



BAA Asteroids & Remote Planets Section Meeting

An Introduction to Astrobiology

29 September, 2019

Peta Bosley



Why Astrobiology?

It's not observational.....

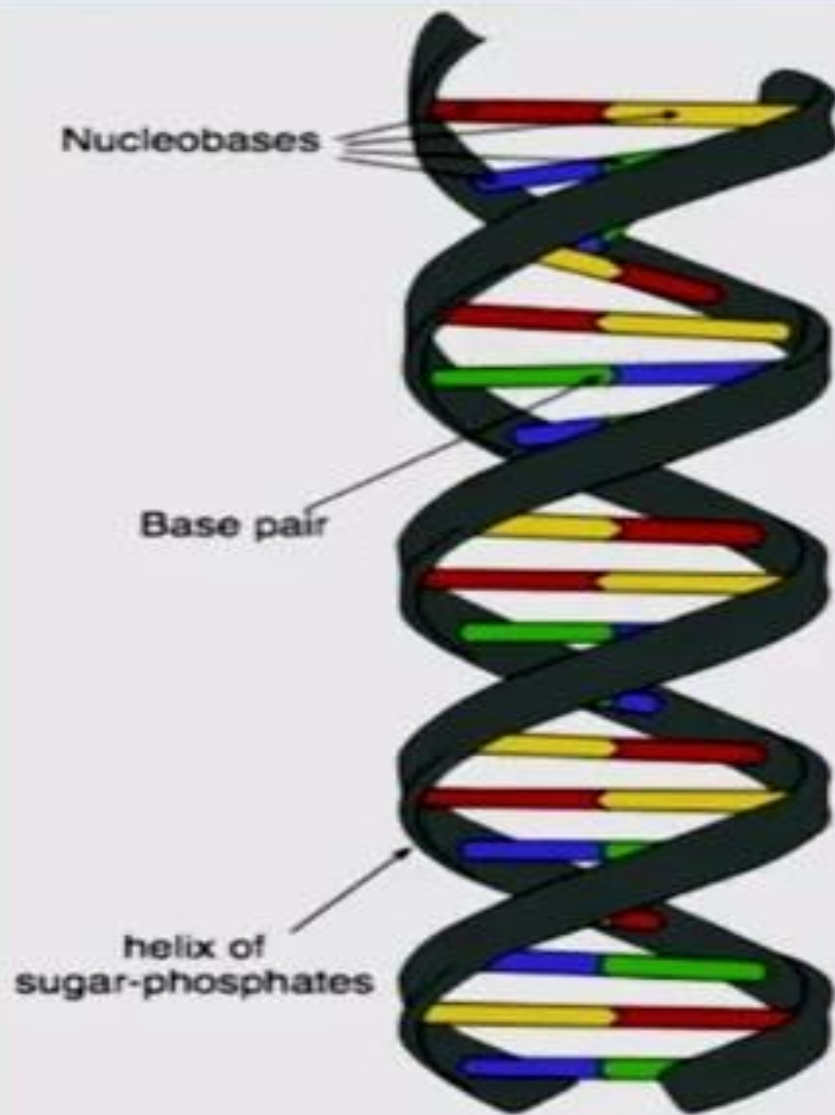


Interest is expanding

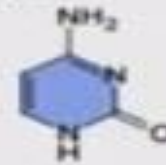


What is Life?

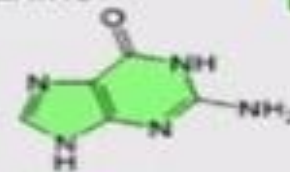
What was early Earth like?



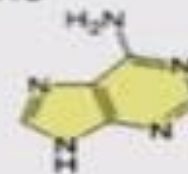
Cytosine



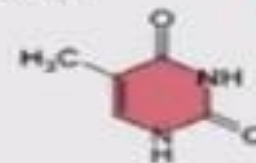
Guanine



Adenine



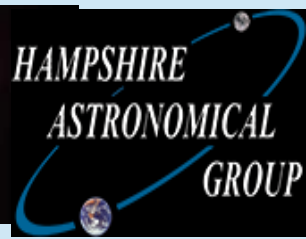
Thymine



Nucleobases
of DNA

DNA

Deoxyribonucleic acid





Early Earth – very different to our environment



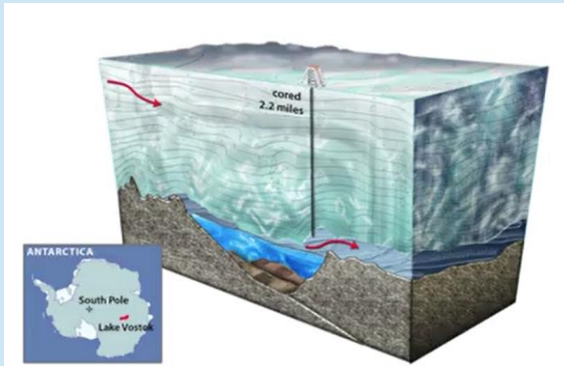
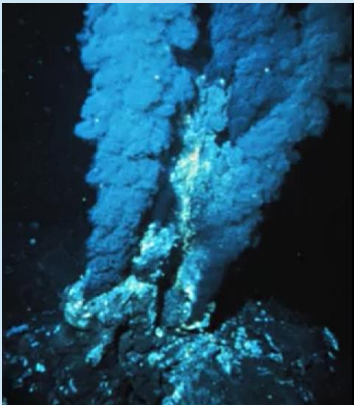
- .Constant volcanic eruptions spewed out methane
- .Fatal levels of UV radiation (no ozone layer then)
- .Dominant early life were unicellular anaerobic bacteria that had no need for oxygen
- .The problem arose when one life-form, the cyanobacteria, developed roughly 2.7 Ga (billion years) ago



- Up to 2.45 Ga / Oxygen levels @ 0.1% current atmosphere
- Great Oxygenation Event / Catastrophe started ~ 2.3 Ga
- Produced as waste product
- Up to 1.85 Ga / Combined in “mass rusting” - banded iron formations on ocean floor
- Carboniferous O₂ levels @ 35%



HAMPSHIRE
ASTRONOMICAL
GROUP





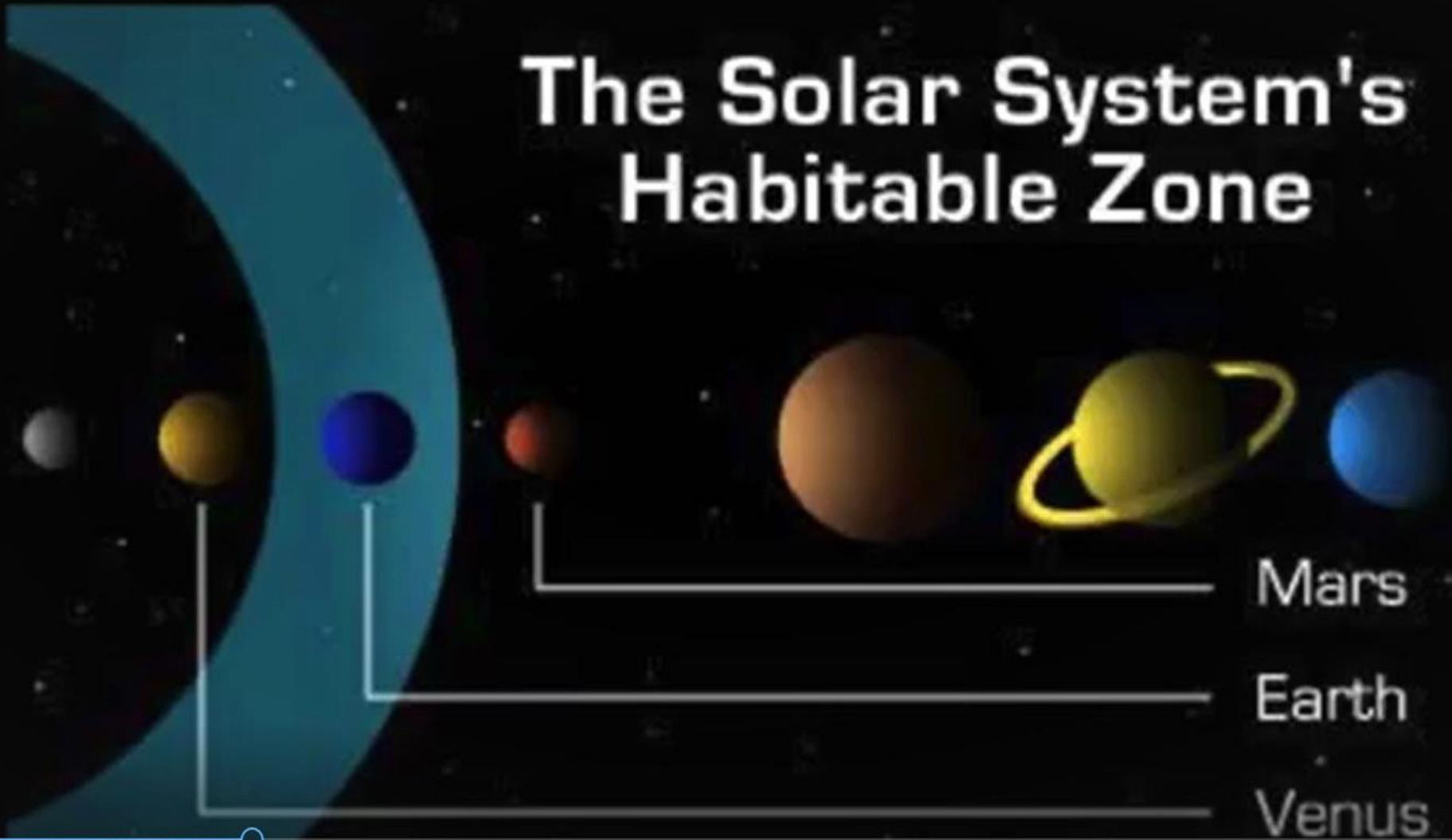


Habitable / “Goldilocks” Zone

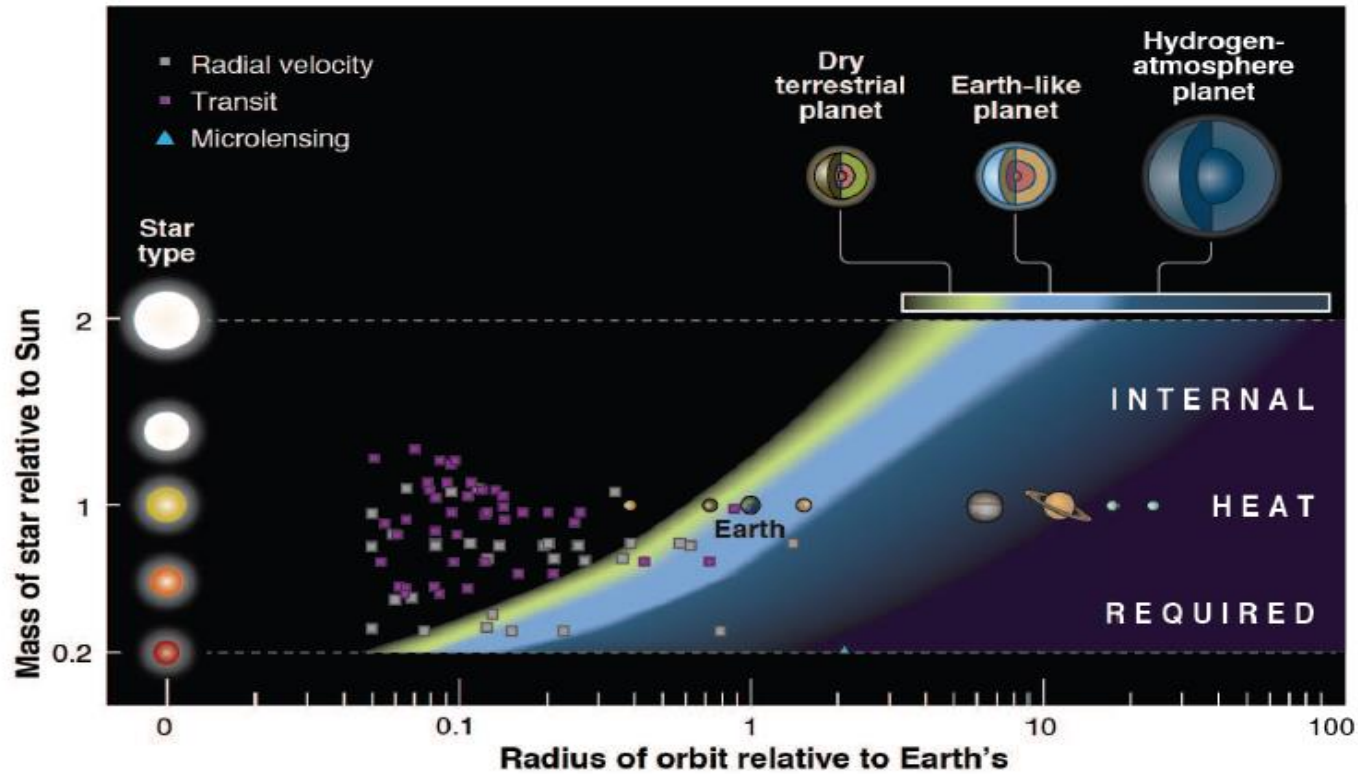
- Traditional thinking was in terms of distance from Primary
- Liquid Water on the Surface
- Distance from Primary varies with star type
- Hot stars have a more distant habitable zone



The Solar System's Habitable Zone

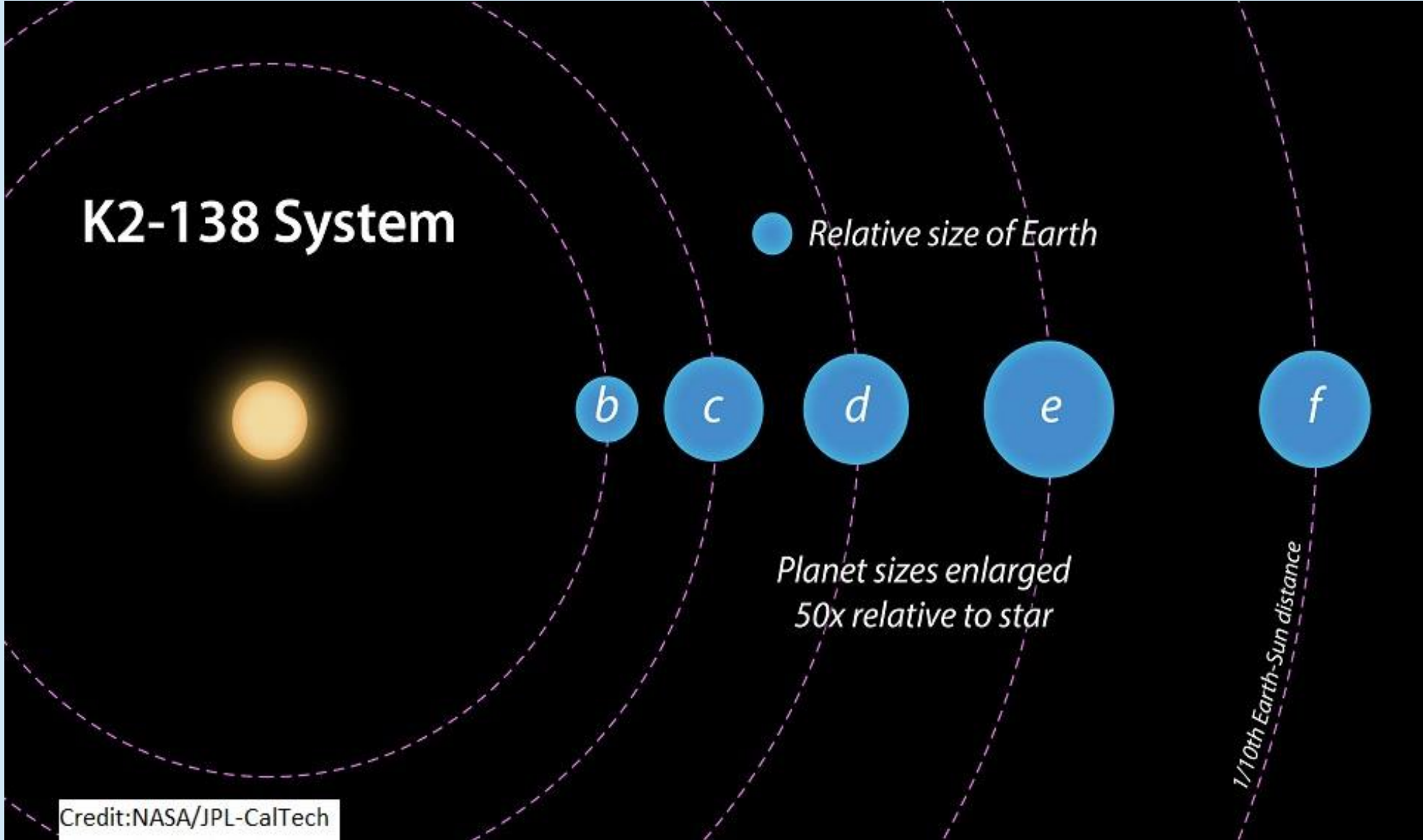


The Habitable Zone



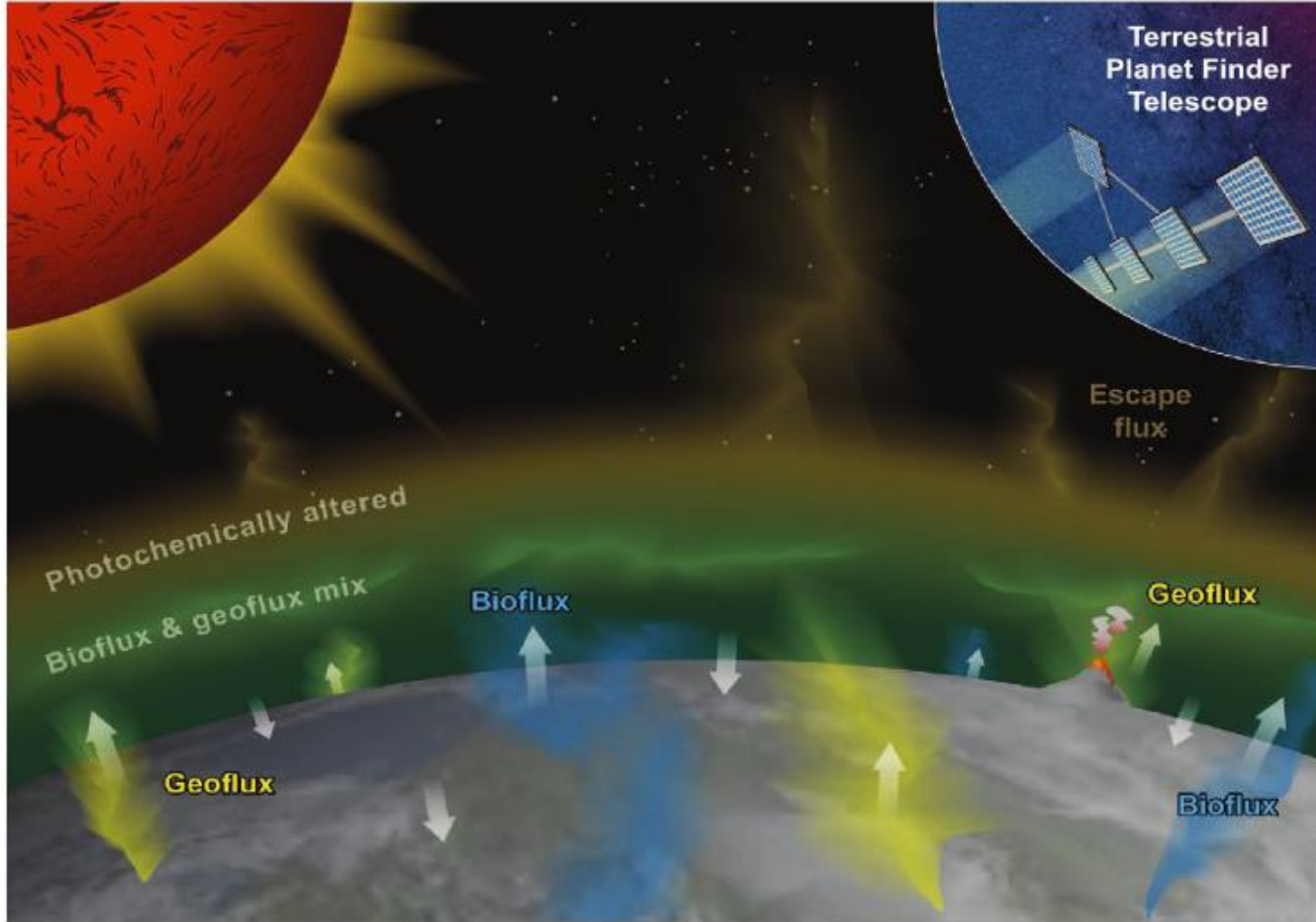
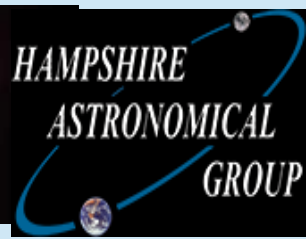
Seager, *Science* 2013

Inner edge: Zsom, Seager, de Wit, arXiv: 1304.3714





- .Biological processes (eg., photosynthesis) leave traces behind
- .These biosignatures can be remotely detected
- .Important biosignatures
 - IR reflections of vegetation
 - Gases in fluctuating states (eg., O₂/O₃ & CH₄)
 - Organic S compounds
- .Caution : non-biological factors can interfere
- .Consider environmental factors
- .Life could still exist if we don't find evidence in the atmosphere





Drake Equation – Number of Communicable Civilisations in our Galaxy

$$N = R * f_p * n_e * f_l * f_i * f_c * L$$

.R – Rate of star formation

.f_p – Fraction of stars with planetary systems

.n_e – Number of planets, per system, suitable for life

.f_l – Fraction of suitable planets, where life exists

.f_i – Fraction of life-bearing planets with intelligent life

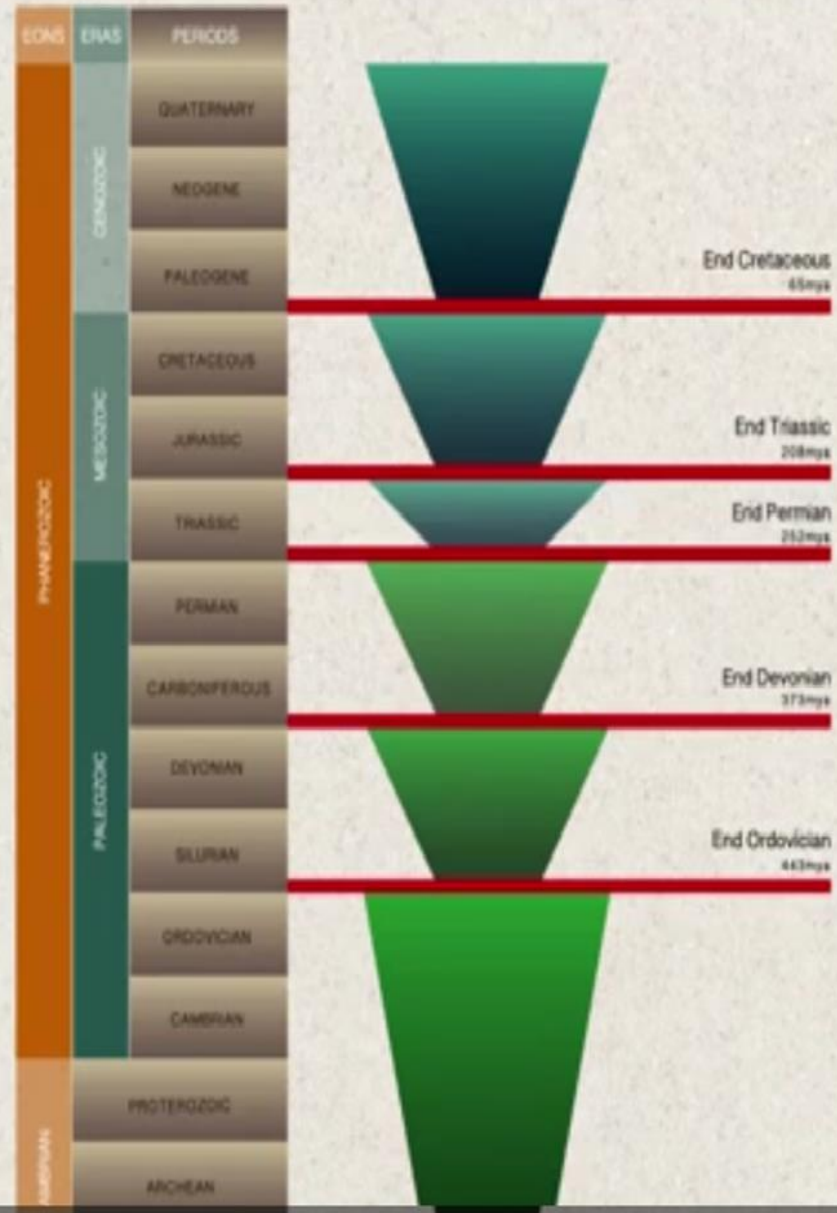
.f_c – Fraction of planets with intelligent life with technology and a capability for/interest in interstellar communication

.L – Average lifetime of a communicable civilisation



Extinction Events

- > 90% of the organisms to have lived on Earth so far are extinct
- There have been 5 major extinction events
- This is important to the Drake Equation
- If life dies out, does it ever get to the technological stage?
- Some scientists suggest that the 6th extinction event is in progress
- Will a technological civilisation survive?



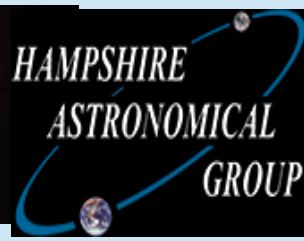


We appear stuck with a terracentric view of biosignature gases

“Nothing would be more tragic in the ... exploration of space than to encounter alien life and fail to recognize it” NRC report 2007



If you'd like to investigate the subject further



- Go back to Uni (probably not feasible for most of us!)

Or

- Study online



Massive

Open

Online

Courses



Online courses offered by many distinguished universities, under the umbrella of various organisations (FutureLearn, edX, Coursera are 3 of the biggest)



Imagining Other Earths

Tutor: Professor David Spergel

Dept of Astrophysics, Princeton

Delivery: Coursera



Astrobiology and the Search for Extraterrestrial Life

Tutor: Professor Charles Cockell

School of Physics & Astronomy

Delivery: Coursera



Astrobiology: Exploring Other Worlds

Tutor: Professor Chris Impey

School of Physics & Astronomy

Delivery: Coursera



Origins - Formation of the Universe, Solar System,
Earth and Life – Copenhagen / Coursera

Astrophysics: Exploring Exoplanets – Australian
National University (ANU) / edX

(Audit has time limits or £40 Certificate)