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BAA Radio Astronomy Section.

Please send all reports and observations to jacook@jacook.plus.com Director Paul Hearn.

RADIO SKY NEWS

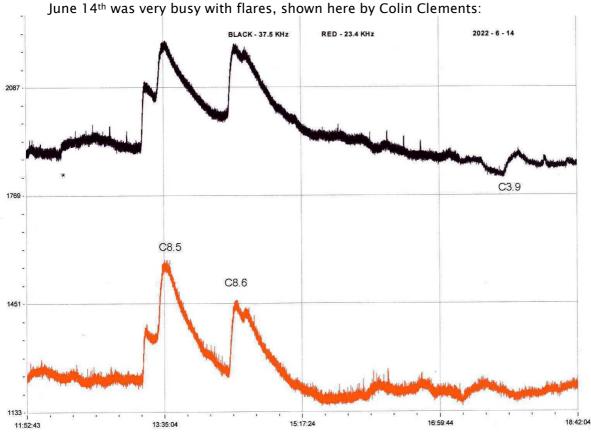
2022 JUNE.

VLF SID OBSERVATIONS.

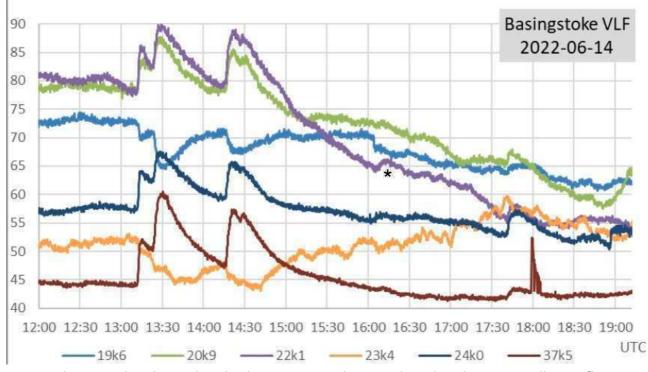
Sunspot counts in June remained high, but flares were generally much weaker than in May. The background X-ray flux also remained high, so the majority of smaller C-class flares were not recorded as SIDs. There are just two M-class flares listed in the SWPC bulletins; an M3.4 peaking at 04:07UT on the 13th was too early for us record, but we did get some recordings of the M1.2 flare on the 10th. This was quite a slow event, and so the VLF effect was not very SID-like. This recording By Steve Parkinson shows the result:



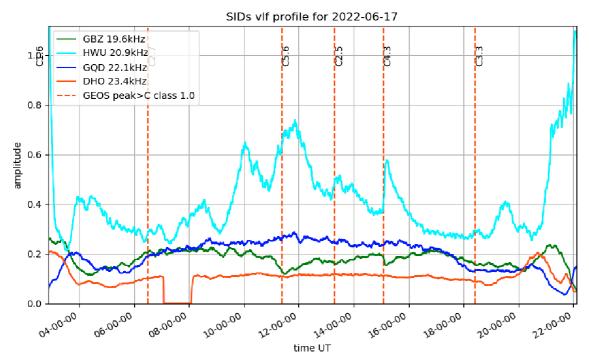
The pink trace is 19.6kHz, showing the M1.2 and C4.1 flares. Red is 23.4kHz, blue 20.9kHz and Green 22.1kHz. The timings are very tricky to determine, the tables showing a peak-time variation from 10:41to 11:02. The GOES X-ray flux is listed as start: 10:11, peak: 10:54, end: 11:14. Looking at the X-ray data, the flux still appears to be above C3 level at 12:00.



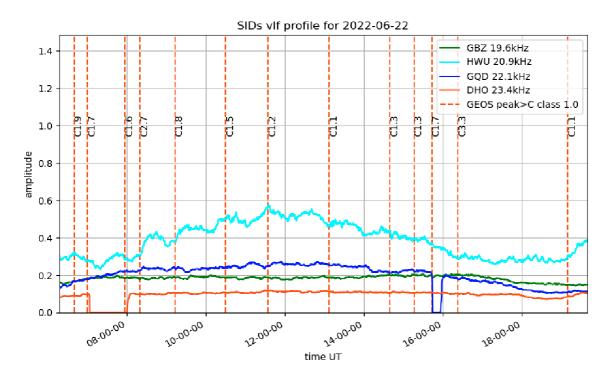
Both the C8.5 and C8.6 flares were double peaked, shown clearly at 23.4kHz and 37.5kHz. The 37.5kHz signal from Grindavik (top trace) also show evidence of the smaller unlisted flare at 12:38UT, indicated on the chart with a '*'.



This recording by Paul Hyde shows very similar SIDs, but also shows a small C1.3 flare at 16:13UT, again marked '*'. It is most clear in the 22.1kHz signal. The C3.9 flare at 17:50 is also visible on several of the signals.



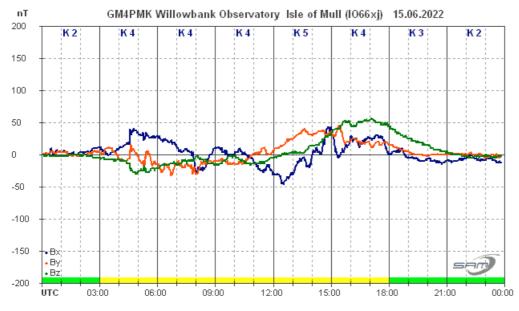
The 17th was also quite busy, shown here by Mark Prescott. 20.9kHz shows SIDs for most of the flares listed, while 23.4kHz has remained unaffected. There is also evidence of some non-solar noise on the 20.9kHz signal, with a strong rise in amplitude at 10:00 that has no X-ray counterpart. Mark's recording from the 22nd also shows some very noisy signals:



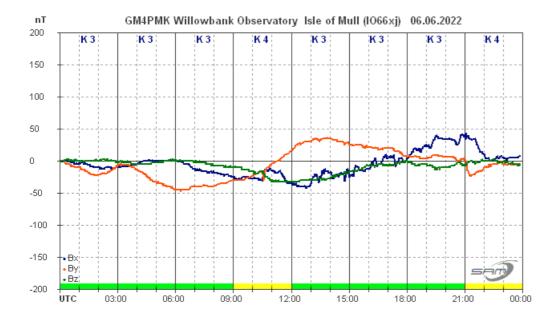
As noted at the start of this report, the high background X-ray flux made it tricky to record the smaller Cclass flares. 20.9kHz here shows the problem well, with a very noisy signal and plenty of hidden flares. 23.4kHz has also remained flat throughout the day.

MAGNETIC OBSERVATIONS.

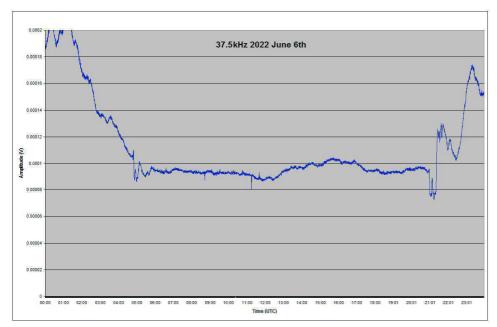
Most of the magnetic activity in June seems to have been from high speed winds, with only two CMEs that had Earth-directed components. The M3.4 flare early in the morning of the 13th did produce a CME, with fairly mild magnetic disturbances recorded on the 15th. Roger Blackwell's recording shows the disturbance during the day, along with what appears to be the arrival shock just after 04:30UT:



This was a very mild CME with only +/-50nT disturbance in the afternoon. It faded out in the evening and had very little effect into the following morning. A filament eruption on the 2nd also produced a CME that was geo-effective, arriving on the 6th.

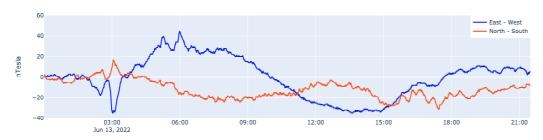


Roger Blackwell's recording shows that it had even less effect than the CME on the 15th. The sudden pulse in the Bx signal (blue) could mark its arrival at 10:30, the disturbance again fading out in the evening and with no disturbance in the morning of the 7th. The M1.2 flare recorded on the 10th does not appear to have produced a CME.

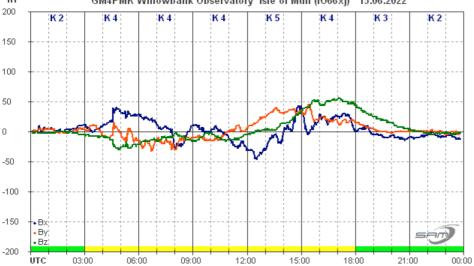


The CME could also be partly responsible for this disturbance to the 37.5kHz signal, recorded on June 6th by Mark Edwards. The early morning transient at 04:50 is well before the magnetic shock arrival, but does match magnetic disturbances recorded in other parts of the world. The pulse just before sunset at 21:00 is much larger, and matches the more widely recorded magnetic disturbances.

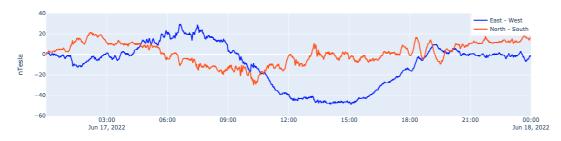
The remainder of the magnetic activity was from high speed solar winds, this recording by Nick Quinn shows a fairly strong change in wind speed around 3AM on the 13th: Steyning Magnetometer (50.8 North, 0.3 West)



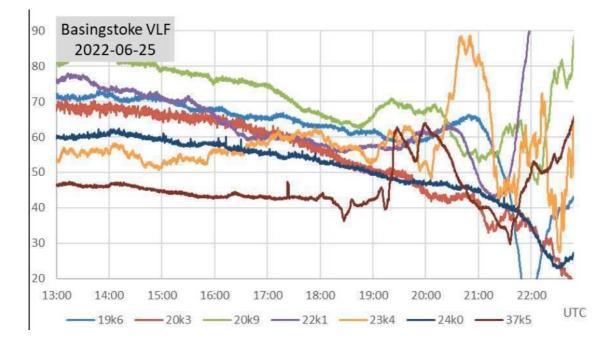
The initial transient is timed to match the M3.4 flare, but as this had very slow rise and fall times, it is unlikely to be an SFE. The following disturbance is rather mild at about +/- 40nT, but it did last for several days. The recording from the 15th is by Roger Blackwell, and the 17th by Nick Quinn: ^{nT} GM4PMK Willowbank Observatory Isle of Mull (1066xj) 15.06.2022



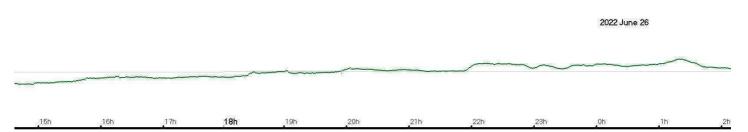
Steyning Magnetometer (50.8 North, 0.3 West)



There will be some variation in intensity at these different locations, the isle of Mull and Steyning near the south coast, but +/- 50nT seems to be the maximum disturbance recorded.



This recording from Paul Hyde shows the VLF signals on the 25th. The European signals are all moving into the sunset by 19:00, the longer trans-Atlantic path at 24kHz stable for a few hours longer. Most notable is the disturbance in the 37.5kHz signal after 18:30 and leading into its local sunset. This appears to be due to the magnetic disturbance from the high speed solar wind, shown in my own recording:

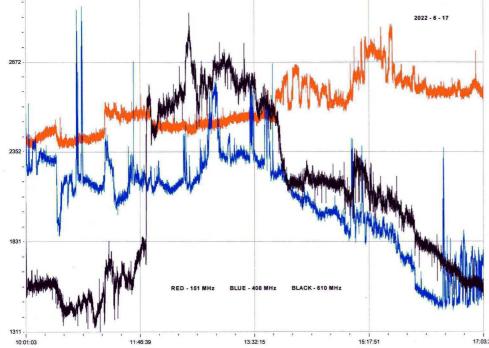


Luckily free from any local interference, a small disturbance can be seen from about 18:30 and lasting into the following morning. My single axis sensor tends to show smaller deviations compared to 2- and 3-axis sensors, with Colin Clements' and Roger Blackwell's recordings showing a stronger disturbance.

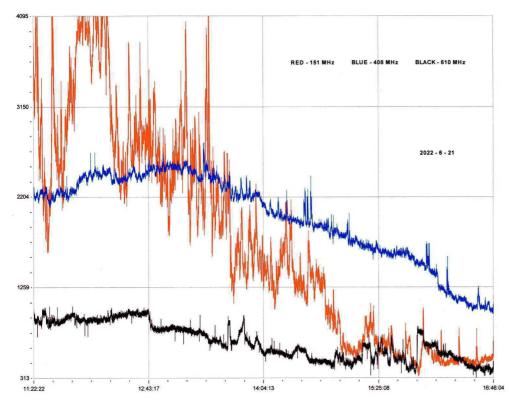
Magnetic observations received from Roger Blackwell, Colin Clements, Nick Quinn and John Cook.

SOLAR EMISSIONS.

Colin Clements did not record any significant VHF emissions following the M1.2 flare on the 10th, but did see some strong emissions from the flares on the 17th:

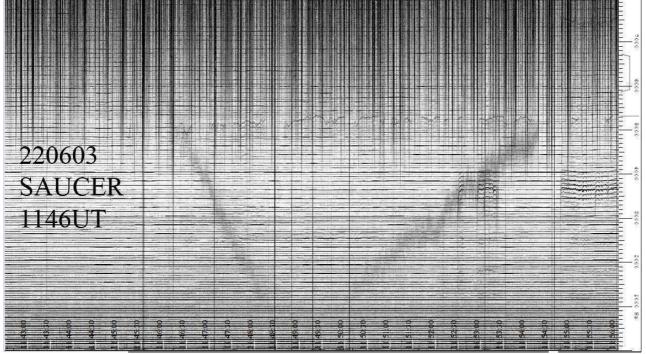


The 610MHz signal (black) shows a very strong signal during the first four flares. 408MHz (blue) is also quite strong, with a delayed signal at 151Mhz (red).

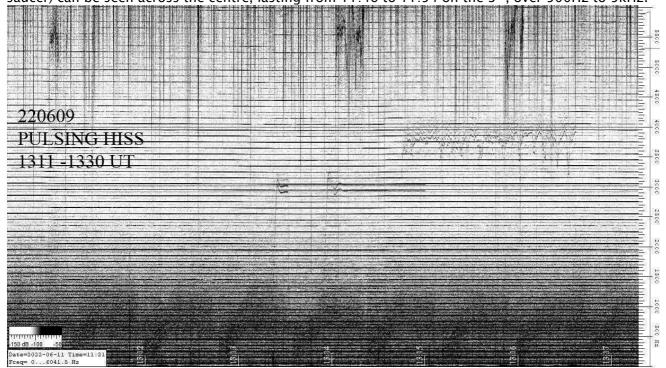


Colin's recording from the 21st is unusual, showing a very strong signal at 151MHz, with very little on the others. It does not match with our flare timings in the afternoon so may well be some interference, although it has not been seen before.

Colin Briden has been recording again at VLF, seeing the usual chirps and hooks previously reported. In June he also recorded some less usual effects that are apparently associated with Aurora.



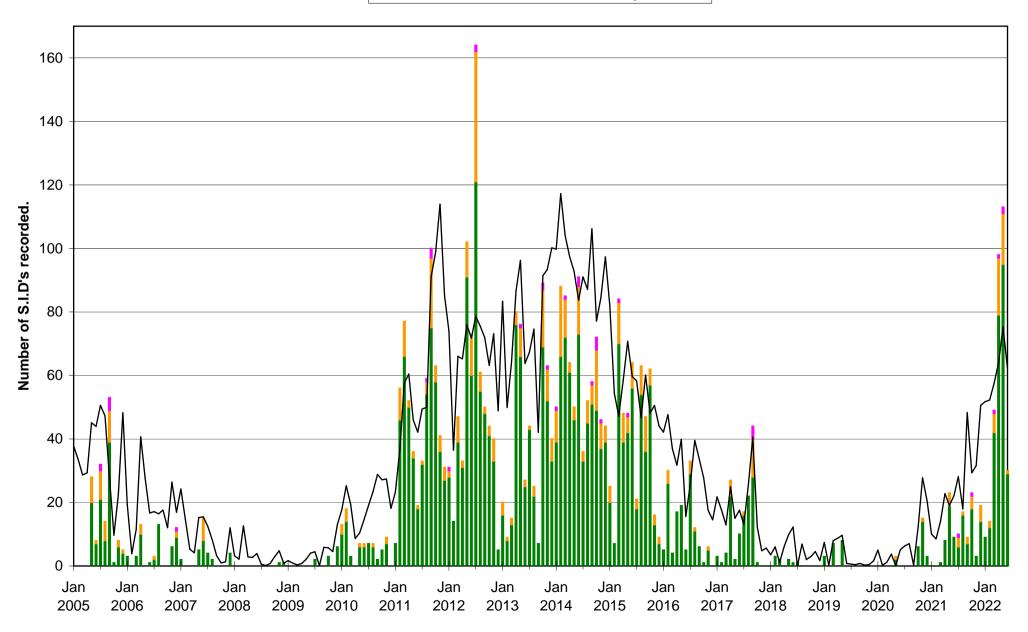
The time and frequency markings are unfortunately not very clear in the picture. Frequency runs vertically, 0Hz at the bottom and 8kHz at the top. Time runs horizontally. A darker V-shaped pattern (known as a saucer) can be seen across the centre, lasting from 11:46 to 11:54 on the 3rd, over 500Hz to 5kHz.



This recording covers 0Hz to 6kHz, and shows a pulsing hiss between 500Hz and 1.5kHz. The period is about 50 seconds. This was seen for about 30 minutes on the 9th. There is evidence of extra noise at higher frequencies on both recordings. This was observed to be from a neighbour's lawn mower, and trains arriving and leaving a nearby railway station.

VLF flare activity 2005/22

C – M – X – Relative sunspot number



BAA Radio Astronomy Section.

2022 JUNE.

	SS	ers	John C	ook (23	.4kHz/22.1	kHz)		Roberto Bat	taiola	Paul I	Hyde (22	.1kHz/2 <i>4k</i>	Hz)	Mark Edv	vards (24	4.0/ <i>19.6</i> / 3	7.5kHz)	Co	lin Cleme	nts (37.5kH	z)
	r class	Observers	Tuned I	adio free	quency rec	eiver,	Maa	dified AAVS0		Spectru	ım Lab /	PC 1.5m f	rame	Crosstrum	alah / D			Tuned F	Radio Fre	quency rece	eivers,
	Xray	sqC	0	.58m fra	me aerial.		IVIOC		J receiver.	-	aer	ial.		Spectrum	I Lad / P	°C 2m loop	o aenai.			ed loop aer	
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9 9	C1.7 C1.9	2								10:08	10:21	10:53	2	10:14	10:27	10:59	2				
10	C3.7	2								10.00	10.21	10.00	-	05:39	05:53	06:11	_ 1+				
10	C4.1	7	08:25	08:41	09:16	2+				08:23	08:42	09:36	2+	08:30	08:40	09:26	2+	08:32	08:50	09:39	2+
10	M1.2	7	10:15	10:45	12:01	3				10:14	10:42	12:00	3	10:16	10:54	12:10	3	09:39	11:02	12:07	3+
10	C2.3	2								14:24	14:35	14:47	1	14:29	14:36	15:11	2				
11	C1.2	1												14:11	14:10	14:17	1-				
13	C8.5	1								00.50	10.00	40.00	0	21:23	21:26	21:29	1-	10:02	10.14	10.51	2.
14 14	? ?	5 1								09:59	10:09	10:32	2	10:01	10:12	10:37	2	10:03	10:14	10:51	2+
14	?	1												12:32	12:39	13:00	1+				
14	?	6	13:13	13:15	?	-				13:10	13:17	?	-	13:13	13:16	?	-	13:19	13:21	13:28	1-
14	C8.5	8	13:24	13:30	?	-				13:23	13:31	14:11	2+	13:22	13:29	?	-	13:28	13:35	14:22	2+
14	C8.6	8	14:17	14:20	15:28	2+				14:16	14:24	?	-	14:17	14:23	14:59	2	14:22	14:28	15:31	2+
14	?	2								14:27	14:30	15:08	2	14:28	14:29	15:19	2+				
14	C1.3	1								16:06	16:13	16:28	1								
14	C3.9	3								17:41	17:47	18:30	2+	17:42	17:50	18:15	2				
16	C4.5	8	13:43	13:44	14:14	1+				13:41	13:46	14:32	2+	13:42	13:46	14:07	1	13:48	13:51	14:04	1-
17	C5.6	6	10:59	11:21	12:13	2+								10:58	11:29	?	-	10:36	11:29	11:36	2+
17 17	? C2.5	2 2								13:06	13:16	13:36	1+	11:43 13:10	11:50 13:26	12:18 13:57	2 2+	11:48	11:53	12:40	2+
17	C2.5 C4.3	2 8	15:01	15:06	15:21	1				14:58	15:07	16:04	2+	15:02	15:20	15:34	2+ 1+	15:05	15:11	15:30	1
17	C3.3	1	10.01	10.00	10.21	ļ				14.00	10.07	10.04	21	18:25	18:27	18:46	1	10.00	10.11	10.00	'
18	C1.8	2												11:38	11:46	12:11	2				
18	?	1												11:53	11:57	12:06	1-				
18	C1.8	1												13:30	13:35	13:57	1+				
18	C1.4	1												17:41	17:48	17:57	1-				
19	C4.0	1												20:00	20:09	20:36	2				
20	C5.7	2												06:17	06:20	06:25	1-				
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22	C3.3	1											-	16:23	16:25	16:47	1				
23	C4.0	1												-	-						
23	C3.6	1												20:07	20:29	20:49	2				
24	C1.5	5	08:55	09:00	09:15	1				08:54	09:06	09:37	2	08:56	09:06	09:37	2	08:58	09:10	09:40	2
25	C1.2	1												14:01	14:03	14:16	1-				

BAA Radio Astronomy Section.

2022 JUNE.

	class		Stev	e Parkin	son (Vario	us)	Andrew T	homas (20.9/22.1/19).6kHz)	Ph	il Rourke	e (23.4kHz)	Mark Prescott (20.9kHz)				Christopher Bailey					
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11	C1.2																						
13 14	C8.5	5					09:41	10:08	10:38	2+				09:55	10:16	10:45	2+						
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14 14	C8.6	5	14:17	14:22	15:15	2+	14:17	14:21	15:57	3				14:16	14:25	14:49	2	14:00	14:22	14:50	2+		
14	C1.3	3																					
14	C3.9)					17:40	17:46	18:03	1													
16	C4.5		13:42	13:47	14:15	2	13:43	13:47	14:18	2				13:43	13:51		2	13:42	13:45	14:15	2		
17 17	C5.6 ?	5					10:56	11:24	12:32	3				10:59	11:53	12:35	3	10:45	11:25	12:12	3		
17	؛ C2.5	5																					
17	C4.3		15:00	15:08	15:34	2	15:01	15:07	16:02	2+				15:03	15:14	16:25	2+	15:00	15:05	15:25	1		
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18	C1.8	3												11:43	11:56	12:11	1+						
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19	C4.0																						
20	C5.7	,					06:05	06:21	06:44	2													
20 20	? ?													15:24	15:33	15:50	1+						
20	: C4.5	5												16:09	16:16	16:31	1						
21	C5.6	5	16:15	16:24	16:42	1+	16:18	16:22	16:46	1+				16:17	16:29		2						
22	C2.7																						
22 23	C3.3 C4.0						09:38	11:26	12:08	3+													
23 23	C4.0						09.30	11.20	12.00	J+													
24	C1.5						08:54	09:11	09:45	2+													
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