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BAA Radio Astronomy Section.

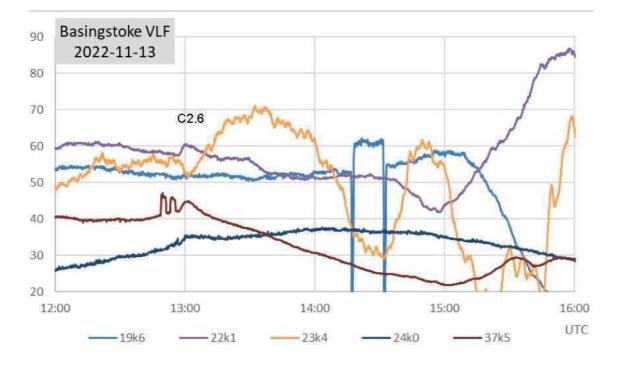
Please send all reports and observations to jacook@jacook.plus.com Director Paul Hearn.

# **RADIO SKY NEWS**

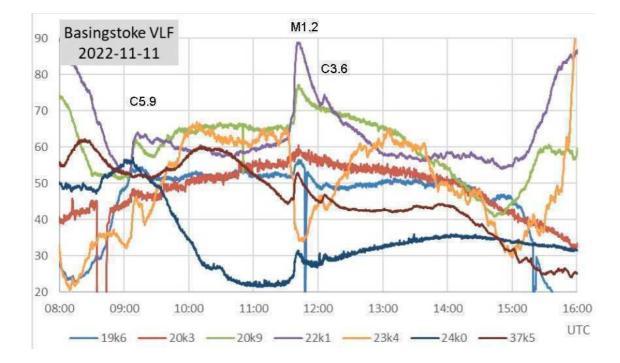
# 2022 NOVEMBER.

## VLF SID OBSERVATIONS.

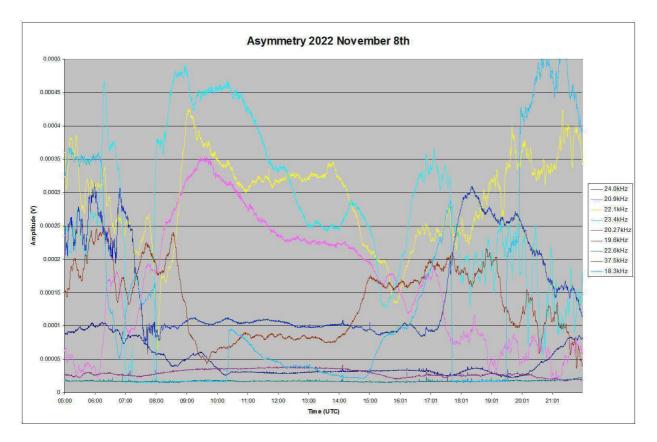
Solar activity has been much lower than in October, with a total of 35 SIDs recorded. Signals have also been very unstable making the smaller flares difficult to detect in the noise. 23.4kHz seems to have been particularly badly affected, while the more northerly / southerly paths have been less affected. Paul Hyde's recording from the 13<sup>th</sup> shows the problem well:



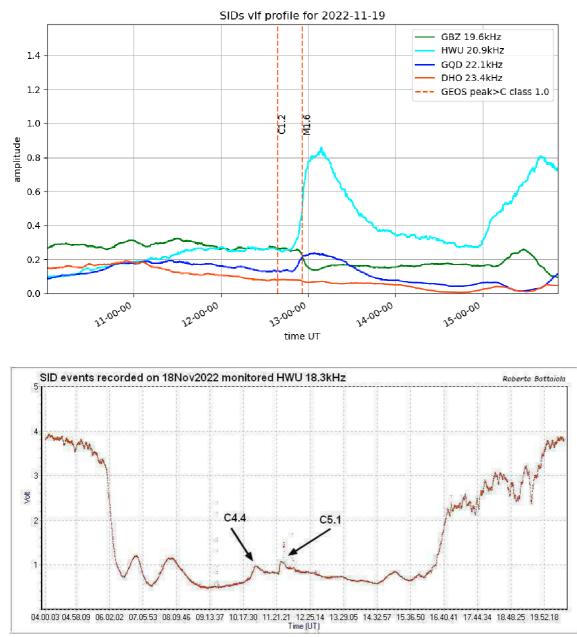
The C2.6 flare stands out well at 22.1 and 37.5kHz, despite some glitches from Grindavik. 23.4kHz shows a good SID, but it is not that obvious on the very unstable signal. It also shows the very early sunset features on the easterly path. Activity on the 11<sup>th</sup> was rather stronger, and so the SIDs are a little easier to see on Paul's recording on the next page. 19.6kHz from Skelton, just a few miles away from the Anthorn site, has not responded as well on either chart, and 23.4kHz shows much the same instability. The M1.2 flare does stand out on most signals, but not 20.3kHz, from Italy. The much smaller C3.6 flare during the M1.2 decay is also well recorded. Both flares were from the same active region.



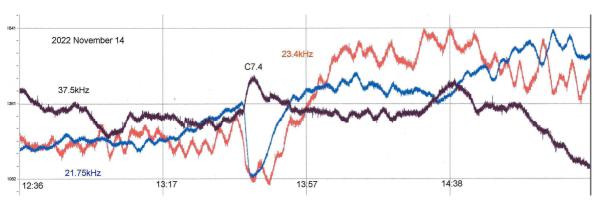
Mark Edwards noted some strong asymmetry in the diurnal curves, particularly on the 8th:



This shows particularly at 23.4kHz (light blue), 22.1kHz (yellow), and 20.9kHz (pink), with an inverted asymmetry at 19.6kHz (brown). My own 23.4kHz recordings also show this asymmetry at 23.4kHz on the 8<sup>th</sup>, and from the 19<sup>th</sup> to the 21<sup>st</sup>. These are mostly days of very low solar activity, although there was an M1.6 in the early afternoon of the 19<sup>th</sup>. This flare shows well at 20.9kHz in the recording by Mark Prescott, although much less asymmetry in the diurnal curve.

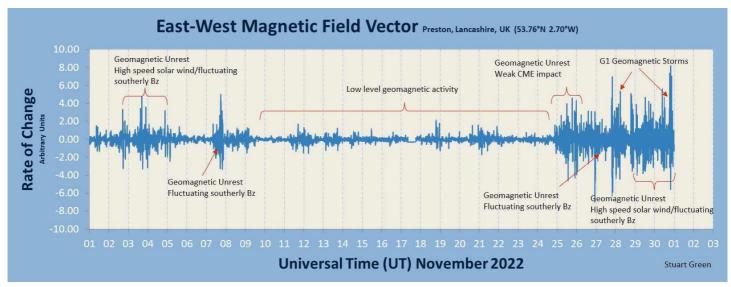


Two much smaller flares have produced very clear SIDs at 18.3kHz in this recording by Roberto Battaiola on the 18<sup>th</sup>.



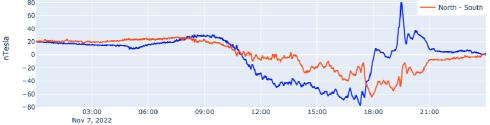
Colin Clements' recording from the 14<sup>th</sup> shows the C7.4 flare, once again on very unstable signals. 23.4kHz also again seems to be the worst affected.

#### MAGNETIC OBSERVATIONS.



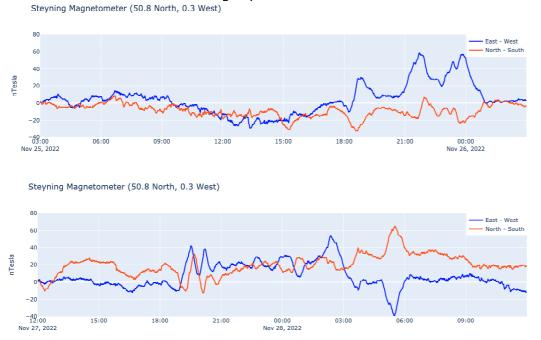
Stuart Green's monthly summary shows a very quiet magnetosphere through most of November, starting with some weak solar wind effects and ending with an extended active period during the last week.



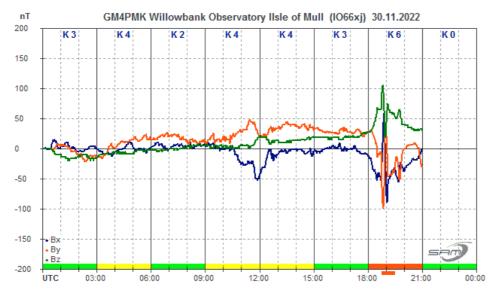


The most active magnetic disturbance was recorded in the afternoon and evening of the 7<sup>th</sup>, shown in the top recording by Mark Edwards, and the lower chart by Nick Quinn. The sensitivities of these two sensors have opposite polarities, but the dominant features are clear in both, with a strong transient at 19:30UT. The source of the disturbance appears to be a brief period of fast solar wind, lasting just 12 hours.

There were a number of CMEs associated with the stronger flares, but they were all from active regions near the solar limb, and so had limited effect on our magnetic field. Stuart Green's summary chart indicates a weak CME impact around the  $25^{th}/26^{th}$ , although its source is not clear. Most of this disturbance seems to have been form a coronal hole high speed wind.



Nick Quinn's recordings from the 25<sup>th</sup> and 27<sup>th</sup>/28<sup>th</sup> show some of this activity. There is no clear sign of a CME impact, indicating that it may have been just a glancing blow.

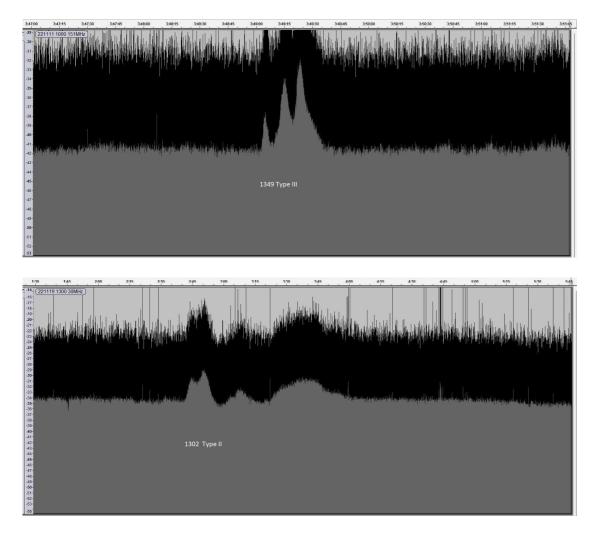


Roger Blackwell's recording from the 30<sup>th</sup> shows further high speed wind effects building up in the evening of the 30<sup>th</sup>, although his recording terminated early for some reason.

Stuart Green's summary chart shows activity based on the rate of change in the magnetic field, while the other charts all show the field intensity over time, accounting for the apparent weaker activity at the end of November that appears to be stronger in the summary chart.

Magnetic observations received from Roger Blackwell, Colin Clements, Stuart Green, Nick Quinn and John Cook.

### SOLAR EMISSIONS.

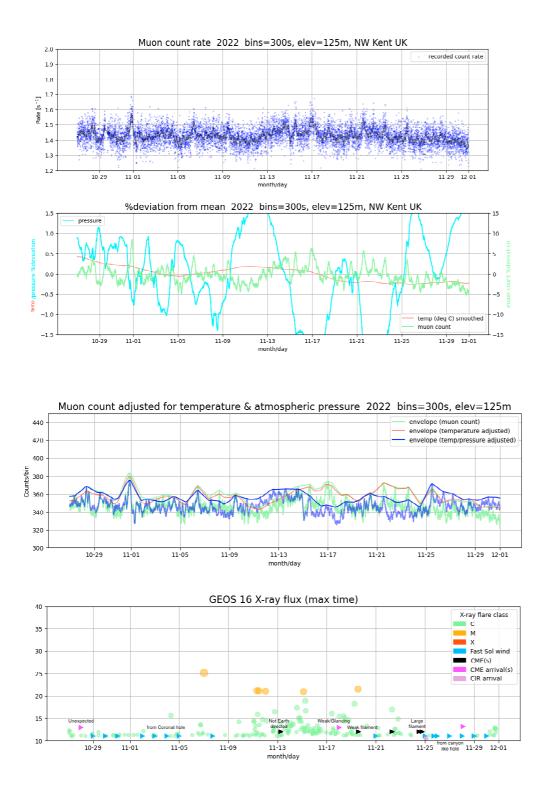


Colin Briden made two solar emission recordings, the first a type III burst at 13:49UT on the 11<sup>th</sup> at 151MHz. Shown in the top image, this lasts for just 30 seconds, with an amplitude of about 10dB. It does not match any of our recorded flares, but does match the timing of a C1.0 flare listed in the SWPC satellite data. The second recording shows a type II burst at 13:02UT on the 19<sup>th</sup> at 30MHz. This lasts for about one minute, with a 5dB amplitude. This one matches the M1.6 flare shown in our SID recordings.

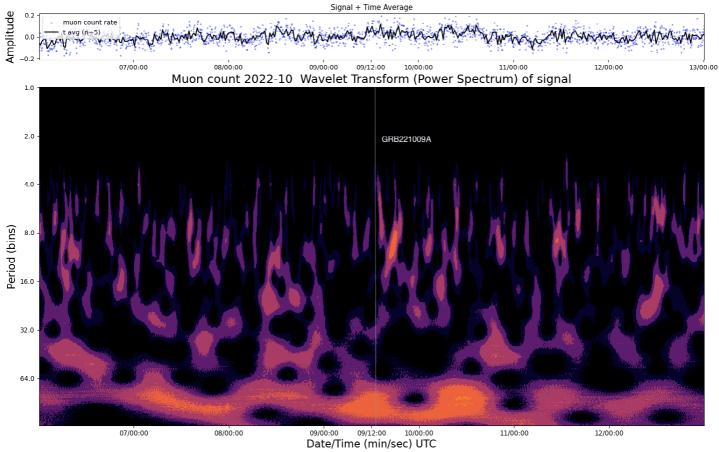
Colin has noticed that the 28-30MHz band is becoming more noisy with interference as the increased solar activity allows better HF propagation. There is also the problem of the sun being much lower in the sky at this time of year. These recordings are currently made using a simple dipole aerial that can easily be trimmed to suit different frequencies. Colin Clements is unable to make VHF recordings while the sun remains below his aerial's horizon.

#### MUON OBSERVATIONS.

Mark Prescott has continued his muon recordings, shown in the charts on the next page. There does not appear to be any stand-out events this time, but it does seem that the general muon count is lower during periods of faster solar wind.



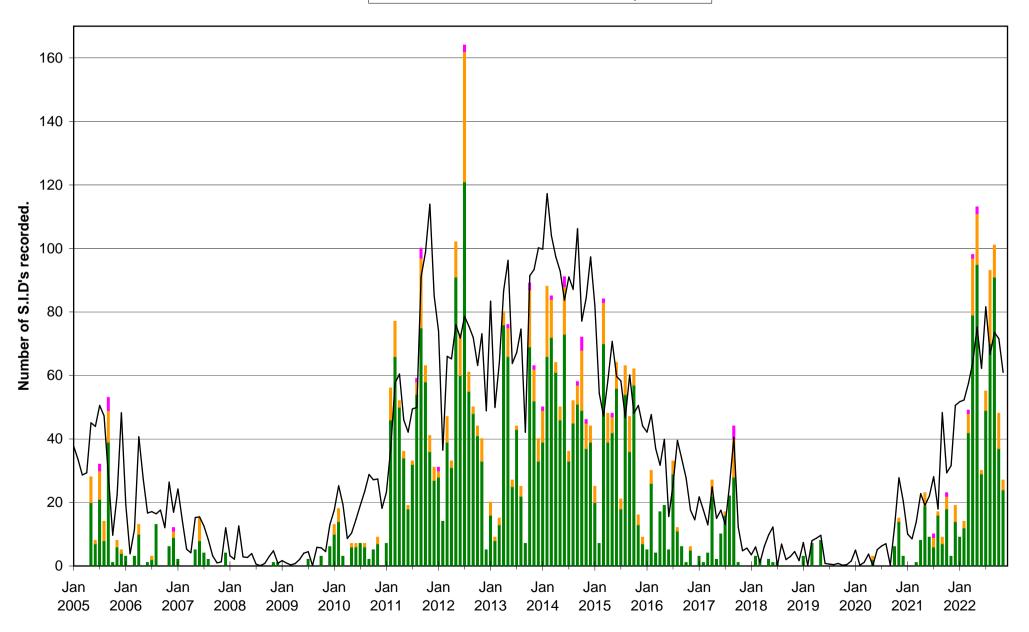
Mark has also been analysing his data during the GRB recorded on October 9<sup>th</sup>. The chart on the next page shows a wavelet transform, indicating a slightly higher muon count on the 9<sup>th</sup> and 10<sup>th</sup>. The GRB timing is marked with a vertical line near the centre of the chart. The vertical axis represents the wavelet period in recording bins, with colour indicating the intensity. There is also a higher count on the 10<sup>th</sup>. There is a general trend of increased intensity during daylight hours compared with night. This represents just a single observation, and so the link to the GRB is probably weak. It is however a useful recording that can be used to compare with other active and quiet periods over the next year or so as we learn to interpret the data.



There will be a solar section Zoom webinar on February 18th, including a radio astronomy contribution from Paul Hearn. Full details can be found on the BAA website.

### VLF flare activity 2005/22

C – M – X – Relative sunspot number



#### BAA Radio Astronomy Section.

#### 2022 NOVEMBER.

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BAA Radio Astronomy Section.

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:	Xray class		Chris	Bailey		Andrew Lutley (23.4kHz)	Peter Meadows (23.4kHz)		
;	Xray cla						Peter Meadows (23.4KHZ)		
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22 C	24.2	13:35	13:59	14:35	2+				

BAA Radio Astronomy Section.

BARTELS DIAGRAM

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ROTATION	KEY:		DISTU	IRBED.			ACTIVE	E		SFE			B, C, M,	X = FLA	RE MAG	NITUDE		Sy	nodic ro (carring		art						
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2545	F 2020 M 2	March 3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	2229 28
2546	29	30	31	2020 Ap 1	oril 2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24
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2551	- 11 F	2234 12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	2020 Se 1	ptembe 2	r 3	4	5	6
2552	7 F	2235 8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	2020 Oc 1	ctober 2	3
2553	4 F	2236 5	6	7	8	9	10	11	12	13	14	15	16 CC	17	18	19	20	21	22	23	24	25	26	27 C	28	29 BCCC	30
2554	31 F	2237 1 B	2020 No 2	ovember 3	4 B	5 CBCC	6 CBC	7 B	8	9	10 C	11 C	12	13	14	15	16	17	18	19	20	21	22 CC	23	24	25	26 C
2555	27 F	28 C	2238 29 CM	30	2020 D 1	ecember 2	3	4	5	6 C	7 C	8	9	10	11	12	13	14 C	15	16	17	18	19	20	21	22	23
2556	24 F	25	2239 26	27	28	29	30	31	2021 Ja 1	inuary 2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19
2557	20 F	21	22	2240 23	24	25	26	27	28	29	30	31	2021 Fe	2	3	4	5	6	7	8	9	10	11	12	13	14	15
2558	16 F	17	18	2241 19 2242	20	21	22	23	24	25	26	27	28	2021 M 1	2	3	4	5 2021 Ap	6 ril	7	8	9 C	10	11	12	13	14
2559	15 F B	16	17	18	19	20	21	22	23	24	25 B	26	27	28	29	30	31	1	2	3	4 2021 Ma	5	6	7	8	9	10
2560	11 F	12	13	14 2244	15	16	17 B	18	19	20 C	21	22 CCCC	23	24	25	26	27	28	29	30	1	2	3	4	5 2021 Ju	6 ne	7 M
2561	8 F CC	9 CC	10	11	12 C 2245	13	14	15	16	17	18	19	20	21 C	22 CCMM	23 CCBM	24	25	26 CCCC	27	28 C	29	30	31	1	2	3
2562	4 F 2021 July	5	6	7	8 CCC 2246	9 CCB	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25 C	26	27	28 CBC	29	30 CC
2563	F 1	2	3 MCXM		5 2021 Ai	6 ugust	7	8	9 CCB	10	11	12	13	14	15	16 C	17	18 C	19	20	21	22	23	24	25	26	27
2564	28 F	29	30	31	1 2248	2	3	4		6 eptembe		8	9	10	11	12	13	14	15	16	17	18	19	20 CCC	21	22 C	23
2565 2566	24 F C	25	26	27 CCCCC	28 MCCC	29 CC 2249 25	30 C	31	1 28	2	3	4 2021 O		6	4	8 CC 5	9 C 6	10	8	12	13	14	15	16	17	18	19
2567	20 F	21 C 18	22	23 MM 20	24	25 2250 22	26 C 23	27	20 C 25	29	27	28	2	3 30		5 2021 No 1			0	9 M	6	11	8	13 9	14	15	16 12
2568	F 13	14	15	16	17	2251 18	19	20	21			CMMX 24	25	26	27	28	29	30	CC 2021 De 1	cember 2	3	4	5	6	7	8	9
2569	F C 10	11	12	13	14	2252 15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	M 2022 Ja 1	anuary 2	3	4	5
2570	F 6	7	8	9	10	11	CC 2253 12	C 13	14	15	CM 16	M 17	MCC 18	19	CC 20	CCCM 21	22	23	24	25	26	27	28	29	30	31	1
2571	F 2022 F 2 F CC	February 3	4 CC	5	6	7	2254 8	9 C	C 10	11	12 MCCC	13	14 CCM	15	16	17	18	19	20	C 21	CC 22	23	C 24	BCCC 25	26	27	C 28 C
2572	F CC 2022 M F CM	March 2	3	4	5	6 C	2255 7 C	8	9	10	11 CCCC	12	13	14 M	15 CCMC	16	17	18	19	20 CC	21	22 CC	23 CCCC	24 C	25	26 C	27 C
2573	F CMC	29 MCCC	30 CCCC	31 CCCX	2022 A 1 CCCC	pril 2 CCMM	3 CC	2256 4 C	5	6 C	7	8 CC	9 C	10	11	12 CBCC	13	14	15 MMCC	16	17 CCCM	18	19	20 CMCC	21 CC	22 CMC	23 C
2574	24 F	25 MCCC	26	27	28 CC	29	30 MMXM	2022 Ma 1 CCCC	ay 2	3	4 BMCM	5 CMMC	6 CC	7 CC	8 CCC	9 C	10 CCX	11 CCMM	12	13	14	15 C	16 CMC	17 CCCC	18	19 CMMM	20
2575	F CCC	22 CC	23 CC	24 CC	25 CCM	26 CC	27 CC	2258 28 CC	29	30	31	2022 Ju 1	ine 2	3	4	5	6	7	8	9 CC	10 CCMC	11 C	12	13 C	14 CCCC	15	16 C
2576	17 F CCCC	18 C CCC	19 C	20 CC	21 C	22 CC	23 CC	2259 24 C	25 C	26	27	28	29	30	2022 Ju 1	ly 2	3 BC	4 C	5 CC	6	7	8 M	9 CCCC	10 CCCC	11 CM	12 CCCC	13 CCC
2577	14 F CMCM	15 M CCC	16 MCM	17 CCCC	18	19 C	20 CCCC	21	2260	23 CCCC	24 C	25	26 C	27 C	28	29	30	31	2022 Au 1	gust 2 CC	3 C	4	5 CC	6	7	8	9
2578	10 F	11 C	12 CC	13 CCCC	14 C	15 CMMM	16 MCC	17 CCMM		19 MCCC	20	21 C	22 C	23	24	25 CCCM	26 CMMM	27 I CMMM	28 CCMM	29 CMMM	30 CCCM	31 C	2022 Se 1 CC	eptembe 2 CCCC	3		5 CCMC
2579	F CC 2022 0	7 October	8	9	10 C	11 CCCC	12 CCCC	13 CCCC	2262 14 MCCC 2263	15	16 MM	17 M	18 CC	19 C	20 MCCC	21 M	22 CCCC	23 CCCM	24 CCCC	25	26 CCC	27 CCCC	28	29 CC	30 CCMM	2022 O 1 CC	2 CMMC
2580	F CMM	4	5 C 2022 No	6 ovember	7 CM	8 C	9 CCCCC	10 CM	11 MMCC	12 CC	13 C	14 M	15 C	16 C	17	18 C	19	20 CC	21	22 CB	23	24 CC	25 C	26 C	27	28	29
2581	30 F	31	1	2	3	4 2022 De	5 ecember	6 C	7 C	8	9	10 C	11 MCCM	12 CC	13 C	14 C	15 CCC	16 CC	17 CC	18 CCC	19 M	20	21	22 CCC	23	24	25
2582	26 F	27	28	29 C	30	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22