Scope \bar{c} moving west towards 43.4°
in 10 minutes ($9^{\text{h}} 0$) steps. Core appears fairly strongly condensed
Central condensation 20" diameter outer cone faint & diffuse 2' diameter
Thin, straight radial spike about 3' long in PA 0° . integrated mag est at 14.5

June 8:0 (night of Saturday-Sunday) Oa-0 white. Transp 6/10 mid-exposure twilight
considerable all through exposure

C/Thomson 1968 J Exposure 110 minutes. In spite of being centered close on
local midnight the summer twilight caused heavy fogging of plate.
Cos LST mid exposure was $17^{\text{h}} 8^{\text{m}} 7.1^{\text{s}}$. Plate centered at $9^{\text{h}} 12.5^{\text{m}} + 52^{\circ} 44'$
scope with moving west towards PA 42.1 in 10 min ($9^{\text{h}} 26$) steps.
There is absolutely no sign of the cone close to the ephemeris position - which
is known to be fairly correct. As can I think be the integrated
mag of the cone is fainter than 15.0 mag.

July 9 22 50 UT ($\text{LST } 17^{\text{h}} 50.5^{\text{m}} - 18^{\text{h}} 20.5^{\text{m}}$) Plate Oa-0 Transp 8/10
M13 Hermin Exposure 30 mins guided by D. Lacombe

The sky cleared suddenly and it was decided to let D.L. have a practice
session on M.13. There was insufficient time for the temperature to get steady -
hence poor focus. But guiding was good.

9/10/54

Plate Aug 6.0. This plate was ^{not} measured but owing to defects in ~~the~~ the
star image the results were discarded partly due to ~~the~~ function change
during exposure

1969

July 31.0 (night of Wednesday-Thursday)

? New Comet was showed up (? by Millman) to say that Alcock had seen on previous night a faint object which he thought might be a comet; but owing to moonlight & overcast, could he had been unable to confirm or discard it. I therefore made 3 exposures of time ranging between 5 and 15 minutes; as owing to rising nearly full moon it was not possible to judge an optimum exposure time as Alcock had suggested a motion of about 21" in 20 minutes in PA

14° 

Exp(1) 15 min on LST mid Exp. 18^h 12^m 36.^s4 moving web

Exp(2) Shutter jammed.


Exp(3) 5 min on LST mid Exp. 19^h 18^m 36.^s4 fixed web.

Both plates were fogged. Plate I very heavily.
(Known Plate II is good enough to show that no object near to mag 10 was present in the suspected location.

Aug 6.0 (night of Tuesday-Wednesday) Oa-0 plate. Transp 67/10 T=62-58° F=7

C/Kibrovich 1969 b This new comet was discovered July 29 at Bajardorf. The announcement was made July 29 - but this is the first clear night since then. Comet reported as mag 14.0 and daily motion available.

Exposure 45 min on LST mid Exp. = 18^h 52^m 30.^s7 Plate center 19^h 21.^m8 + 29^o 0' 1969.5

Drove moving web towards P.A. 22° 35'  in 10 min (10".8) steps.

Guided by H.H. Morgan.

Comet image very strong; quite circular disc 25" diam^x with almost sharp outline & no sign of surrounding outer coma. The ^x nucleus is slightly eccentric. Mag. 14.0

Plate measurement Aug 7.0 C/Kiboutch

Aug 6.92117 $19^h 18^m 35.44^s + 28^\circ 38' 28.9''$ (1950)

Due to their being a considerable discrepancy between the various reductions made from 9 stars (possibly due to comet being some distance from plate center) amounting to 4" or 5" these measurements were discarded & not sent to I.A.V.

Measurement of Plate Aug 8.0 C/Kiboutch

Aug 7.97731 $19^h 16^m 35.54^s + 28^\circ 44' 18.2''$ (1950.0)

Six stars were measured:

B.D. $28^\circ 3264, 28^\circ 3282, 28^\circ 3294$ $D_s = .132616, .746614, .120771$

B.D. $28^\circ 3261, 28^\circ 3262, 28^\circ 3295$ $D_s = .256067, .085636, .658897$

Range $\alpha = 0.13$ $\delta = 0.7$ Direct mean from the 2 reductions was used.

Measurement of Plate Aug 20.0 C/Kiboutch

Aug 19.92204 $19^h 54^m 55.56^s + 29^\circ 24' 39.2''$ (1950)

Nine stars were measured, B.D. $29^\circ 3394, 29^\circ 3409, 29^\circ 3422$ $D_s = .418272, .176997, .404731$

And \rightarrow B.D. $29^\circ 3382, 29^\circ 3410, 29^\circ 3425$ $D_s = .307257, .238994, .453749$

Discarded \rightarrow B.D. $29^\circ 3402, 29^\circ 3427, 28^\circ 3145$ $D_s = .704933, .151292, .143775$

The measurements were made on the center of the coma; but later it was noticed that the central condensation was slightly eccentric in the coma: requiring a correction of $\alpha - 0.05$ & $\delta - 1.5''$

The direct mean of reductions (1) & (2) was taken & the above correction added.

The 3rd reduction with poorer images was discarded

Range Pds (1+2) $\alpha = 0.8$ $\delta = 0.2$ Reduction (1) & (2) $\alpha = 2.9$ $\delta = 0.6$

1969

Aug 7.0 (night of Wednesday-Thursday) Oa-0 plate Transp 7-8/10 $T=60-57^{\circ}$ F7

C/Kishoutch 19696 Exh. 45 min cor LST mid Exh $18^{\text{h}} 57^{\text{m}} 30.7^{\text{s}}$

Plate centered at $19^{\text{h}} 14^{\text{m}} 0^{\text{s}} + 29^{\circ} 11.0'$ (1969) Drove with moving comb toward PA $14^{\circ} 15'$ in 5 min (8.5) steps. Guided by H.H. Morgan. (P.A. $14^{\circ} 15'$)

The comet is immediately following the trail of a bright star. The coma is much more diffuse than on previous night about 45" diameter, irregular in shape with faint nucleus. Est integrated mag = 14.0

August 8.0 (night of Thursday-Friday) Oa-0 plate Transp 6/10 $T=60^{\circ}$ F=7

C/Kishoutch 19696 Exh. 65 min cor LST mid Exh $20^{\text{h}} 20^{\text{m}} 31.1^{\text{s}}$
(22)

Plate centered at $19^{\text{h}} 11.0^{\text{m}} + 29^{\circ} 10'$ (1969.0). Drove with moving comb toward PA $14^{\circ} 15'$ in 5 min (8.5) steps R.L.W.

The comet image between two faint star trails is very like last night: diffuse and 50" diameter & elongated (? short tail) toward PA 145° , faint nucleus. Est integrated mag = 14.0

[Aug 14.0 C/Fujikawa See next page]

August 20.0 (night of Tuesday-Wednesday) Oa-0 plate. Transp 8/10 $T=58^{\circ}$ F=8

C/Kishoutch 19696 Exh. 40 min cor LST mid Exh $19^{\text{h}} 50^{\text{m}} 1.7^{\text{s}}$

Plate centered at $18^{\text{h}} 55.6^{\text{m}} + 29^{\circ} 28'$. Drove with moving comb toward P.A. $6^{\circ} 10'$ in 5 min (4.8) steps. Guided by H.H. Morgan.

Comet image is close alongside, but separated from a faint star trail. The coma now is again extremely condensed: a compact little disc 20" diameter without any sign of outer coma or tail. It is even smaller & more compact than on Aug 6.0 (see I.A.V. Circular 2166)

Plate measurement Sept 4.0 C/Khoralek

Sept 3-86895 $18^{\circ} 31' 44.28'' + 29^{\circ} 19' 57.0''$ (1950)

Nine stars were measured; but 3 were discarded

used \rightarrow B.D. $29^{\circ} 3280, 28^{\circ} 3020, 29^{\circ} 3302$ Ds = $.289408, .519885, .191507$

\rightarrow B.D. $29^{\circ} 3275, 29^{\circ} 3288, 28^{\circ} 3027$ Ds = $.101765, .760928, .137307$

Discarded \rightarrow B.D. $28^{\circ} 3016, 29^{\circ} 3285, 28^{\circ} 3030$ Ds = $.124768, .540478, .234753$

Range between used reductions $\Delta = 0.7'' \delta = 0.8$

(Range between all 3 reductions including discarded one $\Delta = 3.4'' \delta = 0.8$)

1969

(1969 Wed-Thursday)

Aug 14.0 / New Comet discovered by Fujikawa, at Onahara Japan Aug 12.7 Mag 11.0

C/Fujikawa 1969d Oa-O Plates Transp 6/10 Bal v. low altitude

and only a short time before Twilight

This was the only attempt made at Ootaka to photograph this comet as it was only visible low in the eastern sky just before dawn and was getting nearer to the sun. Two plates were exposed

Exp I 8 min on DOT mid exposure $23^{\text{h}} 39^{\text{m}} 53.4^{\text{s}}$


Exp II interrupted by clouds, intermittent between LST $0^{\text{h}} 12^{\text{m}} 0^{\text{s}} \rightarrow 0^{\text{h}} 21^{\text{m}} 0^{\text{s}}$

Totaling is all about $6\frac{3}{4}$ minutes.

Neither plate showed any sign of comet after a careful search
Comet certainly fainter than mag 11.

Sept. 4.0 (1969 Wed-Thursday) Oa-O Plate Transp 6/10 T=61-60 F=7

C/Kishimoto 1969b. Exposure 45 min on DOT mid Exposure = $19^{\text{h}} 32^{\text{m}} 30.4^{\text{s}}$

Plate centered at $18^{\text{h}} 32^{\text{m}} 5 + 29^{\circ} 19'$ (1969.7) Drift with focusing web towards PA 4°  in 5 minutes (3.9) steps. Guided by David Larcombe

The comet is very small but image is strong. The head is very compact & dense about 15" in diameter with no surrounding haze. There is a sharp tail: 2 straight faint narrow streams towards PA 150° 90" long & towards PA 140° 60" long.

Est. integrated Mag = 13.7

Plate measurement Sept 20.0 P/Honda Mura Pajdusotova 1969e

Sept 20-16810 $9^h 41^m 50^s.69 + 14^\circ 37' 28''.7$ (1950)

9 stars measured BD $15^\circ 2103, 14^\circ 2141, 14^\circ 2146$ Ds: $.756600, .733334, .110065$

BD $14^\circ 2129, 15^\circ 2107, 14^\circ 2143$ Ds: $.033537, .303858, .362575$

BD $15^\circ 2098, 14^\circ 2135, 15^\circ 2116$ Ds: $.444580, .298570, .256960$

The measurements are poor because of the wedge-shaped trails & the agreement between the 3 reductions was unusually poor

a direct mean of the 3 reductions was used

Range $\alpha = 3''.6$ $\delta = 3''.5$

Observations (5-9) 1969 see Next Page \rightarrow

Observation of Pleiades 5 observed: 1 Starhawk & 4 Refractors

Plate measurement Oct 2.0 C/Kishonlik 1969b.

Oct. 1.81431 $18^h 6^m 6^s.30 + 27^\circ 50' 5''.6$ (1950)

Nine stars measured \rightarrow BD $28^\circ 2926, 27^\circ 2961, 27^\circ 2969$ Ds: $.276557, .620941, .102505$

only 6 stars used \rightarrow BD $27^\circ 2956, 28^\circ 2928, 27^\circ 2974$ Ds: $.586433, .198293, .215276$

(3 discarded) BD $27^\circ 2946, 28^\circ 2939, 27^\circ 2975$ Ds: $.501440, .188672, .309888$

Range in 2 used reduction $\alpha = 0''.1$ $\delta = 1''.3$

Range in 3 reduction including discarded one $\alpha = 1''.7$ $\delta = 2''.1$

ly69

Sept. 20.0 (night of Friday-Saturday) P/Honda-Markos-Pajdoschova 1969e This comet was discovered by Markos on August 12 at mag 14 & only 6' from predicted position. Because of weather I was unable to try for this comet until this early morning. The comet was low in the sky shortly before dawn.

P/H-M-P 1969e Oa-O plate. Transp 6/10 - but low altitude

$T_{46^{\circ}-45^{\circ}}$ F12.

Exposure 20 mins. Coe DST mid Exh. $3^m 46^s$ in 31.5^s guided by H.A. Morgan. Drove with mirror wet towards P.A. $46^{\circ} 32'$ \nearrow in $2\frac{1}{2}$ min (7"0) steps. Plate centered on $9^h 42^m 25^s + 14^{\circ} 40'$ (1969.5).

The image of the comet is strong; moderately well condensed coma 30" diameter. Outer coma 2' diameter. Integrated mag est at 8.5

Venus is a few degrees S of the comet. The plate is rather poor for moment as owing to changing transparency the star trails are wedge-shaped.

Oct 2.0 (night of Wed-Thursday) Oa-O plate Transp 5/10 but passing Cloud $T_{47^{\circ}47'}$ F11.

C/Kohoutek 1969b Exposure 30 mins - Coe DST mid Exh. $20^m 4^s$ 0.6

Plate centered $18^h 7^m 1^s + 27^{\circ} 50'$ (1969.8) Drove with mirror wet towards P.A. $29^{\circ} 30'$ \searrow in 10 min (3"2) steps. An intended 40 min exp was stopped by cloud after 30 mins. Integrated mag est at 13.0

Coma small condensed 20" diameter Tail narrow fan in P.A. $120^{\circ}-130^{\circ}$ about 60" long

Occultation of Pleiades Sept 30 (night of Monday - Tuesday)

Occultation (5) 1969 Sept. 30 Z.C. 0541 Mag 4.0 Dis. Bright limb P.A. 84°

Observed Disappearance LST $2^h 25^m 30.5$ (+cor Clock 2.0 F) cor LST = $2^h 25^m 28.5$

\therefore U.T. Observed Dis = $2^h 0^m 45.5$ Passing clouds, star faded out, Modestly good quality.
(1.5 before prediction) [N.A.O. Prelim predicted from cor limb = $+2.98$ - bad]

Occultation (6) 1969 Sept. 30 Z.C. 0537 Mag 3.8 Rappharu D. limb P.A. 185°

Observed Rappharu LST $2^h 39^m 45.0$ (+cor Clock 2.0 F) cor Obs LST = $2^h 39^m 43.0$

\therefore U.T. Observed Rappharu = $2^h 15^m 6.3$ Passing clouds: Observation rather poor.

(0.3 before prediction - no prediction) [N.A.O. Prelim, used from cor limb $+24.4$ - hopeless]

Occultation (7) 1969 Sept. 30 Z.C. 0536 Mag 5.4 Rappharu D. limb P.A. 235°

Observed Rappharu LST $3^h 4^m 38.0$ (+cor Clock 2.0 F) cor Obs LST = $3^h 4^m 36.0$

\therefore U.T. Observed Rappharu = $2^h 39^m 55.2$ Passing clouds: Observation very poor.

(1.2 after prediction) [N.A.O. Preliminary predicted from corrected limb = $+5.17$ - v. bad]

Occultation (8) 1969 Sept. 30 Z.C. 0539 Mag 4.4 Rappharu D. limb P.A. 265°

Observed Rappharu LST = $3^h 27^m 12.0$ (+cor Clock 2.0 F) cor Obs LST = $3^h 27^m 10.0$

\therefore U.T. Observed Rappharu = $3^h 2^m 25.5$ Passing clouds: Observation fairly good.

(16.5 before prediction) [N.A.O. Prelim predicted from cor limb = -0.68]

Occultation (9) 1969 Sept. 30 Z.C. 0541 Mag 4.0 Rappharu D. limb P.A. 233°

Observed Rappharu LST = $3^h 42^m 18.0$ (+cor Clock 2.0 F) cor Obs LST = $3^h 42^m 16.0$

\therefore U.T. Observed Rappharu = $3^h 17^m 29.0$ Passing clouds: Observation fairly good.

(10.0 before prediction) [N.A.O. Prelim predicted from cor limb = $+0.40$]

All the above occultations were observed under difficult conditions due to
rubble, passing clouds, varying fluctuations in brightness of stars. The disappearance: very
early & the reappearance: late. For all reappearance measurements were set on
the position of star at reappearance. The first 2 occultations (in BAA Handbook)
& the last 2 were not observed: we got 5 out of 9.

1969

Oct. 3.0 (night. Thursday - Friday) Oa-O plate Trump 7-8/10 T: 57°-53° F=9

C/ Khorvich 1969 b Exposure 30 min Cor BST mid Exp 20^h 52^m 59^s.6

Star with moving web towards P.A. 29° 46' ↘ in 10 min (2"9) steps.

Plate centred at 18^h 6^m 8^s + 27° 47'. This plate was taken because of the great transparency (this is poor seeing) to compare state of comet tail with that of previous night. No measurement for position was made.

Coma condensed 20" diameter with little outer coma. Short jet tail 60" long in P.A. about 115°

Oct 14.8 (night) Tuesday - Wednesday Plate Oa-O. Trump 7/10 But v. low alt. & Skpt. Tashq

C/ Tago - Sato - Kasaba 1969 g This new comet was discovered

on Oct 10.4 by the above 3 astronomers at Tokyo Observatory. Tonight was the first opportunity for observing it: but it was extremely low above near the sunset and moving South. It was only possible to give short exposure.

Exposure 10 min Cor BST mid Exp was 20^h 15^m 6^s.1 T=56°-55° F=9

Intended to observe with moving web towards PA 55° 30' ↘ in 5 min (8"2) steps.

Plate centred at 16^h 27^m 6^s - 4° 39' (1969). The plate shows a definite image of comet. The central coma only moderately condensed about 30" diameter.

The outer coma is very diffuse about 4' diameter. Int. guided mag at about 9.5 guiding was difficult as only good guide star was v. faint - observed between webs.

on fixed web. C/ Tago - Sato - Kasaba 1969 g :-

Plate measured give: Oct 14.77958: 16^h 23^m 16^s.61 - 4° 24' 37".5 (1950)

Min stars measured B.D. -4° 4107, -3° 3939, -4° 4110 Ds: 222167, 255514, 522318 } find mean of 3 reductions

BD -4° 4102, -3° 3940, -4° 4113 Ds: 245978, 257573, 496450 } Ray 2 = 2"2 δ = 1"7

BD -4° 4101, -3° 3943, -4° 4118 Ds: 425427, 297884, 266689 } alt. = 15° 50' 40" Refs. negligible.

Plate measurement: Oct 15.0 C/Khorshid 1969b

Oct. 14.84952 $18^{\circ} 2' 11.47'' + 27^{\circ} 9' 9.2'' (1950)$

Nine stars were measured:

B.D. $26^{\circ} 3158, 27^{\circ} 2944, 26^{\circ} 3180$ Ds = $.383335, .441466, .175210$

B.D. $26^{\circ} 3149, 27^{\circ} 2941, 26^{\circ} 3184$ Ds = $.410958, .298963, .290078$

B.D. $27^{\circ} 2938, 26^{\circ} 3168, 27^{\circ} 2961$ Ds = $.543062, .170917, .286019$

Best mean of the three reductions were adopted.

Range $d = 2.0$ $S = 2.1$

Occult (10) 1969 Oct 20 Z.C. 3256 Mag 6.2 Dec. 82. P.A. 104°

Observed in LST $21^{\text{h}} 21^{\text{m}} 31.1$ (+on clock 0.25 hrs) or in LST $21^{\text{h}} 21^{\text{m}} 31.3$

Obs. U.T. Dec 19^h 35^m 12.6 very good quality, observation.

(5.4 before prediction) [N.A.O. position recorded from cos. limb = $+1.24''$]

This occultation was observed in duplicate by H. H. Morgan using 4" Wray (same clock) (Position was $19^{\text{h}} 35^{\text{m}} 12.9$ U.T.)

[N.A.O. limb recorded from cos limb = $+0.14''$]

Plate measurement Oct. 29.8 C/Khorshid 1969b

Oct. 29.80818 $18^{\circ} 3' 4.38'' + 26^{\circ} 44' 27.2'' (1950)$

Nine stars measured B.D. $26^{\circ} 3158, 26^{\circ} 3168, 26^{\circ} 3180$ Ds = $.525571, .262936, .211491$

B.D. $27^{\circ} 2938, 26^{\circ} 3159, 26^{\circ} 3184$ Ds = $.339112, .303649, .357240$

B.D. $26^{\circ} 3148, 25^{\circ} 2426, 27^{\circ} 2961$ Ds = $.365674, .276142, .358295$

Best mean of the 3 reductions were adopted.

Range $d = 2.1$ $S = 2.2$

1969.

Oct. 15.0 (night of Tau-Wed) continued.

C Kishimoto 1969b Oa-O Plate Transp 6/10 T=55-57° F9.

Exposure 30 min. around LST of Expos 21^h 46^m 6^s.1 Motion at moment is negligible in 30 minutes - so drove with fixed web. Plate centered at 18^h 2^m 54^s + 26° 48' Central condensation strong about 20" diameter

Tail in PA 90° about 45" long. From the central condensation two curved streamer curves towards PA 135° and PA 180° each about 45" long

& both rather narrow - especially those towards PA 180°

Drizzle suggesting a spiral arrangement. Est integrated Mag = 12.8



Oct 29.8 (night of Wednesday-Thursday)

I C/Tago-Sato-Kosaka 1969g Plate Oa-O Transp 7/10. But too light

in view of the good transparency an attempt was made to

get a further plate of this comet for position before it got too far below the equator. This was specially desirable as the comet is due to return to the northern hemisphere early next year as a fairly bright comet - and the orbit is so far poorly determined.

T=50° F11

Exposure 4 minutes. Corrected LST mid Expos = 20^h 36^m 58^s.9

Drove with free drive. Plate centered 16^h 38^m.5 - 12° 16'

Owing to the short exposure, too light & extremely low altitude no sign of comet

II C/Ushworth 1969b Oa-O Plate. Transp 6-7/10 T=46-44 F11

Exposure 15 min (moon was rising) LST at mid Exp = 21^h 45^m 32^s.9 Drove with fixed web (motion negligible). Plate centered at 18^h 4^m.2 + 26° 46'

Good image faint central condensation 15" diameter Outer coma 30" diameter. Then a streamer a spoke towards PA 140° & a linear one to PA 220°. No definite tail. Est. integrated Mag = 12.6

on plate

Plate Measurement Nov. 5.0 P/Fayy 1969 a

Nov. 5.015182 $\alpha = 5^h 29^m 40^s.83$ $\delta = +10^\circ 30' 1''.7$ (1950)

Nine stars measured B.D. $+10^\circ 803, 10^\circ 806, 11^\circ 854$ Ds: .185798, .682875, .131328

B.D. $+10^\circ 798, 11^\circ 837, 9^\circ 892$ Ds: .333807, .441284, .224909

B.D. $+10^\circ 795, 11^\circ 853, 9^\circ 876$ Ds: .363572, .424494, .211935

Direct mean of the 3 reductions. Range $\alpha = 1''.3$ $\delta = 1''.2$

Plate Measurement Nov. 26.7 C/Khoult

8 Nine stars measured: B.D. $+27^\circ 2994, 27^\circ 3000, 27^\circ 3001$ Ds: .168279, .430202, .421519

B.D. $28^\circ 2971, 27^\circ 2997, 27^\circ 3005$ Ds: .251941, .311914, .436146

B.D. $28^\circ 2970, 27^\circ 2995, 27^\circ 3002$ Ds: .190555, .138943, .670500

Nov. 26.76188 $18^h 17^m 22^s.59$ $+27^\circ 41' 55''.2$ (1950)

Direct mean of the 3 reductions accepted.

Range $\alpha = 2''.0$ $\delta = 2''.9$

1969

Nov. 5.0 (night of Tuesday - Wednesday)

P/Faye 1969 a This comet was recovered by Elizabeth Roemer on May 17 quite close to its predicted position at May 18 - about 1 mag fainter than was predicted. I stupidly failed to try to look for this comet sooner expecting its magnitude to continue 1 mag fainter than predicted; so that by Nov 5 it should still be about mag 13.6 - when I decided to have a go.

Exposure 10 minutes Plate Oa-O Transp 7/10 $T=37^{\circ}$ $F 14.5$
Cor LST mid Exp = $3^{\text{h}} 8^{\text{m}} 6.2^{\text{s}}$. As comet was moving at $6.4''$ in 10 mins towards PA $62^{\circ} 30'$ \swarrow I was able to drive with fixed web. Plate was centred at $5^{\text{h}} 30^{\text{m}}.4 + 10^{\circ} 40'$

The comet image is faint on the plate: diffuse coma $30''$ diameter, with near-stellar central condensation - not very dense. Integrated Mag est at 11.5

Nov 26.7 (night of Wednesday - Thursday)

C/K Shortish 1969 b Oa-O Plate Transp 6-7/10 near-full moon rising
 $T=32.5$ $F 15$.

Exposure 20 minutes. Cor LST mid exposure = $22^{\text{h}} 29^{\text{m}} 4.4^{\text{s}}$. Drive with moving web towards PA 26° \swarrow in 10 min ($4.5''$) steps. Plate centred at $18^{\text{h}} 18^{\text{m}}.0 + 27^{\circ} 42'$ (1969.7) The comet is rather faint: central coma $30''$ slightly hazy mediantly, concentrated ~ no outer coma. Tail about $60''$ in PA 80°

Est hazy Mag = 12.5

Plate Measurement Nov. 29.8 C/Kohoutek

Nov. 29.75440 $\alpha = 18^h 19^m 45^s.84$ $\delta = +27^\circ 58' 30''.6$ (1950)

Nine stars were measured B.D. 27°3001, 27°3005, 28°2985 Ds: .173953, .300194, .525853

B.D. 28°2978, 27°3002, 28°2993 Ds: .151166, .465779, .383056

B.D. 27°3000, 28°2979, 27°3017 Ds: .360011, .243454, .396533

Range $d = 0''.4$ $\delta = 2''.2$

Plate measurement Dec. 4.75 C/Kohoutek 1966.

Dec. 4.75108 $\alpha = 18^h 24^m 4^s.58$ $\delta = +28^\circ 31' 8''.8$ (1950)

Nine stars were measured, but only six were finally used:

B.D. 29°3253, 28°2993, 28°3004 Ds: .131168, .398458, .470375

used \rightarrow B.D. 29°3248, 28°2995, 28°3011 Ds: ~~.398458~~ .228728, .306856, .464416

discarded \rightarrow B.D. 29°3252, 27°3017, 28°3003 Ds: ~~.470375~~ .201110, .342320, .456572

Range over the 2 reductions used $d = 0''.1$ $\delta = 0''.1$

over all 3 reductions including the discarded one $d = 2''.4$ $\delta = 0''.1$

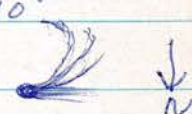
1969.

Nov. 29.8 (night of Saturday-Sunday) Oa-0 plate Transp 6-7/10 $T=30-30^{\circ}$ $F=16$

C/Kishonich 1969 6. Exposure 20 minutes CorLST mid Expt $22^{\text{h}} 30^{\text{m}} 6.6$

Drone with mirror, wheel turned PA $28^{\circ} 45'$ \nearrow in 10 min (4.5) steps.

Plate centered at $18^{\text{h}} 21.5^{\text{m}} + 28^{\circ} 9'$ (1969)

The comet image is rather faint & is partially obscured by a star almost on top of it. The coma faintly concentrated is about 30" diameter & short tail extends to PA 90 extending 60" + (cut off by star); there are also two or three very fine curved streamers curving up towards PA 180° and up to 120" long. The appearance is similar to that of Oct. 15 - again suggesting a spiral structure.  List interpreted May 12.4

Dec. 4.75 (night of Thursday-Friday) Oa-0 plate Transp 6-7/10 $T=31-30^{\circ}$ $F=16$

C/Kishonich 1969 6.

Seen, v.v. bad.

Exposure 20 mins CorLST mid Exposure = $22^{\text{h}} 45^{\text{m}} 2.2$ Drone with mirror wheel towards PA 31.5 \nearrow in 10 min (5.9) steps. Plate was centered at $18^{\text{h}} 24^{\text{m}} 45^{\text{s}} + 28^{\circ} 32'$. Central condensation 20" diameter, outer coma about 40" diameter. Short tail 60" long PA 90 - vague & diffuse in middle of a broad diffuse fan extending from about PA 40° - PA 120° - the foggiest & lack of detail in tail probably due to extremely bad seeing. The main part of the fan tail lies between two streamers: the chief one 2" long PA 110° & a smaller, shorter one in PA 70° . List interpreted May ≈ 12.1

1969 Dec. 9. Polaris Seymer Plate under a Pole Star

On - 0 plate.

ten Magnitude Scale

5 Exposures of varying length were given, moving the plate slightly between exposures

Exp(1)	Exposure 9 mins 0 sec.	LAST $1^h 27^m 1^s - 1^h 36^m 1^s$
(2)	3 min 0 sec.	$1^h 37^m 1^s - 1^h 40^m 1^s$
(3)	1 min 0 sec.	$1^h 41^m 1^s - 1^h 42^m 1^s$
(4)	20 secs.	$1^h 43^m 1^s - 1^h 43^m 21^s$
(5)	7 sec.	$1^h 44^m 1^s - 1^h 44^m 8^s$

1969

Occultation (11) 1969

Dec. 13. Z.C. 3171 Mag 3.8 Durchmesser D.L. P.A. 10°

Observed Dis L.S.T. $23^h 33^m 42.2$ R.L.W. $23^h 33^m 42.5$ R.H. South (+ cor 9.3 Clock fast)

Cor. Obs. Dis L&T $23^h 33^m 32.9$ R.H.W. $23^h 33^m 33.2$ R.H. South

∴ U.T. Obs. Dis $18^h 14^m 33.4$ R.L.W. $18^h 14^m 33.7$ R.H. South

(3.4sm after Pred.) (good quality, duration) (3.7 after Pred.)

[Prelim Prediction from Cor. Lin = +0.20 Min = +0.12 R.H.S. — N.A.O.]

Occultation (12) 1969

Dec. 14 Z.C. 3310 Mag 6.4 Durchmesser D.L. P.A. 17°

Observed Dis L.S.T. $22^h 11^m 59.0$ (+ cor 8.95 Clock fast)

Observed Cor. L&T $22^h 11^m 50.15$

∴ Observed Dis U.T. = $16^h 49^m 8.5$ (3.5 before Pred.) ~~good quality~~

(only moderately good quality: owing to slip in calculating Pred Time of Difference & making it 1^m 30 sec too late, I was not expecting Dis. so soon & was taken unawares!)

[Prelim Prediction from Cor. Lin = -0.31 N.A.O.]

1969 Summary of Observations during the Year.

Plot Measured C/Khorakli

Jan 4.75657 $19^{\circ} 1^{\prime} 7.99^{\prime\prime} + 34^{\circ} 52' 3.4^{\prime\prime}$ (1950).

Nine stars were measured and reduced but only 3 stars (one reduction)

was used, as the other images were poor - mainly due to plate being over of center then

used. \rightarrow B.D $34^{\circ} 3399, 34^{\circ} 3401, 34^{\circ} 3409$ D_s = .071088, .858846, .070066

B.D $34^{\circ} 3388, 34^{\circ} 3403, 34^{\circ} 3410$ D_s = .503319, .214508, .282173

discarded \rightarrow B.D $34^{\circ} 3396, 35^{\circ} 3456, 34^{\circ} 3412$ D_s = .248847, .361028, .390126

Total range (3 reductions) $L = 3.4$ $S = 2.4$

(The reduced value was in $L = 2.2$, in $S = 1.3$ from ten mean of the 3 reductions).

1970.

Jan 4.75 (night of Sunday-Monday)

0a-0 plate Transit 5/10 some light cloud.

C/Khorshid 1969 G. Exposure 20 minutes. cor LST mid Exposure = $4^h 24^m 55^s 10.4$

Swiss with moving web towards P.A. $43^{\circ} 45'$ in 10 min (8.1) stars.

Plate arrived at $19^h 12^m 35^s + 34^{\circ} 54'$ (1970)

The plate is not quite perfectly focused & the drawing is rather poor.

The image of the comet is faint. Central condensation $20''$ faint stars - there is only the faintest suggestion of an outer coma. There is also a suggestion of a very faint tail towards P.A. 110° & 140° (two separate streams). The comet appears definitely fainter than when last observed (Dec 4) and log Mag est at 12.5.

Jan 24.75 (night of Saturday-Sunday). This was the last opportunity here for observing

C/Tapp-Sato-Kosaka since last seen before going too far South on Oct 14.8

It is now beginning to come north within reach of this latitude. It was in conjunction on Dec. 21 but is still predicted to be around mag 3.6 ^{- about 1 mag fainter than that} ~~at~~ conjunction.

① Tapp-Sato-Kosaka 1969 g 0a-0 plate. V_g low altitude near sunset and a great deal of passing cloud.

We only managed to get one short exposure between clouds R.H. South garden.

Exposure 1 min. cor LST mid Exposure $4^h 43^m 59^s 0.7$ Swiss is fixed web.

Plate arrived at $1^h 2^m 40^s - 1^{\circ} 30'.5$. The sky clouded over again so no gaps were large enough to be worth using.

Very few stars are visible on the plate; but the comet image is sharp near centre. Outer coma about $2.5'$ diameter gradually condensing towards centre. Definite short exposure a suggestion of a tail towards P.A. 340°

Plate not worth measuring.

log Mag est at 4.0

1970.

Jan 25.75 (night of Sunday - Monday)

1

C/Tape - Satō - Koraka 1969 g. The weather was rather helplessly unsettled - a tremendous lot of cloud with only few & mainly very short gaps occurring incredibly. We managed to get 2 exposures - one very short and another of about 10 minutes (but even that was greatly reduced by passing cloud. The short exposure was not good enough for precise measurement; and the longer exposure not so long as we hoped for & not long enough to show very much in the way of tail structure.

Exposure I 1 minute $\alpha = 0$ hlat. Passing cloud, Poor Transparency.

on L.S.T. mid Exposure = $2^{\text{h}} 44^{\text{m}} 31.3^{\text{s}}$. Stars fixed well. Plate centered at $1^{\text{h}} 11^{\text{m}} 36^{\text{s}} + 2^{\circ} 0'$. Image of comet in clear Central Zone $30''$ diameter & outer coma 2.5 diameter. Circular, no sign of tail.

Exposure II 10 minutes $\alpha = 0$ hlat. Passing cloud & poor transparency.

on L.S.T. mid Exposure = $3^{\text{h}} 18^{\text{m}} 19.7^{\text{s}}$ Stars with moving well towards PA $58^{\circ} 40'$ in 1 minute ($10''$) steps. Plate centered at $1^{\text{h}} 11^{\text{m}} 36^{\text{s}} + 2^{\circ} 0'$ (1970)

The stain trails are broken & wedge shaped due to passing cloud

The inner coma strongly condensed about 2.5 diameter. Outer coma traced to about $6'$ diameter. Tail about 1° long in PA 75° . Faint & very diffuse fan shaped - but center narrow. No definite structure.

1970

Jan 26.75 (night of Monday-Tuesday) One bad luck with the weather continues. Continuous uphill passing cloud with short unpredictable gaps. In the context of selecting a site (plus the very rapid reaction & 1 minute diving steps) the RA was unconsciously set - so the comet is right on edge of plate - (also in the process of changing the position of the coils part of the exposure seems to have been made moving along one direction web & part along the other direction web!)

① / Tago-Sato-Kosaka 1969 g Oa-0 plate Transit 6/10 but much interruption due to passing cloud. Exptm = 23 min / ^(17 min) cor LST mid Expt = $2^h 52^m 29.5^s$
Drove with moving web towards PA $58^\circ 5'$ Δ in 1 min ($9'' 5$) steps.

Plate roughly centered (near edge of plate) was near $1^h 29^m 50^s + 5' 33''$
The heavily condensed inner coma was 2.5 diameters. Outer coma to about 7.5 diameters
Tail about $1^\circ 30'$ long in P.A. 75° - Tail narrow at root about 1.5 width spreading out into a narrow diffuse fan. (Tail extends beyond edge of plate)

Feb 4.75 (night of Wednesday-Thursday) This was the first night with a clear unbroken sky since this comet came into view on 7 Jan 75!

① / Tago-Sato-Kosaka 1969 g

Two exposures were given one short 3 min & one long 35 min: Oa-0 Transit 6/10

1) Expt 1) Expt. 3 min wr. LST mid Expt. = $3^h 10^m 29.5^s$. Drove with fixed web.

Plate centered at $2^h 16^m 50.3^s + 26^\circ 28'$ Comet very strong - definitely trailed

Outer coma about 5' diameter faint tail towards PA 85° about 10' long.

2) Expt. 2) Expt. 35 min wr. LST mid Expt. $4^h 24^m 29.5^s$. Oa-0. Transit 6-7/10 T $37^\circ 35'$ F 14

Drove towards PA 54° Δ with moving web in 1 minute ($4'' 7$) steps. Plate

centered at $2^h 16^m 50^s + 26^\circ 28'$. The long exposure shows heavily condensed inner coma about 4' diameter, and an outer coma quite transverse diameter of about 10'. The outer coma

Plate Mammak Minor Hunt (for identification) (Plate I 3 min. exposure)
Feb. 4.76564 $L = 2^h 10^m 45.61^s$ $S = +25^\circ 30' 54.8''$ (1950)

Six stars were measured and the direct mean of the 2 reductions taken as result.

B.D. $+25^\circ 363, 24^\circ 324, 25^\circ 377$ $D_s = .379266, .418390, .202345$

B.D. $+25^\circ 358, 24^\circ 319, 25^\circ 378$ $D_s = .282321, .424241, .293437$

Range $L = 2.2, S = 0.7$

1470

Feb. 4. 75 (cont.) Plate to C/Topo Sato Ranch & Minor Planet ? What shows some indication of being "hooded," symmetrically with respect to direction of tail. The main tail is a very diffuse streamer towards P.A. 80° & 2.5 long - it appears to be double in part of its length, a shorter & even more diffuse tail is about 0.5 long centered around P.A. 130° . There is no sharp detail as within the head or the tail of the comet.

? Minor Planet. During the examination of this plate it was noticed that 3 star-trails $1\frac{3}{4}$ S of the comet - & just to the right of plate center - were definitely not quite parallel to one another. The star trails are long (about $2\frac{1}{2}$); and it was clear that the central trail must be of a moving minor planet (there was no hypothesis about it, which excluded another comet). It was then found possible to locate the object on the 1st of the 2 plates (3 mins exposure) which had been taken about 5 minutes earlier. From this an approximate daily motion was deduced. (As this drawing was not made till the following morning it could not be further confirmed to same night.) The exact exposure was measured for precise position.

Feb. 5 ? Minor Planet. We contacted Milton & also Taylor (H.M. Naval Res. Office) as to the position, daily motion & approximate magnitude also set out to identify which ~~one~~ it was.

Feb. 5. 75 (night of Thursday - Friday). The sky was overcast, but there were very occasional & transient gaps. So the telescope was set on the indicated position & I waited for a chance to expose. We managed to expose one plate but only for a few seconds before it clouded up. LST Exp. about $3^h 24^m 0^s$. Centered on $2^h 14^m 50^s + 25^\circ 38'.0$. Owing to short exposure Minor Planet was at extreme limit of visibility on plate & could not be used for definite measurement.

Plate Measurement of ? Minor Planet (Plate II)

Fl. 6-80104 $\lambda = 2^h 13^m 25.07^s$ $\delta = +25^\circ 29' 17''.2$ (1950.0)

Six stars were measured

B.D. $+25^\circ 36.3, +24^\circ 33.0, +25^\circ 37.7$ Ds: $.159697, .231089, .609215$

B.D. $+24^\circ 32.4, +26^\circ 37.8, +25^\circ 37.8$ Ds: $.470890, .147643, .381462$

The direct mean of the 2 reductions was taken as the result.

Range $\lambda = 1''.9$ $\delta = 1''.7$

Plate Measurement of ? Minor Planet

Fl. 7-78585 $\lambda = 2^h 14^m 43.11^s$ $\delta = +25^\circ 28' 36''.4$ (1950)

Five stars were measured, one of which was used in both reductions

B.D. $+24^\circ 32.4, +25^\circ 37.7, +25^\circ 37.8$ Ds: $.215569, .270416, .514015$

B.D. $+24^\circ 32.5, +25^\circ 37.7, +24^\circ 33.5$ Ds: $.0171557, .503537, .324905$

A straight mean of the 2 reductions was taken as final result.

Range $\lambda = 3''.0$ $\delta = 1''.3$

1970

Feb. 6 (Snow) (Plate for identification of minor planet) Passing clouds are frustrating but larger gaps and were able to get 2 good exposures on ? minor planet.

Feb. 6-75 (mid of Friday-Saturday) Oc-Ophiu Transit 6/10 $T=35^{\circ}-33^{\circ}$ $F=15$

Exposure I 6 mins. on SST mid Exp. $3^h 28^m 0^s.1$

Exposure II 3 mins on SST mid Exp. $4^h 9^m 30^s.1$

Both exposures were driven with fixed wire and plate centered on $2^h 14^m 0^s + 25^{\circ} 38'.5$

The ? minor planet gives a good image on 6th plate & has disappeared from the position occupied on Feb. 4-75. The shorter exposure Plate II was measured for precise location.

The comet is visible on 6th plate toward the North following course of plate.

Feb. 7-75 (mid of Saturday-Sunday) Third plate for identification of minor planet.

Minor Planet Exposure 7 mins. Passing cloud 2 mins, 5 mins clear. Oc-Ophiu Transit 4-5/10

on SST mid Exposure = $3^h 51^m 30^s.3$. Grove with fixed wire Center $2^h 14^m 0^s + 25^{\circ} 38'$

again we were lucky: under passing cloud prior to exposure & clouding completely soon after it. The image of the minor planet is good.

The photographic magnitude of this object on all plates taken was estimated at 11.0 - about .5 mag fainter than $R.D. 25^{\circ} 378$

I sent these 3 positions to Marsden: none of the published ephemerides agreed with the object - so he calculated an ephemeris orbit from my 3 positions & finally identified it by the elements. It turned out to be Interamnia No. 704 & had been in opposition on Oct 27, 1969.

Occultation 1970 (1)

1970 Feb. 15 Z.C. 885 Mag 5.6 Distance 8.1. P.A. 132°

Observed disappearance LT 6^h 59^m 22^s.5 + cor (clock 1.5 min.)

cor LT 6^h 59^m 24^s.0

∴ Observed V.T. of disappearance = 21^h 27^m 33^s.1 good quality observation

(2.9 before predicted V.T.) Preliminary residual from cor limb (N.A.O.) = +1.23

In spite of my having classed this as "good quality" the N.A.O. residual is surprisingly large!

1970.

Feb. 26^g night of ^{Wednesday-Thursday} ~~Wednesday-Thursday~~ Two plates exposed (1) C/Kashanich (2) C/Tayo-Sato-Kascha.
(1) C/Kashanich ⁽¹⁹⁶⁹⁻⁶⁾ Exposure 22.5 min. Ca-0. Transh 7/10 T 34°-33° F 15
Cor. Δ T mid-exposure = $5^{\text{h}} 41^{\text{m}} 18^{\text{s}} 0$, Snow with moving well towards
P.A. $39^{\circ} 15'$ Δ in $2\frac{1}{2}$ minutes ($6^{\text{m}} 0$) steps. Plate centered at
 $L = 21^{\text{h}} 35^{\text{m}} 9$ $\delta = +60^{\circ} 58'$

Plate shows v. strong image of comet: Heavily condensed inner coma $30''$ diameter with
little or no outer coma. The tail is in form of a wide diffuse fan extending
between P.A. 29° & 10° . The strongest part is a wide streamer in P.A. 23° , $2'$ long.
The edges of the fan are heavily condensed to form 2 bordering streamers
about P.A. 29° & P.A. 10° about $1'$ long. Est. integrated Photo mag = 12.0

(2) C/Tayo-Sato-Kascha 1969 g. Ca-0 Plate Transh 7/10 T 33°-32° F 15.
Exposure 20 min. Mid cor. Δ T Exp = $6^{\text{h}} 54^{\text{m}} 3^{\text{s}}$. Snow with
moving well towards P.A. $37^{\circ} 23'$ Δ in 5 minutes ($7^{\text{m}} 8$) steps.
Plate centered at $3^{\text{h}} 32^{\text{m}} 5 + 43^{\circ} 0'$

The image is strong: Inner coma about $45''$ diameter heavily condensed. Outer coma
about $2.5'$ diameter. There is a very diffuse wide fan-shaped tail with an angle
of about 80° centered about P.A. 130° and extending widely for about $7'$.

March 3.0 (night of Friday-Tuesday) Ca-0 Plate Transh 6-7/10 T = 34° F 15
C/Tayo-Sato-Kascha 1969 g. Exposure 20 min. Mid cor. Δ T Exp. $5^{\text{h}} 48^{\text{m}} 1^{\text{s}} 4$
Snow with moving well towards P.A. $33^{\circ} 5'$ Δ in 5 minutes ($6^{\text{m}} 8$) steps. Centered at $3^{\text{h}} 45^{\text{m}} 1 + 44^{\circ} 38'$
Comet image shows inner coma about $30''$ diameter faintly heavily condensed with outer
coma to about $3'$ diam. Coma appears circular - no suggestion of tail.

1970.

Mar 6.0 (night of Thursday - Friday) Two plates were exposed on the comet


B/Tape Sato-Karacha 19.69 g Oa-O plate

I Exposure 20 min (R.L.W.) Tranch 6-7/10 T=32°-30° F. 16

Mid exposure 6^h 33^m 59^s.3

II Exposure 20 min (H. Morgan) Tranch 7/10 T=30°-29° F. 16.

Mid exposure 7^h 19^m 59^s.3

Both plates driven with moving web towards PA 31°20'  in 5 minutes (6.3) steps; and centered on 3^h 53.0 + 45° 30.5 (1970)

~~Results~~ The comet image is strong on both plates and differs from one another only in minor respects.

Plate I Inner coma faintly heavily condensed 30" diameter. Outer coma is about 2.5 in diameter but is diffusely elongated towards the f side with a broad radial streamer-like condensation extending for about 1' in P.A. 90°

Plate II Inner coma faintly heavily condensed 30" diameter. Outer coma is about 2.5 in diameter and as in plate I is diffusely elongated on the f side; but there is now no sign of the radial concentration towards P.A. 90° Instead there is now a fine linear spike in P.A. 350° 1.5 long.

Mar 7.0 (night of Friday - Saturday) Oa-O plate Tranch 6/10 T=35°-32° F. 15


C/Tape Sato-Karacha 19.69 g Exposure 30 min Mid exposure 6^h 16^m 59^s.5
Exposure with moving web towards P.A. 30°40'  in 5 min (6.2) steps.

Plate centered at 3^h 55.2 + 45° 43'

The comet image is strong; but is unfortunately superimposed on 3 of web of faint stars. One can only say the appearance is essentially same as last night: inner coma 30" diameter, outer coma 2.5 diameter & elongated on f side.

1970

March 28. C/Bennett 1969 i, diameter Sec 1969 in southern hemisphere; due to each perihelion 0.54 A.U. around March 20 and achieve first magnitude brightness. Up till end of March 1970 would only be observed in S. hemisphere. The first possible date for viewing it at work was March 28 when it would be just above N. horizon just before dawn. Unfortunately on that date it would be out of reach of 6" lens, behind a tree.

Visual description only: a flecking glint of ten comet was got through field glasses & the branches of a tree. It appeared about magnitude 1.0 or other description possible.

The early mornings of March 29, 30 & 31 were completely overcast.

April 16 C/Bennett 1969 i Kodak Da-0 plate (Whole Plate). Transp 7/11 but thin hazy cloud
E.V. of March 30 - April 1 (incl. of Tues - Wednesday) $T = 33^{\circ} - 32^{\circ}$ F 14.

Comet v. bright (est. about mag 1.0) visible twilight considerably slightly above tail N.E.

Exposure 7 mins. for $\text{Mid-L.S.T.} = 16^{\text{h}} 16^{\text{m}} 32^{\text{s}}.4$ approx V.T. $1^{\text{h}} 3^{\text{m}} 50^{\text{s}}$

Score with moving web on comet's nucleus (centered in micrometer web-square)

Motion in 1 minute = 7.2 towards P.A. $76^{\circ} 24'$ Δ . Approx Plate center = $22^{\text{h}} 22^{\text{m}} 2 + 19^{\circ} 45'$

Comet was fairly low altitude R.L.W. drove - assisted by Mark Ethington.

Observations from here written up by O. G. Buczyński from
of the Cadair Snow Observatory in Lancashire

(C/BANKS 1969 I)

1970 APRIL 2.17 (AT 2.15574) (clock 0.3 slow) + density strip
exposed 5 min

α 22^h 25^m 10^s δ = +22° 27.5 (1970)

P.A. 75° 12' circ. = 21° 28'

8 min = 2148 KMS = 55.3

in $\frac{1}{2}$ min = 6.9 steps

exposed 21 min R.W. drive on nucleus T = 28 f 17

TRANSPARENT $\frac{4}{10}$ L.S.T. 16^h 4^m 3^s - 16^h 25^m 3^s exposed = J.T. 3^h 47^m 16^s

(C/BANKS 1969 I)

1970 APRIL 4.17 (416042) density strip exposed 5 min.

α 22^h 31^m 45^s δ 27° 32.5 (1970)

P.A. 72° 40' circ. = 21° 0'

8 min = 51.2 = 1.99 KMS

(in $\frac{1}{2}$ steps = 1 min = 6.4 steps)

partly covered transparency 5-6%

exposed 40 min M.S.A. drive on nucleus

L.S.T. 16. 9.12 - 16. 49.12 T = $\frac{36}{35} + 14$

exposed 15^h 51.0^s J.T.

C/BARNETT 1969 I

APRIL 1970 APRIL 7.17 α $22^{\circ} 42^{\prime} 45^{\prime\prime}$ + $300^{\circ} 15^{\prime} 0^{\prime\prime}$ (1970)

P.A. $68^{\circ} 42^{\prime}$ Circle 28°

8 min motion = $44.5 = 1.727$ REV

$\frac{1}{2}$ = 1 min = 5.6 steps GUIDE STAR 24 min P
 3° N.

1 EXPOSURE 33 MIN LST. 15.10.1 \rightarrow 15.43.1 TRANSPARENT 7.8/10

2 .. 25 MIN LST. 16.22.1 - 16.47.1 TAP 33 - 31

APRIL 7.10.882 MATH DRIVE BIRTH EXPOSURE

APRIL 7.15591

EXPOSURE 2 $3^{\circ} 44^{\prime} 30^{\prime\prime}$ LST. APRIL 7.156

C/BARNETT 1969 I

APRIL 1970 APRIL 9.17 EXPOSURE 1 AT 9.11548

2 AT 9.15708

α $22^{\circ} 50^{\prime} 50^{\prime\prime}$ + $38^{\circ} 10^{\prime}$ (1970)

P.A. $66^{\circ} 12^{\prime}$ Circle $30^{\circ} 28^{\prime}$

16 min motion = $80.6 = 3.130$ REV

GUIDE STAR 24.5 P 1° SW of center

$22^{\circ} 26^{\prime} 30^{\prime\prime}$ + $39^{\circ} 40^{\prime}$

8 min motion = $40.3 = 1.565$ REV

$\frac{1}{2}$ = 1 min = 5.0 steps

ex ① Transparency $\pm 4/10$ LOW ALTITUDE & HAZE
 ex ② - $\pm 6/10$ MORE OR LESS OK OF HAZE
 BEGINNING TWILIGHT.

EXPOSURE ① 22 min L.S.T. 15.33.1 - 15.55.1
 ② 18 min L.S.T. 16.35.1 - 16.53.1

TEMP = 28-27 F = 16

ex ② APRIL 9.15.0 3^h. 36^m. 9^s U.T.

C/BARNETT 1969 I

1970 APRIL 10.17 (AP 10.08920)

α 22.55.6 + 39° 54.0 (1970)

P.A. 64° 45' CIRCLE = 32° 0'

16 min motion = 3.00 ROT in $\frac{1}{8}$ = 2 min = 9.6 steps

EXPOSURE 20 min Temp 30° F/6

LST 15. 0.1 - 15. 20.1 RLW Transparency 4-6/10

C/BARNETT 1969 I

APRIL 1970 AP 11.17 (AP 11.13426)

GOOD STAR

α 22.59.25 δ + 41° 33.7 (1970) 16 min

P.A. 64° 28' circle reading 32° 12' 10' N

16 min motion = 72° 10' = 2.80 ROT in 2 min $\frac{1}{8}$ = 9.0 steps

Observed 22 min L.S.T. 16.21.1 - 16.30.1 obs H. Masgala
Transparency $\sqrt{10}$ 8/10 Temp = 33° - 33° f = 15

C/Bennett 1969 I

1970 APRIL 14.17 (AP 14.12L400)

α 23.12.55 + 45° 58.0 (1970)

PA 59.57 Circle 37.13'

6 min motion = 62.6 = 2.43 min $\div \frac{1}{8} = 2 \text{ min} = 7.83 \text{ J/s}$

90.00 ftab 20 m f 1.25 N

Observed 22 min L.S.T. 16.5.1 - 16.27.1

Temp = 37° f = 14

Suddenly hinged up 2 min after starting. G.S. gradually difficult
sometimes invisible but occasionally bright Transparency $\sqrt{10}$