

2010 AUGUST

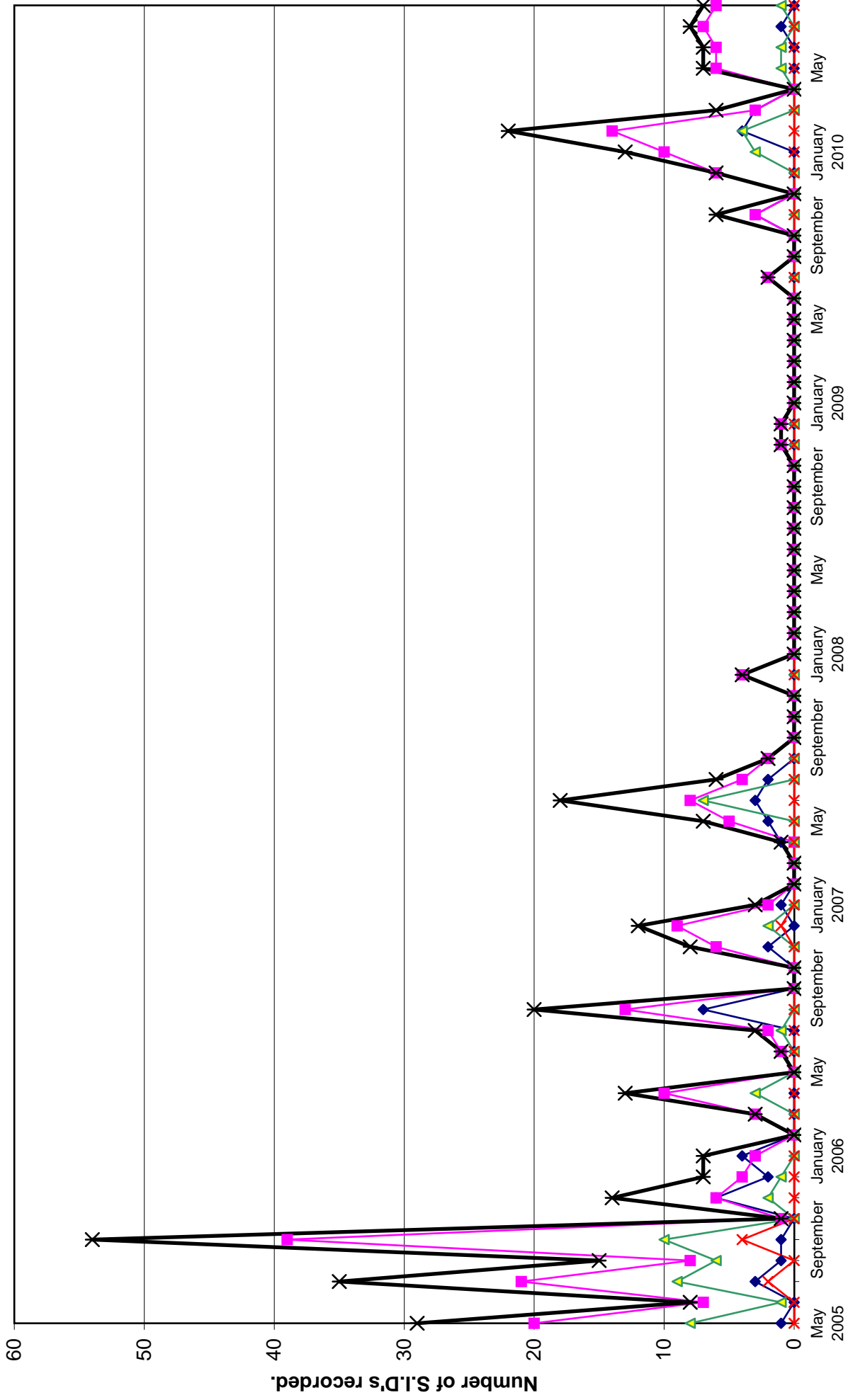
DAY	Ray class	Observers	John Cook (23.4/22.1kHz) Tuned radio frequency receiver, 0.58m frame aerial.	Roberto Battaiola (21.75kHz) Modified AAVSO receiver.	Nigel Curtis (23.4kHz) Gyrator receiver, shielded loop aerial.	Bob Middlefell (22.1kHz) Tuned radio frequency receiver, 0.5m frame aerial.	Mark Edwards (22.1kHz) Spectrum Lab / PC 2m loop aerial.
			START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)
1	C3.2	4	07:50 08:35 ? -				07:55 09:01 09:58 3
7	M1.0	4	18:07 18:20 18:44 2				18:10 18:23 18:45 2
14	C4.4	6	09:45 10:07 10:30 2				09:45 10:07 11:39 3
14	C1.6	1					18:06 18:11 18:18 1-
15	C5.4	5					18:25 18:£1 18:46 1
16	C1.4	4	16:37 16:41 16:50 1-				
18	C4.5	1					

DAY	Ray class	Observers	Colin Clements (23.4kHz) AAVSO receiver, 0.76m screened loop aerial.	Karen Holland (19.5kHz) Tuned radio frequency receiver, 0.58m frame aerial.	Mike King (20.9kHz) AAVSO receiver. loop aerial.	John Wardle (23.4kHz) Gyrator MKII receiver, 1m loop aerial.	Peter King (18.3kHz) Own designed receiver, 1.4m loop aerial.
			START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)
1	C3.2						
7	M1.0						
14	C4.4		09:39 10:14 11:14 3				07:55 08:25 09:35 3
14	C1.6						17:55 18:25 18:45 2+
15	C5.4		18:22 18:32 18:47 1				09:40 10:00 10:30 2+
16	C1.4						18:25 18:30 18:40 1-
18	C4.5						16:35 16:40 16:50 1-

DAY	Ray class	Observers	Paul Hyde (22.1kHz) Tuned radio frequency receiver, 0.96m frame aerial.	Gordon Flander (18.3kHz) PC sound card.	John Elliott (18.3kHz) Tuned radio frequency receiver, 0.5m frame aerial.	Martyn Kinder (18.2kHz) Tuned radio frequency receiver, 0.58m frame aerial.	Mark Hom (23.4kHz) Tuned radio frequency receiver, 0.58m frame aerial.
			START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)	START PEAK END (UT)
1	C3.2		07:41 08:22 08:44 2+				
7	M1.0		18:09 18:25 18:48 2				
14	C4.4		09:45 10:06 11:00 2+			09:37 10:04 ? -	
14	C1.6					18:25 18:32 18:47 1	
15	C5.4		18:25 18:31 18:42 1-			16:35 16:46 17:05 1+	
16	C1.4		16:38 16:45 17:02 1			04:45 ? ? -	
18	C4.5						

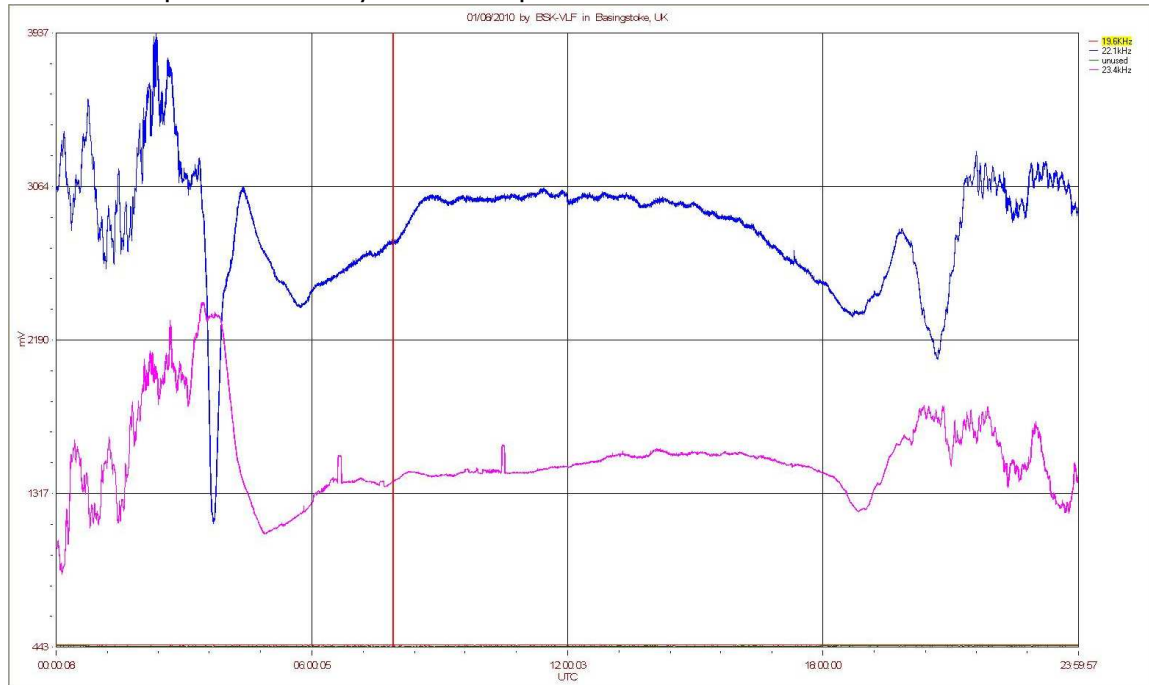
There were some unusual recordings this month, details shown further on. In addition to those listed above, there were numerous B-class flares in the first half of August, with further C-class events over night. Nothing is reported in the GOES record for the 19th to 25th, and the 29th. The most energetic event of the month was the M1 flare on the 7th.

VLF flare activity 2005/10.

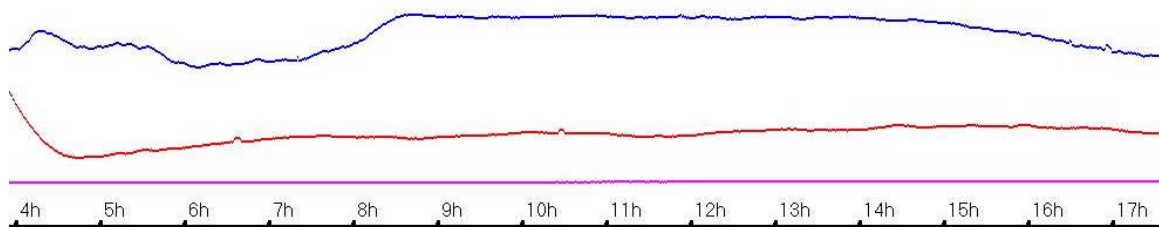


## August events.

The sun produced several rather unusual flares during August, the first being a C3.2 flare on the 1<sup>st</sup>. The GOES data records it as a slow flare. We are used to seeing sudden changes in ionisation from flares, giving rise to our Sudden Ionospheric Disturbances but this one produced a very Slow Ionospheric Disturbance.

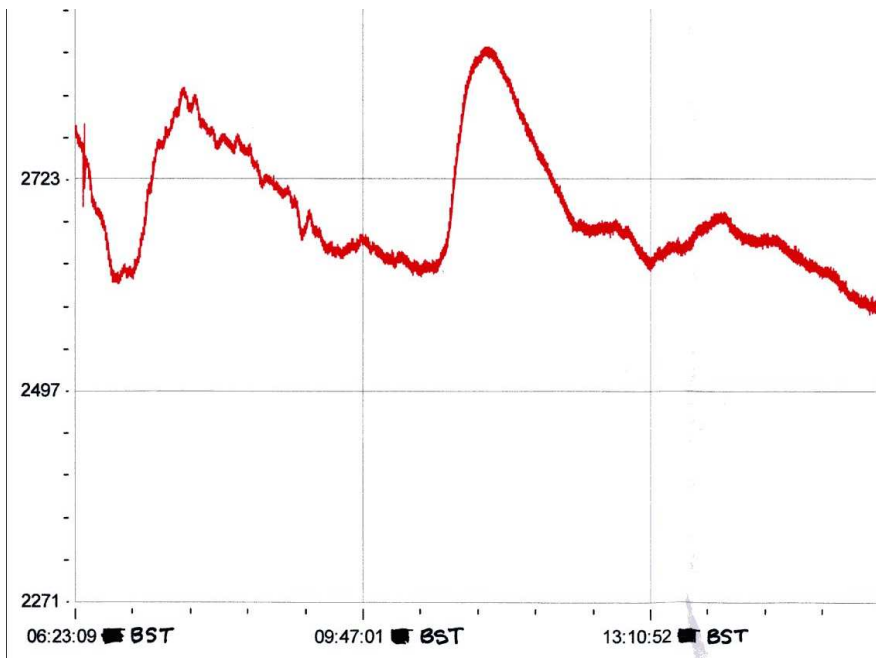


This chart by Paul Hyde shows its untypical recording. The blue trace at 22.1 kHz shows a gentle rise and decay starting at 07:41 UT. The red trace at 23.4 kHz shows much less effect.

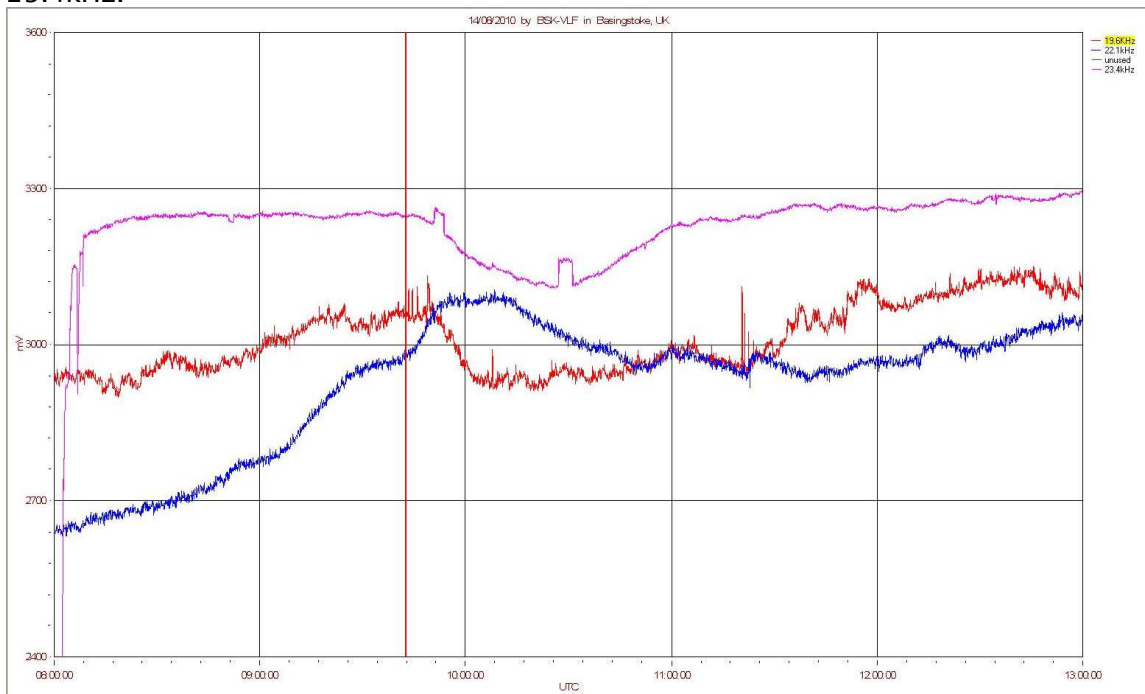


My recording shows similar results. Blue is again 22.1 kHz, showing a gentle rise followed by a nearly flat signal for the next 5 hours. 23.4 kHz (red) is not recorded at all. The space weather centre indicates that region 1092 was responsible for this long duration slow flare, which also produced a type IV radio sweep, an 890sfu Tenflare, and a full halo CME. A disappearing solar filament shortly before this event adds to the general confusion!

The C4.4 flare on the 14<sup>th</sup>. caused a similarly unusual SID.

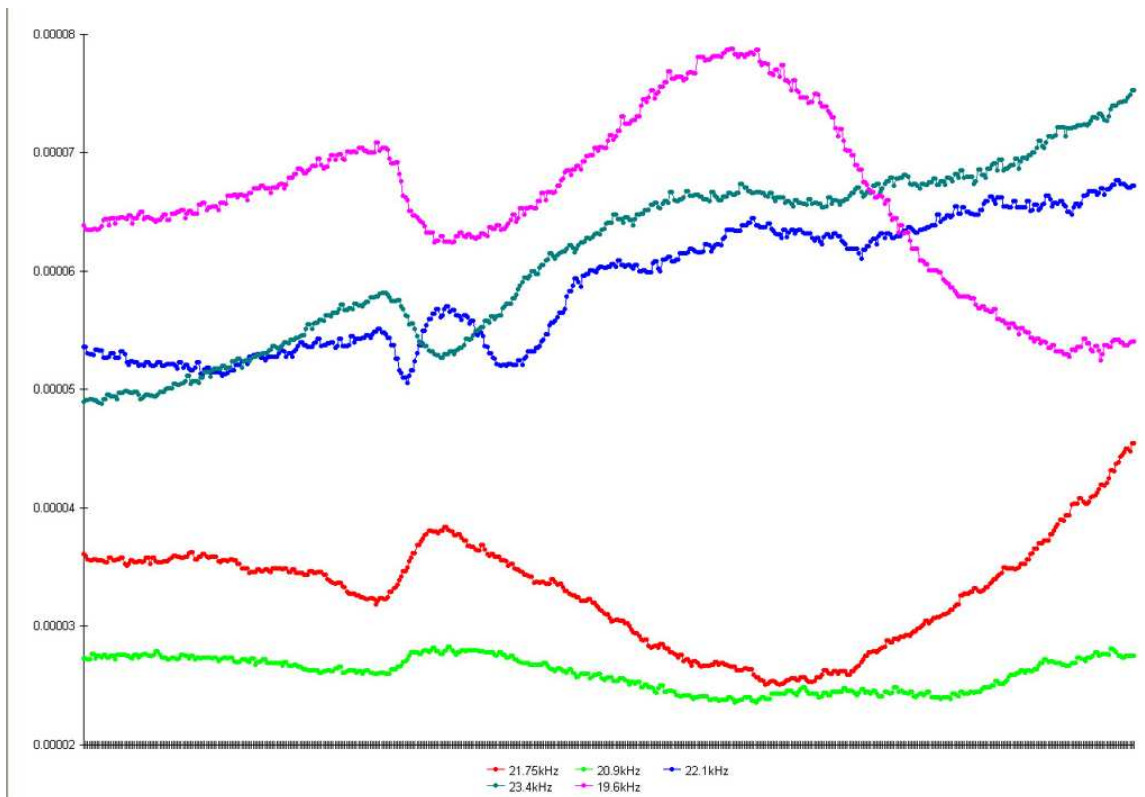


The chart from Colin Clements shows a nearly symmetrical response to this flare at 23.4kHz.



Paul Hyde recorded this chart, red is 19.6kHz, blue 22.1, and magenta is 23.4kHz. The space weather centre described this as a C4 slow flare, with a type II radio sweep, a proton event, and full halo CME.

The flare on the 15<sup>th</sup>. was also listed as a slow flare, but produced a more distinctive SID as shown in Mark Edwards recording:



Magenta is 19.6kHz, green is 20.9, Red is 21.75, blue is 22.1, and cyan is 23.4kHz.

### Smoke detectors.

Not another way to detect solar flares, but rather an important way to protect us and our radio equipment. In a very near miss, one of our observers was awakened by a smoke detector to find a thick grey fog enveloping the radio room. It seems that a fault had developed in an old oscilloscope which was gently steaming away close to catching fire. The smell of burning had saturated the contents and decoration of the room, requiring re-painting and a complete clean-out. Luckily that was the extent of the damage.

I recently had a free (courtesy of the local council) fire check of the house. My old smoke detector was replaced with a new one, and my VLF receiver system was closely inspected before being given the OK. If your council offers this service, then it is well worth taking up. It is also worth fitting thermal fuses to power supplies, so that any over-heating is quickly detected. Rather lose an observation than risk a fire!